

Simulation and reconstruction for Theia LBL

Guang Yang June 9 2023



Beyond the Theia white paper

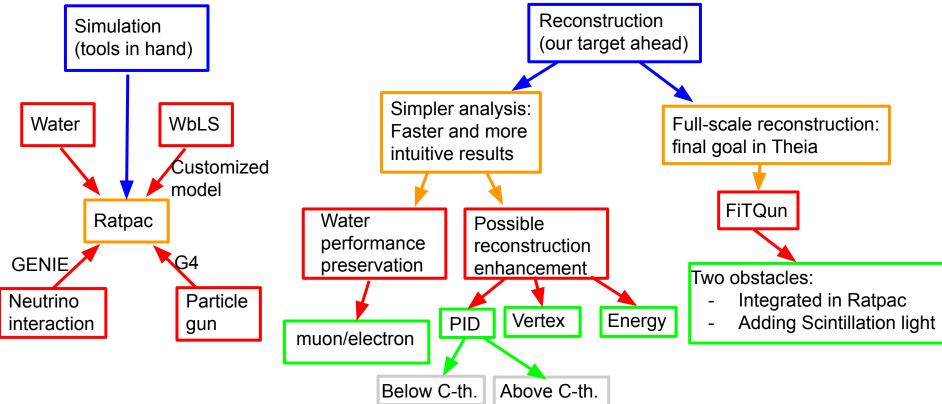
In the Theia white paper, we assumed the same reconstruction performance as water.

- Did not simulate and reconstruct WbLS events.
- Use existing FiTQun reconstruction on SK samples and weight the flux. The main improvement compared to LBNE was including more x-ring y-decay samples and reducing the background with multi-variable technique.

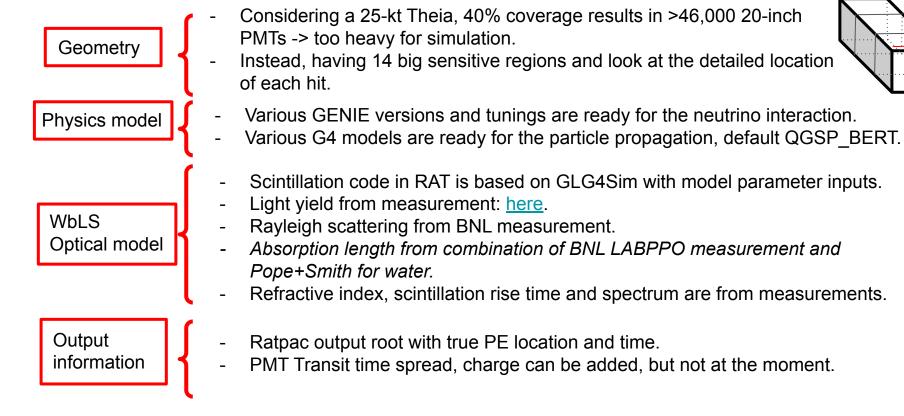
To work on the actual WbLS, simulation and reconstruction from scratch are needed.



Simulation and reconstruction highlight



Simulation with ratpac



Short-term goal

Ideally, we can define good metrics to see the performance improvement as function of scintillation decay time while preserving the Cherenkov ring clarity.

Less ideally, we demonstrate the reconstruction improvement with the scintillation light without damaging the Cherenkov ring significantly.



Topics



Cherenkov ring clarity in presence of scintillation light

Vertex reconstruction

PID below Cherenkov threshold

Energy reconstruction



PID above Cherenkov threshold



Electron/pi0 separation above the Ch. threshold

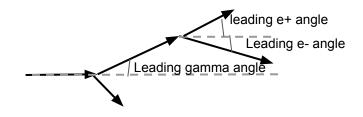
Box shape detector (25 kt Theia)

NCpi0 is a major background in the nue appearance channel for the LBL study.

Both NCpi0 and nue appearance can provide a single ring.

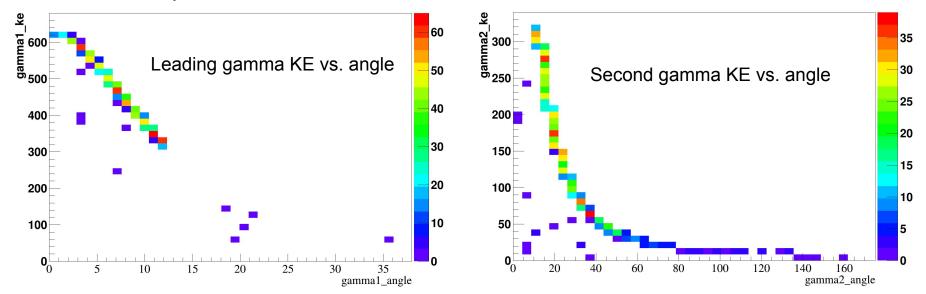
Pi0 and electron particle guns are generated at detector center with various energies.

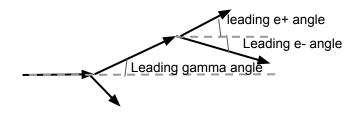
Both pi0 and electron are pointing to the same direction.



