

CIPANP 2018

Neutrino Mass and Neutrino Mixing parallel session

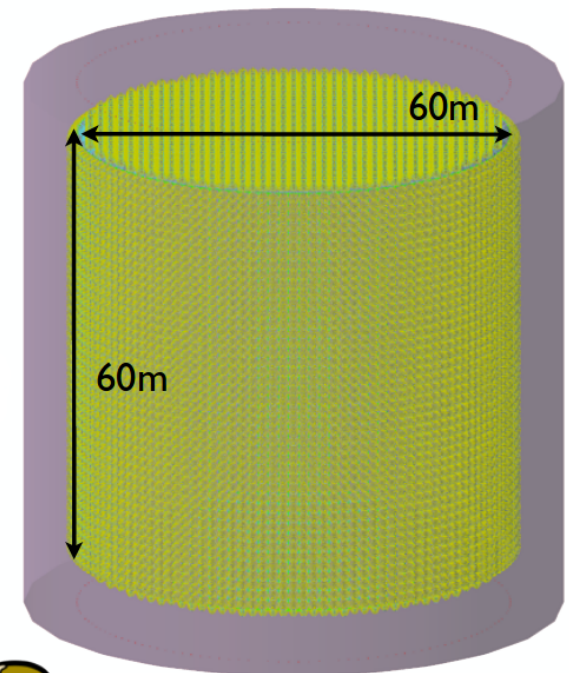
Theia:

A multi-purpose water-based
liquid scintillator detector

Vincent Fischer

University of California at Davis

- Greek mythology: Theia (Θεία) is the Titan goddess of light, mother of Helios (Sun), Eos (Dawn), and Selene (Moon)
- **Theia** is a project for a large multi-purpose water-based liquid scintillator detector
- **~50 kilotonnes** of water-based liquid scintillator (**WbLS**)
- **Physics goals:** CP violation, mass hierarchy, solar neutrinos, double-beta decay, sterile neutrinos, geoneutrinos, etc...
- Use of fast-timing photosensors known as **LAPPDs**



More details: [arXiv:1409.5864](https://arxiv.org/abs/1409.5864)






~60 proto-collaborators in more than 6 countries

4 meetings already occurred (2 in the US, 2 in Germany)

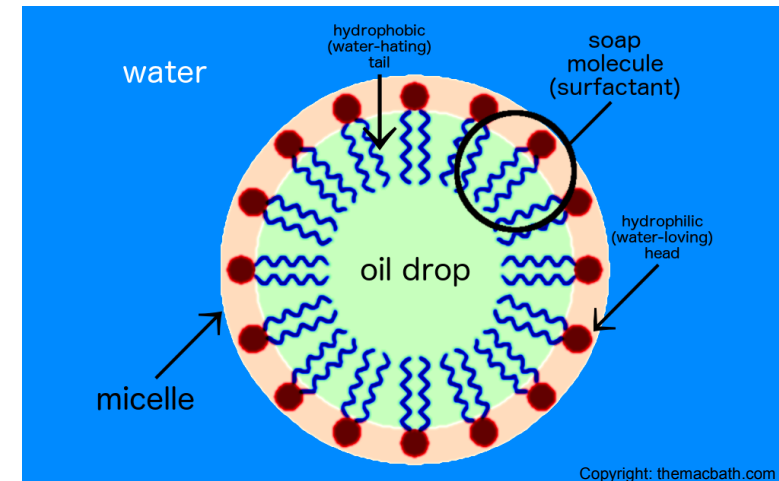
Last meeting (April) was at UC Davis with a large attendance and interests from other collaborations
(ANNIE, WATCHMAN)



Theia workshop

12-14 April 2018
University of California, Davis
US/Pacific timezone

- **Water-based Liquid Scintillator (WbLS)** is a mixture of pure water and oil-based liquid scintillator
- While water and oil don't mix, WbLS is made using a **surfactant** (soap-like) such as PRS* (hydrophilic head and hydrophobic tail) to hold the scintillator molecules in water in a “**micelle**” structure
- **Combines the advantages** of water (low light attenuation, low cost) and liquid scintillator (high light yield)
- Emission of **prompt Cherenkov** light and **delayed scintillation** light
- **Tunable LS content** for a broad range of physics goals
- **Low cost** and **environmentally-friendlier** than pure LS
- Strong R&D effort ongoing at **Brookhaven and Berkeley Nat. Labs** and **UC Davis**