Checks on true phase space

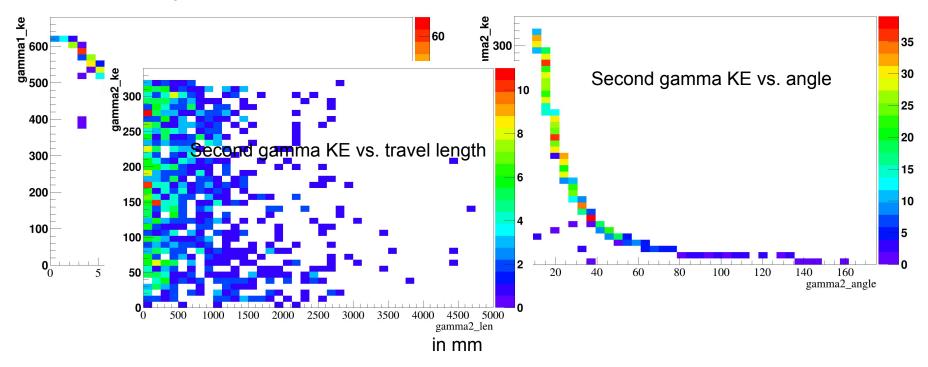
500 MeV KE pi0

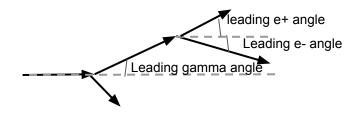




Checks on true phase space

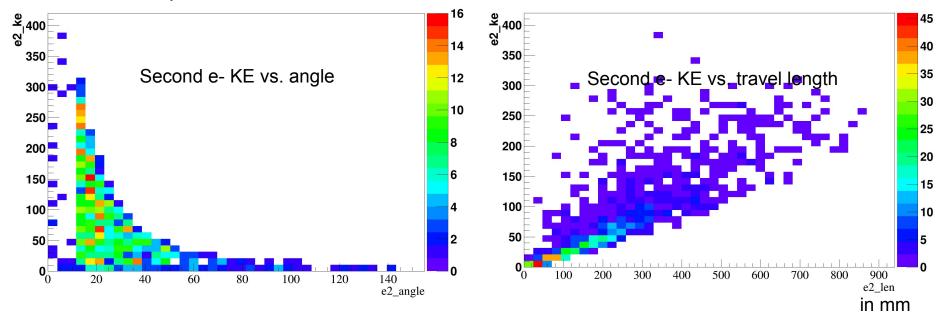
500 MeV KE pi0





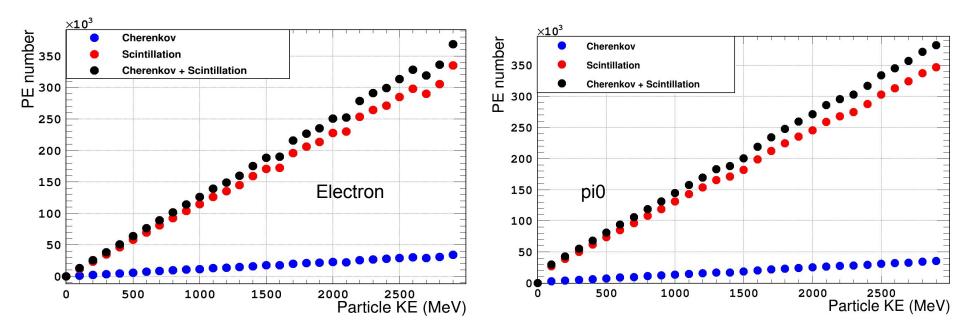
Checks on true phase space

500 MeV KE pi0



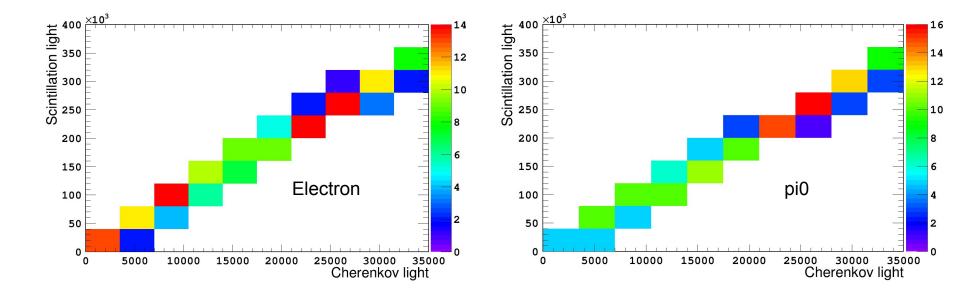


PE vs. energy

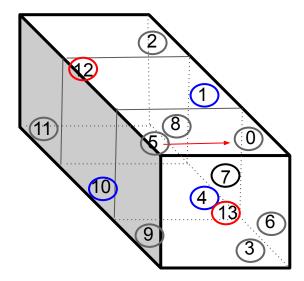




Scintillation light vs Cherenkov light



Defining geometry



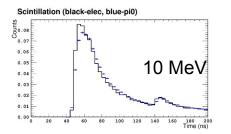
-14 faces are defined. -The result will be on each face.

- 7 is the face that particles pointing to.
- 10,4,1 are the faces surrounding the particle.
- 12 and 13 are the far-away faces.
- 0,2,3,5,6,8,9,11 are the corner faces.

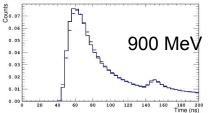


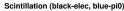


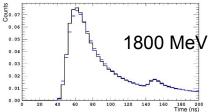
1D scintillation time

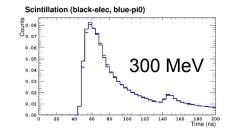


Scintillation (black-elec, blue-pi0)

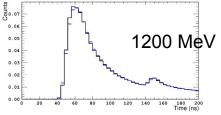




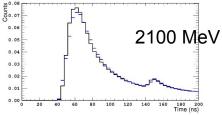






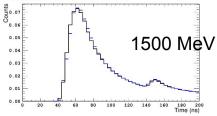




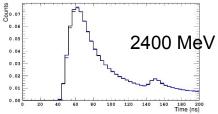


Scintillation (black-elec, blue-pi0) \$0.08 So.07 0.06 600 MeV 0.05 0.04 0.03 0.02 0.01 0.00 b 20 40 60 80 100 120 140 160 180 20 Time (ns) 200

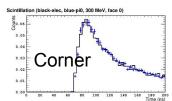
Scintillation (black-elec, blue-pi0)

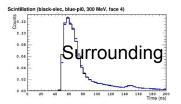


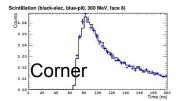
Scintillation (black-elec, blue-pi0)

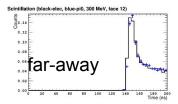


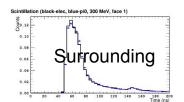
1D time on each fact (300 MeV, blue pi0, black electron)

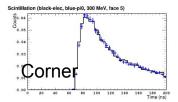


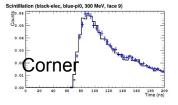


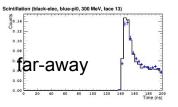


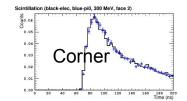


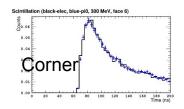


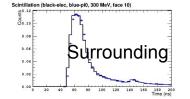


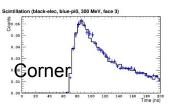


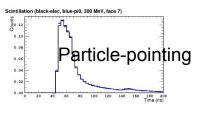


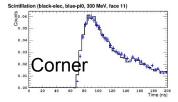




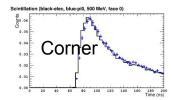


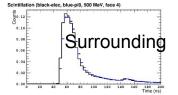


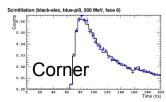




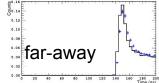
1D time on each fact (500 MeV, blue pi0, black electron)