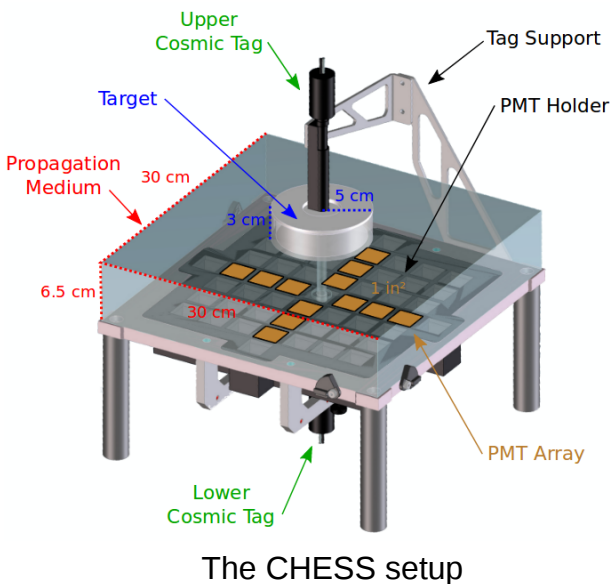


Micelle structure in water

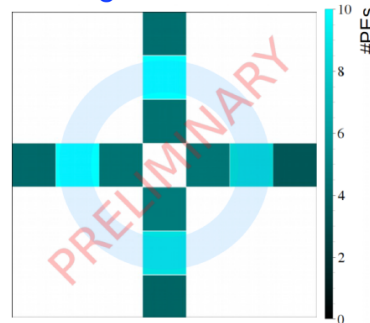


Former UCD graduate student **Morgan Askins** at BNL

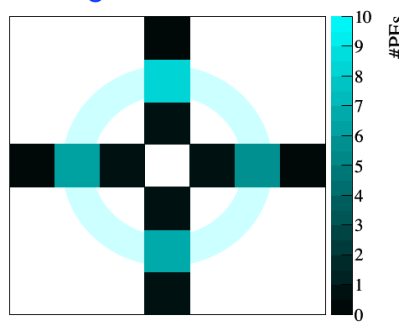
- **Time and charge reconstruction** show separation of Cherenkov (directional) and scintillation light (isotropic) in simulations
- Ongoing tests performed with **CHES** (Eur. Phys. J. C (2017) 77: 811) to **separate both components**



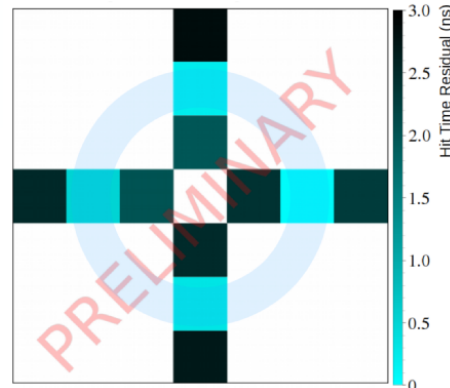
Average of WbLS data set



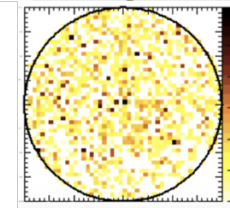
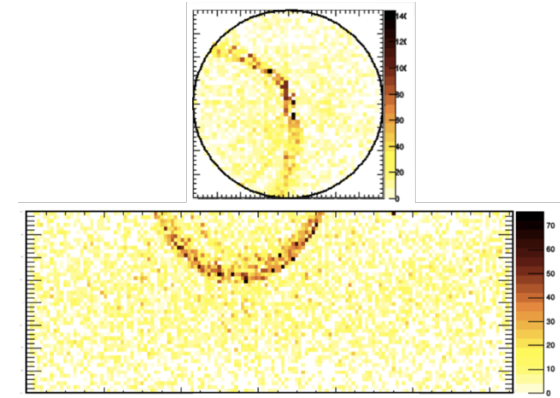
Average of water data set