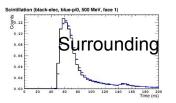


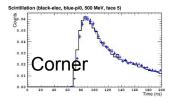


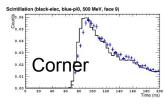


Scintillation (black-elec, blue-pi0, 500 MeV, face 12)





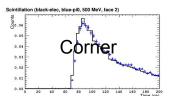


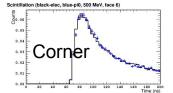


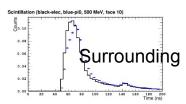
Scinitizion (black-elec, blue-pi0, 500 MeV, face 13)

20 40 60 80 100 120 140 160 180 200 Time (ns)

tar-awav

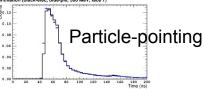




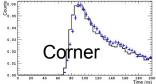


Beneficial control of the second seco

Scintillation (black-elec, blue-pi0, 500 MeV, face 3)



Scintillation (black-elec, blue-pi0, 500 MeV, face 11)



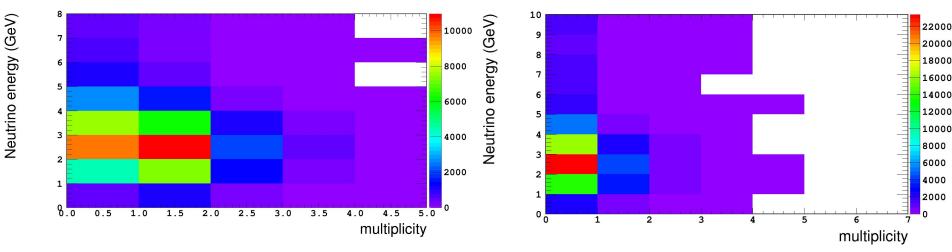


Neutron tagging



Neutron multiplicity in the high-energy events

Antineutrino CC



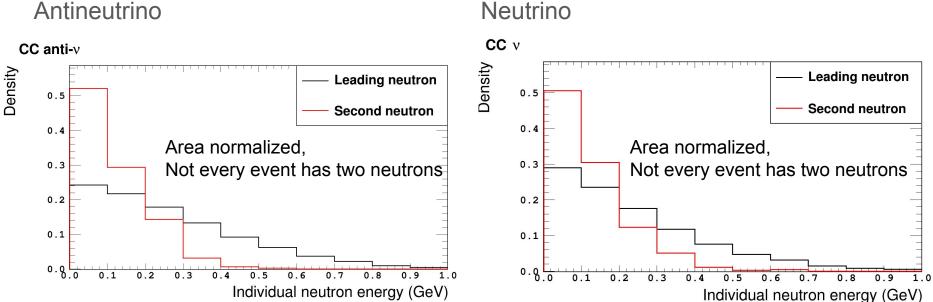
Neutrino CC

For antineutrino interaction, we usually get at least one neutron in the final state.

If neutron capture tagged, the neutrino sign can be determined at a certain level.



Neutron energy in the high-energy events



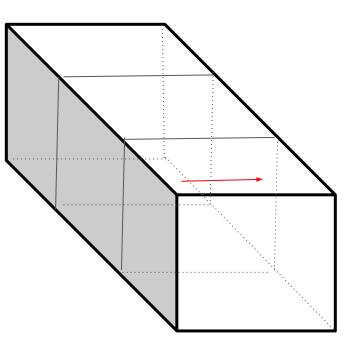
Neutrino



Neutron generation

Neutron particle gun generated at the center of the detector

Default neutron interaction model in geant4.10 (Bertini)



nPE vs. time

10

10²

10²

10

10

10²

103

104

103

WbLS (5pct) Sc.

WbLS (5pct) Ch.

10⁵ Time^{10⁶}(ns)

¹⁰⁵ Time¹⁰⁶(ns)

WbLS (5pct) Sc.

WbLS (5pct) Ch.

- Water Sc.

Water Ch.

Water Sc.

Water Ch.

They was

104

1-D-hat

10³

50 MeV neutron

읽 표 10⁶

105

104

10³

10²

10

. 10⁷

106

10⁵

104

10³

10²

10

ST 107

10⁶

105

10⁴

10³

10²

10

1

1

350 MeV neutron

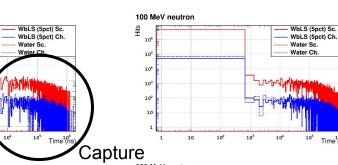
1

200 MeV neutron

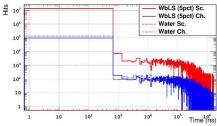
For each energy (plot), 1,000 events accumulated

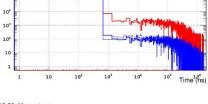
10⁵ Time¹⁰⁶(ns)

¹⁰⁵ Time¹⁰⁶(ns)



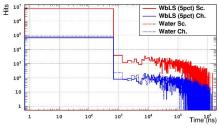
250 MeV neutron



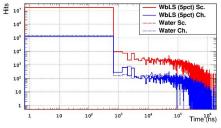


400 MeV neutron Hits WbLS (5pct) Sc. WbLS (5pct) Ch. 10 Water Sc. 10 Water Ch. 105 10⁴ 10³ 10² 10 10² 10³ 10⁴ 1 10

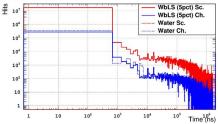
150 MeV neutron



300 MeV neutron



450 MeV neutron







Neutron capture efficiency

More than 95% of neutrons captured in the detector

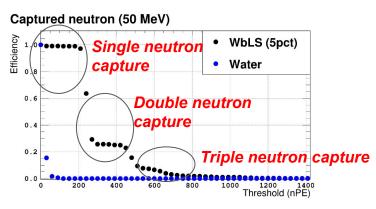
A cut of 1us used:

- Less than 1 us: neutron elastic and inelastic scattering
- More than 1 us: neutron capture process

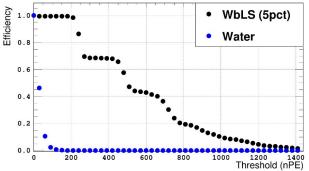
One neutron can result in multiple-neutrons in the detector

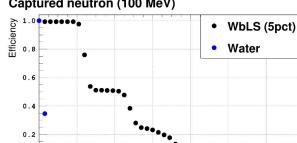
Neutron capture efficiency vs. threshold cut after 1us

Sum all pe after 1 us in each event









600

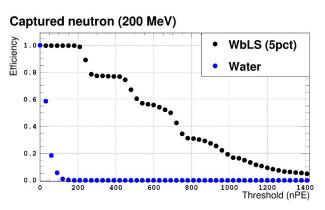
800

400

200

Captured neutron (100 MeV)

0.0



1200 1400 Threshold (nPE)

1400

1000

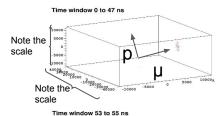




Cherenkov ring clarity

How does a 1 GeV neutrino event looks like 493 KE Proton

375 KE u- and



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instanting and

-10000

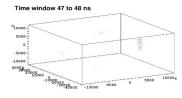
-10000

5000

-5000

10000

10000



Time window 49 to 50 ns

Time window 61 to 70 ns

-10000 -20000 -30000 -40000

Time window 201 to 300 ns

-10020000

Time window 2001 to 3000 ns

-10000 -20000 -30000 -10000

10000

5000

-5000

5000-

-5000

40000 30000 20000 10000

-10000

10000

5000

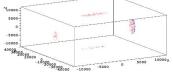
-5000-

10000

40000 30000 20000 10000

40000 38000 20000 10000

1000



-5000

5000

-5000

-10000

-10000

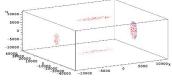
10000

10000

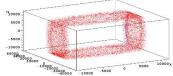
10000.

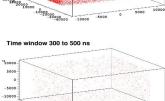
5000





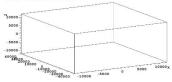
Time window 71 to 100 ns

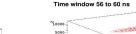




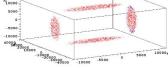
400000 300000 200000 -10000 -30000 -30000 -40000 -10000



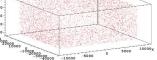




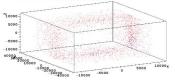
N1000







Time window 1001 to 2000 ns







Na000

5000

-5000

-10000

N1000

5000

-5000

40000 30000 20000 10000

-10000

10000

5000

-5000

-10000

40000 30000 10000

0

40000 30000 20000 10000

-10000 -20000 -30000 -40000

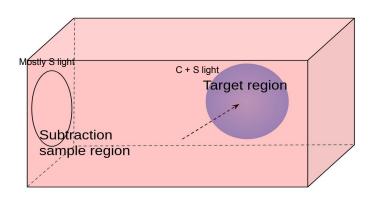
Time window 101 to 150 ns

-10000 -20000 -300000 -40000

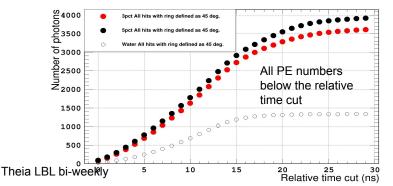
Time window 501 to 1000 ns



Clarity of the water detector information-> muon ring



Total number of photons in ring (500 MeV muon)

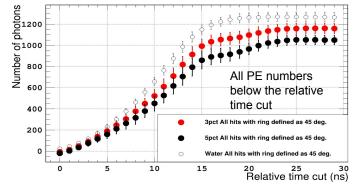


-Define ring clarity as:

Total light in ring - light in far region

- In-ring light will be C+S
- Far region will be primarily S
- The (solid-angle corrected) subtraction should yield the net C signal
- Demonstrate clean identification of Cherenkov ring at few % LS loading for 500-MeV muons

Total number of photons in ring - light in far away region (500 MeV muon)



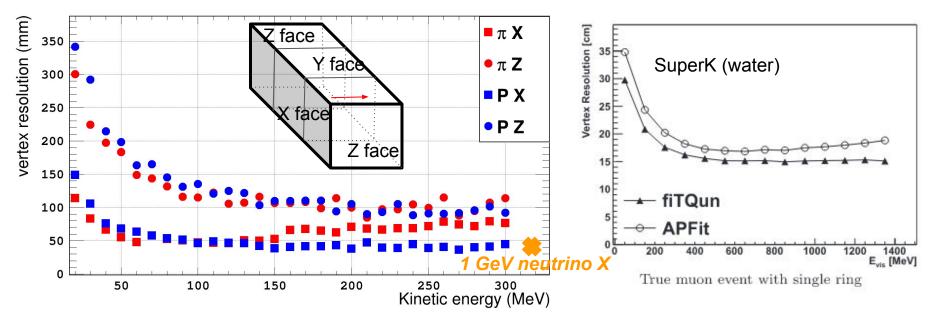


Vertex resolution



Vertex resolution for 5% WbLS with proton/pion

Use the earliest light on each face to determine the vertex location (true at 0)

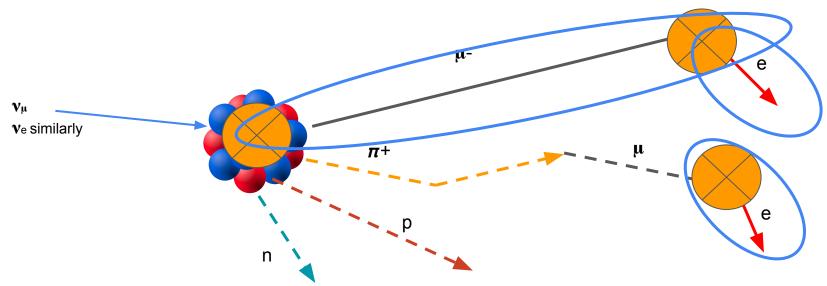




PID below Cherenkov threshold



Scinillation light information



- Without the slow scintillator, we may be able to identify the muon and michels.
- We might aim to improve the reconstruction of the particles below the Cherenkov threshold -> keep in mind the neutrino energy is our final goal.