Number of PEs detected

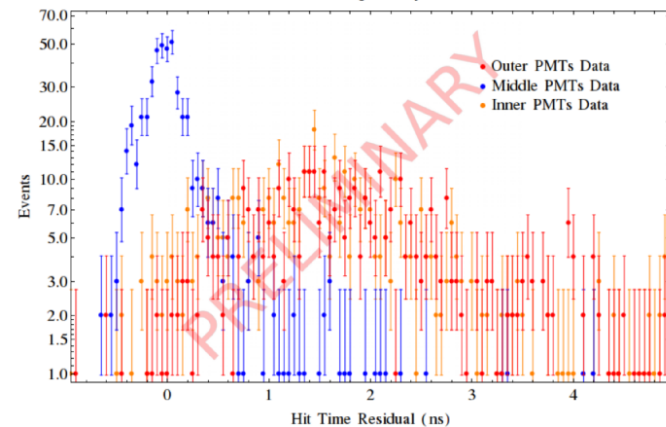


Mean hit time residuals



Simulated CC interaction in WbLS

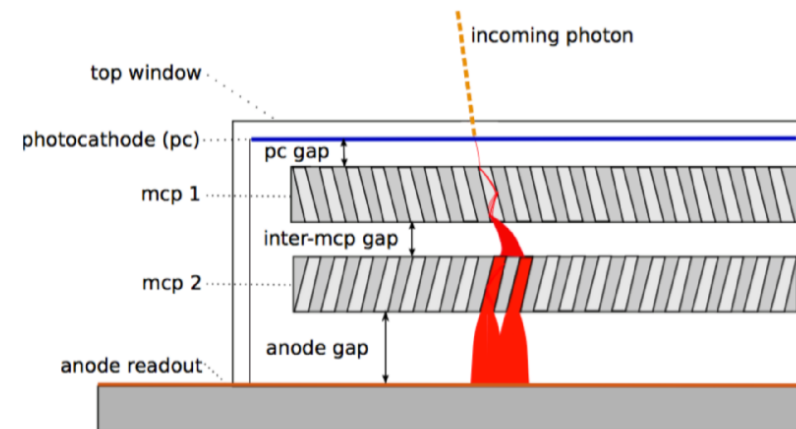
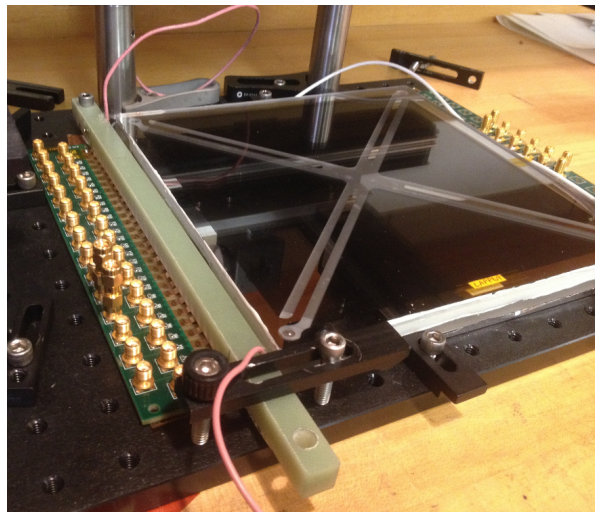
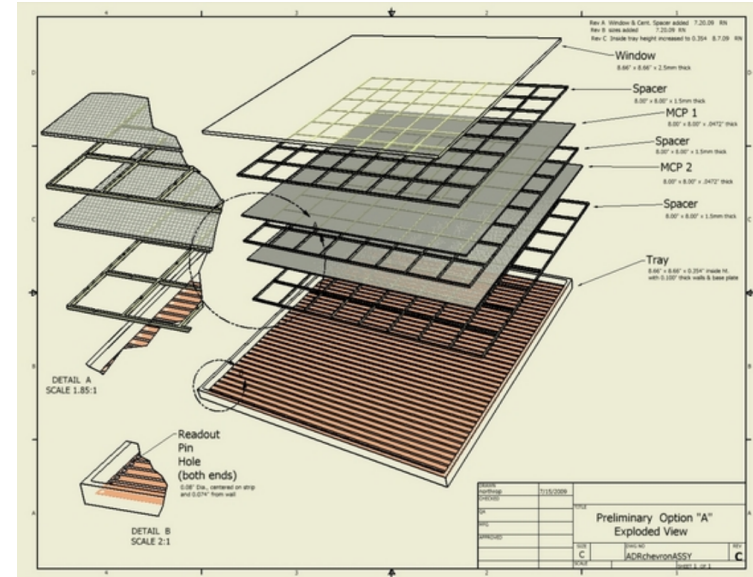
Credit: Leon Pickard (UC Davis)