Particle gun with Proton/**π**+

Each charged particle's PE-KE looks linear. A workflow for CCQE could be:

- C light: muon
- S light: total muon ->proton KE

 $CC1\pi$ + with pion above C threshold:

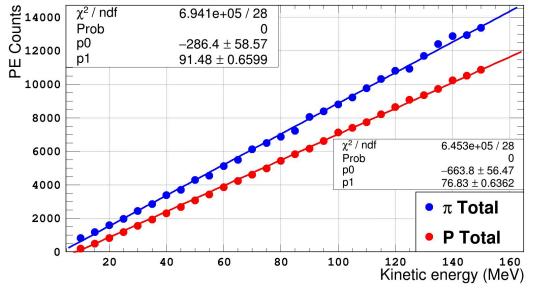
- Similar to QE

 $CC1\pi$ + with pion below C threshold:

- C light: muon
- S light: total muon-> sum of light from proton and pion
- Either assume proton/pion have the same PE-KE

Need PID to identify interaction channel.

5% WbLS

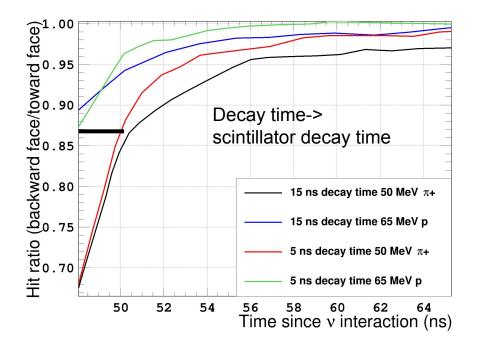






PID : proton/pi+ below Cherenkov threshold

Study the proton/pi+ separation in a more systematic way -> In the Neutrino 2022, we just showed one specific case.

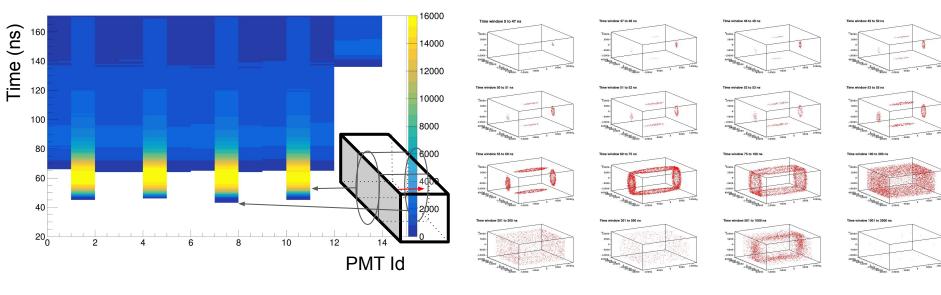


- Given a decay time, for each proton energy, find the corresponding pi+ energy that produce similar light amount (typically about 20 MeV lower).
- 2. Set the hit ratio cut at a few time slices.
- 3. Obtain efficiency and purity for proton with each hit ratio cut for each energy.

Proton/**π**+ separation

As a start, separate the whole detector to 14 "PMTs"

Hope is to use the scintillation light asymmetry due to particle traveling.



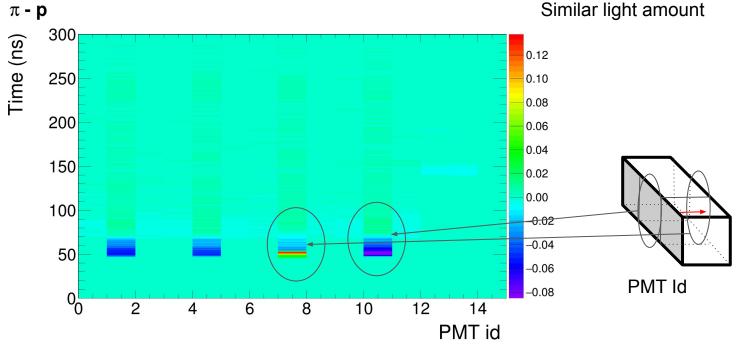
80 MeV π+

Proton/π+ separation

As a start, separate the whole detector to 14 "PMTs"

g.

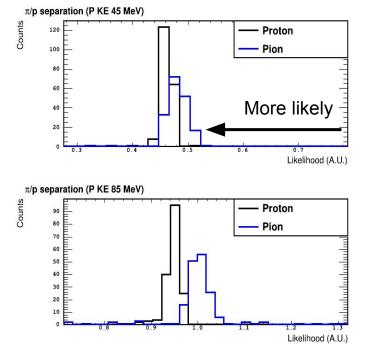
Hope is to use the scintillation light asymmetry due to particle traveling.

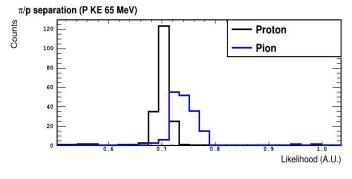


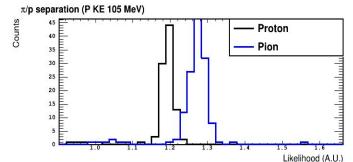


Using the Eos likelihood framework

Taking proton PDFs to fit for proton and pion sample (5 ns decay time).







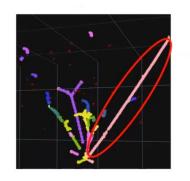
Energy resolution

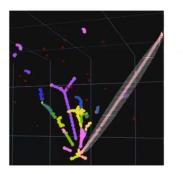


Energy reconstruction

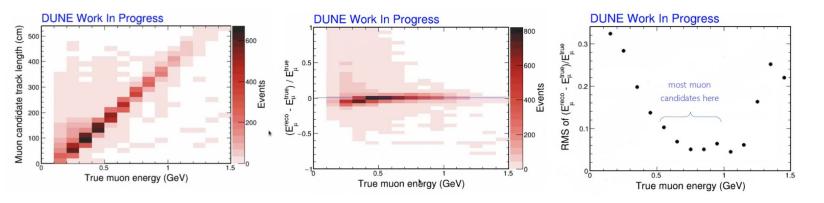
How DUNE LAr works

$$E_v(v_\mu) =$$





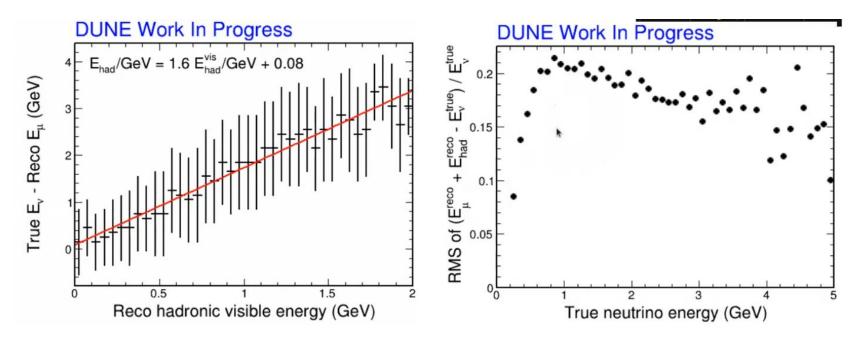
muon energy (range) everything else (calorimetric)





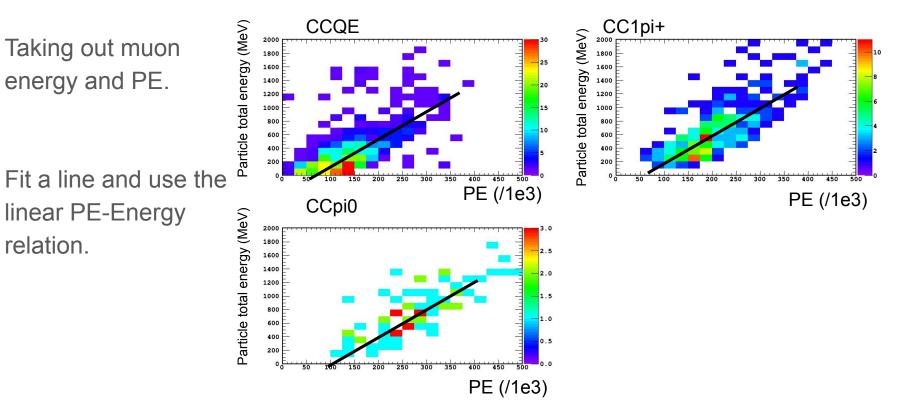
Energy reconstruction

DUNE hadronic energy





Everything except muon in 5% WbLS

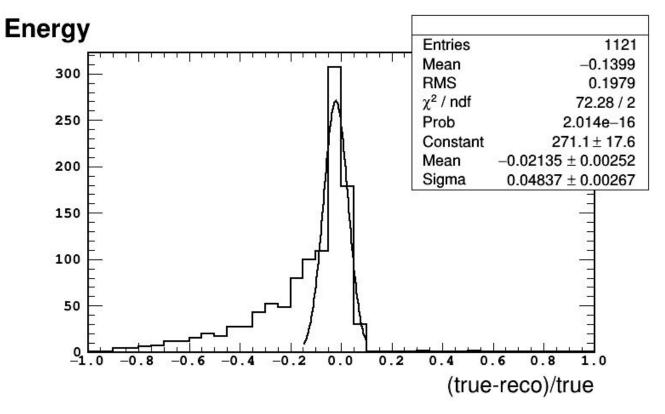




Energy resolution

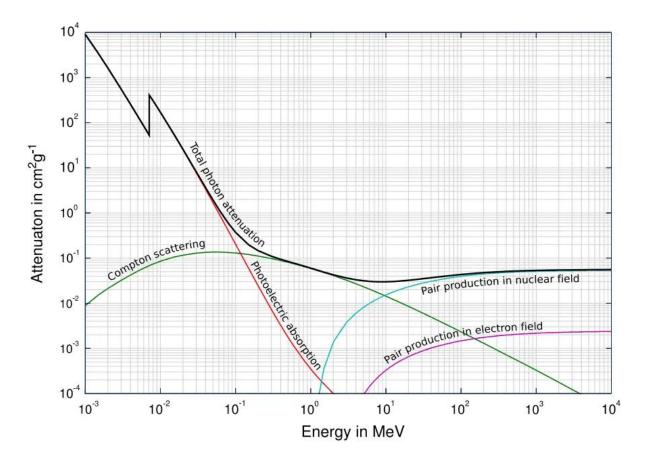
CCQE + CC1pi+ + CC1pi0 combined

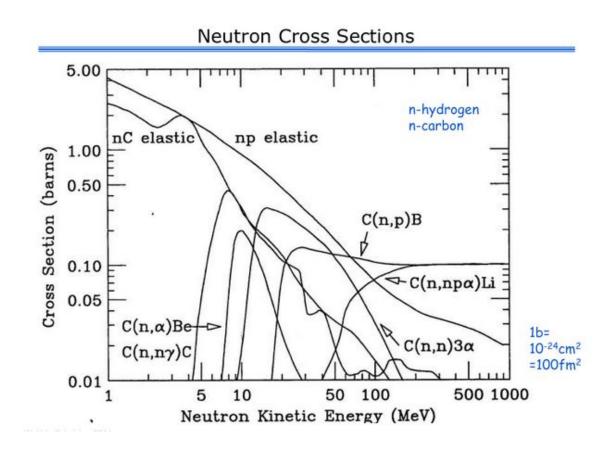
I don't have any correction like what DUNE does. Will look into that. I expect at least similar resolution.





Backups





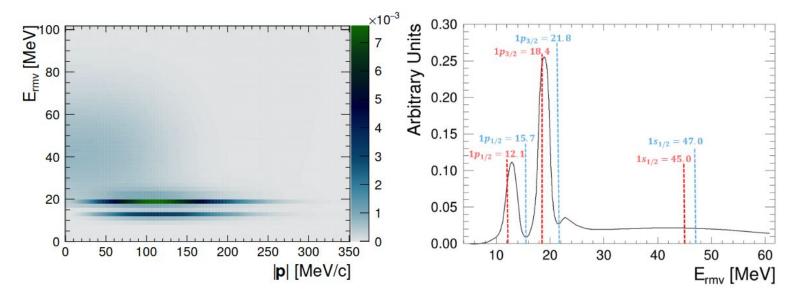


Fig. 8: The two-dimensional probability density distribution for the spectral function for oxygen in NEUT [43] (left), and the projection onto the removal energy axis (right). On the left, the darker colour represents a higher probability of finding an initial-state nucleon with a particular removal energy and momentum. The two sharp p-shells at $E_{rmv} \sim 12$ MeV and $E_{rmv} \sim 18$ MeV, and the larger diffuse s-shell at $E_{rmv} \sim 20 - 65$ MeV and $|\mathbf{p}| < 100$ MeV/c, are visible. The predictions for the shell positions from another model [58] are overlaid on the right with dashed lines, for protons (red) and neutrons (blue). The energy in MeV is labelled for each prediction.

FiTQun



A full reconstruction providing information of neutrino interaction vertex, number of rings, and momentum and PID of each ring.

For a single ring, there are seven reconstructed quantities: location (3), momentum (1), direction (2), time (1), denoted as \mathbf{x} .

$$\mathcal{L}(\mathbf{x}) = \prod_{\text{unhit}} (1 - P(i \text{ hit}; \mathbf{x})) \times \prod_{\text{hit}} P(i \text{ hit}; \mathbf{x}) f_q(q_i; \mathbf{x}) f_t(t_i; \mathbf{x})$$

P is the likelihood the PMT got hit; q,t are measured charged and time in pmts; fq is the charge profile; ft is the time profile;

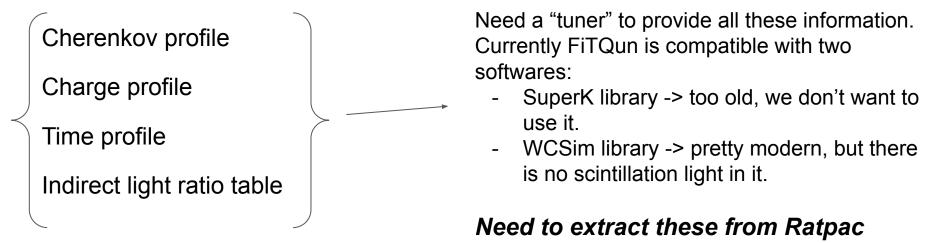


Four steps

- 1. Simulate different kinds of events
- 2. Extract needed information from the simulation
- 3. Convert to the FiTQun-style input
- 4. Run FiTQun with test samples



Input for FiTQun



Since we are using different material and detector geometry, all these might need to be re-generated, although would be nice to use existing profiles..



Workflow: each step can be called a milestone

Need to extract information from Ratpac -> a big chunk of hack, tested part of them, they seem to work.

Simulate a large amount of events for each profiles.

Expand FiTQun to include the scintillation light

Need a lot of validations: looking at water first.



Summary

Simulation is ready for event reconstruction studies.

Largely, the there are two streams:

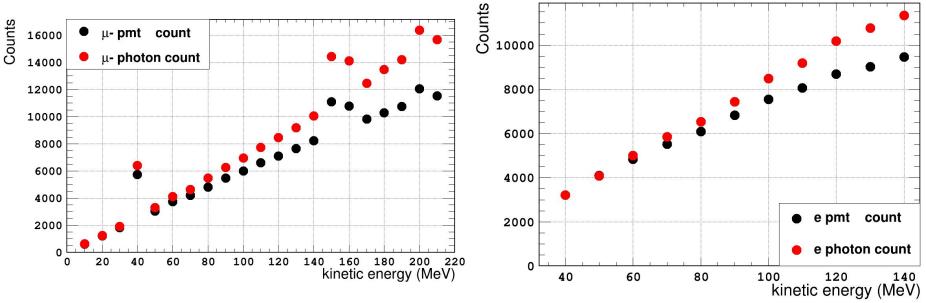
- Simple physics studies about basic performance: fast but isolated
- FiTQun full reconstruction: complete but difficult

In the future meetings, Gian and I may present some detailed work on some branches.



Particle gun with muon and electron

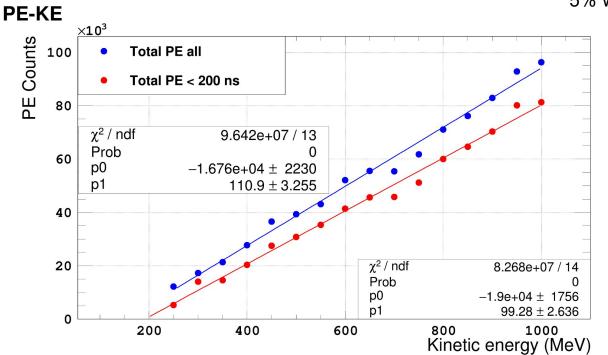
5% WbLS



Single event at each KE used, so fluctuation is strong.



PE-KE for true $CC0\pi$



5% WbLS

Parameter	Name in ratdb	Measured (yes/no)	Comment	Reference
Light yield	LIGHT_YIELD	yes		https://doi.org/10.1140/epjc/ s10052-020-8418-4
Rayleigh Scattering	RSLENGTH	yes		private communication with BNL
Absorption Length	ABSLENGTH	no	combination of BNL data for LABPPO and Pope+Smith for water	
Refractive index	RINDEX	yes		
Reeemission probability	REEMISSION_ PROB	no	0.8 for w< 345nm, 0 for w >370	
Scintillation rise time	SCINT_RISE_TI ME	yes		
Scintillation time profile for betas	SCINTWAVEFO RM	yes		https://doi.org/10.1039/ D0MA00055H
Scintillation time profile for alphas	SCINTWAVEFO RMalpha	no	used same as SCINTWAVEFORM	
Birk's constant for betas	SCINTMOD	no	SNO+ (measurements for WbLS undergoing)	
Birk's constant for alphas	SCINTMODalph a	no	SNO+ (measurements for WbLS undergoing)	
Birk's constant for neutrons	SCINTMODneu tron	no	SNO+ (measurements for WbLS undergoing)	
Scintillation emission spectrum	SCINTILLATION	yes		https://doi.org/10.1039/ D0MA00055H
Scintillation emission spectrum for wavelength shifters	SCINTILLATION _WLS	yes	used same as SCINTILLATION, LAB->PPO energy transfer should be non-radiative, so	

https://docs.google.com/spreadsheets/d/1QqpolQU69itKQxvAd-bvHCZZQVoRMGxC6zsBcNU4Yfo/edit?usp=sharing

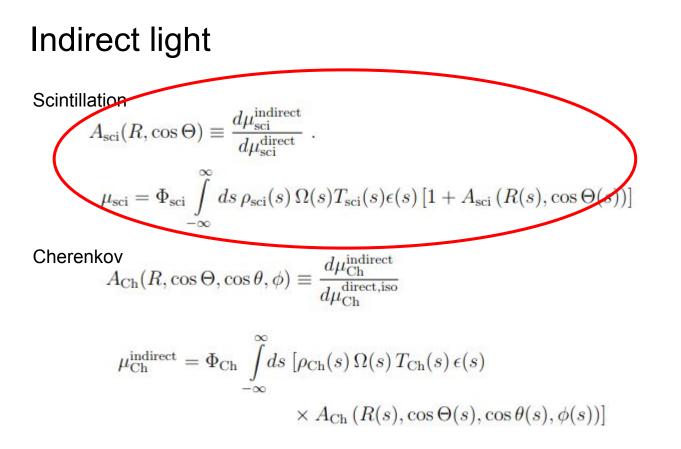
$$F(\mathbf{x}) \equiv -\log \mathcal{L}(\mathbf{x}) \equiv F_q(\mathbf{x}) + F_t(\mathbf{x})$$

$$F_q(\mathbf{x}) \equiv -\sum_{\text{unhit}} \log(1 - P(i \text{ hit}; \mu_i)) - \sum_{\text{hit}} \log(P(i \text{ hit}; \mu_i)f_q(q_i; \mu_i))$$
The μ is the predicted mean charge.
$$arXiv. \ 0902.2222$$
Scintillation light
$$\mu_{\text{point,sci}} = \Phi_{\text{sci}} \Omega(r) T_{\text{sci}}(r) \epsilon(\eta)$$

$$LY \qquad \text{Solid angle Transmission acceptance}$$

$$\mu_{\text{sci}} = \Phi_{\text{sci}} \int_{-\infty}^{\infty} ds \ \rho_{\text{sci}}(s) \Omega(s) T_{\text{sci}}(s) \epsilon(s)$$
Similarly,
$$\mu_{\text{Ch}} = \Phi_{\text{Ch}} \int_{-\infty}^{\infty} ds \rho_{\text{Ch}}(s) \Omega(s) T_{\text{Ch}}(s) \epsilon(s) g(\cos \theta(s); s)$$

$$Ch. \text{ angular profile}$$



Two layers of questions

- How can we use the scintillation information?

-> As soon as we can separate out the Cherenkov light largely, we can use all the remaining light.

- How would the Slow-Scintillator help?

-> Slow-scintillator may give additional information, but how?

Obviously, show-scintillator can separate out the Cherenkov light better.

What do we do with water? -> T2K

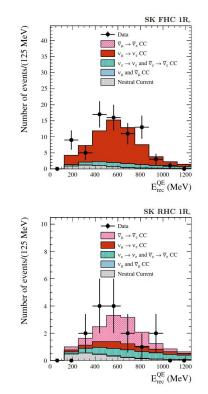
CP: nue appearance

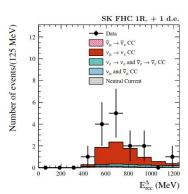
Nu mode

- 1 e-like ring 0 decay
- 1 e-like ring 1 decay

Antinu mode

- 1 e-like ring 0 decay





What do we do with water? -> Theia white paper

CP sensitivity

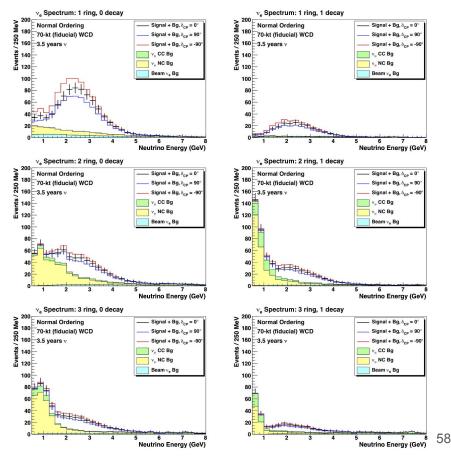
Nu mode:

- 1,2,3 rings with 0,1 decays

Antinu mode:

- 1,2,3 rings with 0 decay

We have information of the primary Cherenkov ring and decay Cherenkov ring.

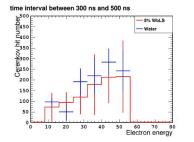


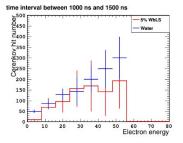
Clarity of the water detector information-> michel ring

The same trick as the muon ring, 1 GeV neutrino

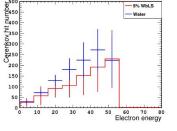
All light in ring - light in an away region with the same solid angle

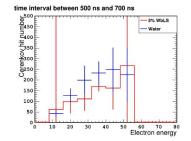
Error bar shows the event-by-event deviation Note that the neutrino energy spectrum is broad.

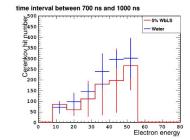




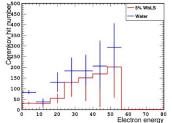








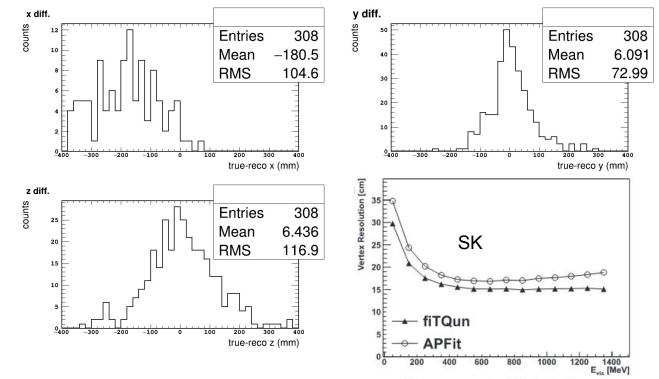
time interval between 3000 ns and 5000 ns



A clear improvement -> vertex resolution

Only for true CC1π+ channel with DUNE flux;

vertex determined by the first hit on each face.



True muon event with single ring

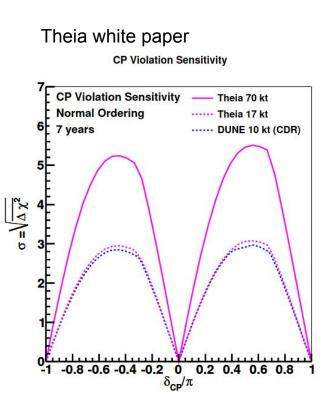
FiTQun

Being used in Super-K, T2K and WCSim

It has also been used in the Theia long-baseline result.

Performance -> Atmospheric Neutrino Oscillation Analysis with Improved Event Reconstruction in Super-Kamiokande IV: arXiv. 1901.03230

Principle -> The extended-track reconstruction for MiniBooNE: arXiv. 0902.2222



Beyond that

Even with all the input information, previously, although FiTQun contains scintillation light information for each track, it did not really work with the particles below the threshold.

We will have low energy proton, pion etc. -> it needs some effort to look into those events below the Cherenkov threshold in FiTQun. In principle, it should be straightforward.

Geometry

Theia 25 kt letter box inside DUNE cavern: 20 m x 18 m x 69 m

500 MeV muons at the end of the detector