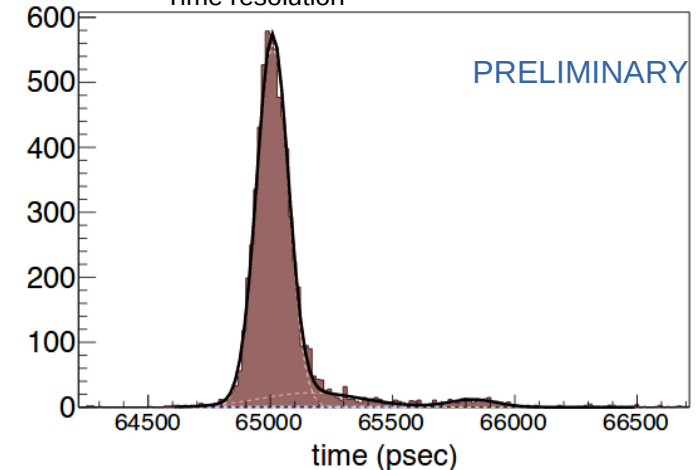
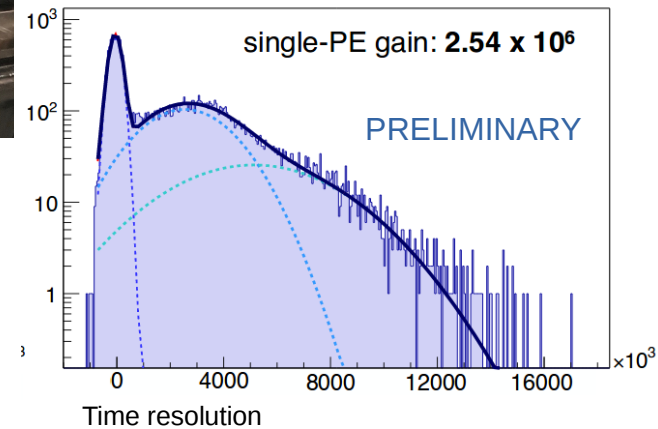
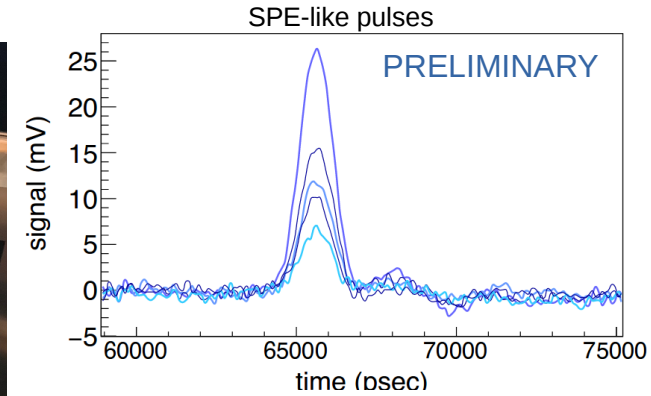
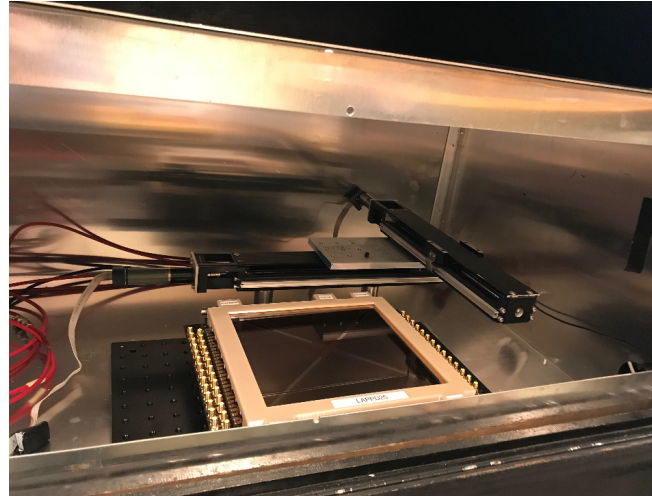
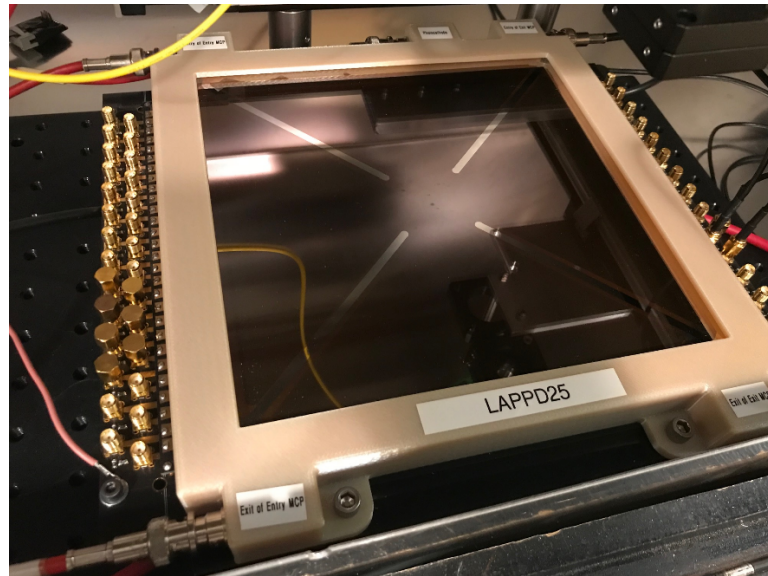


Hit time residuals

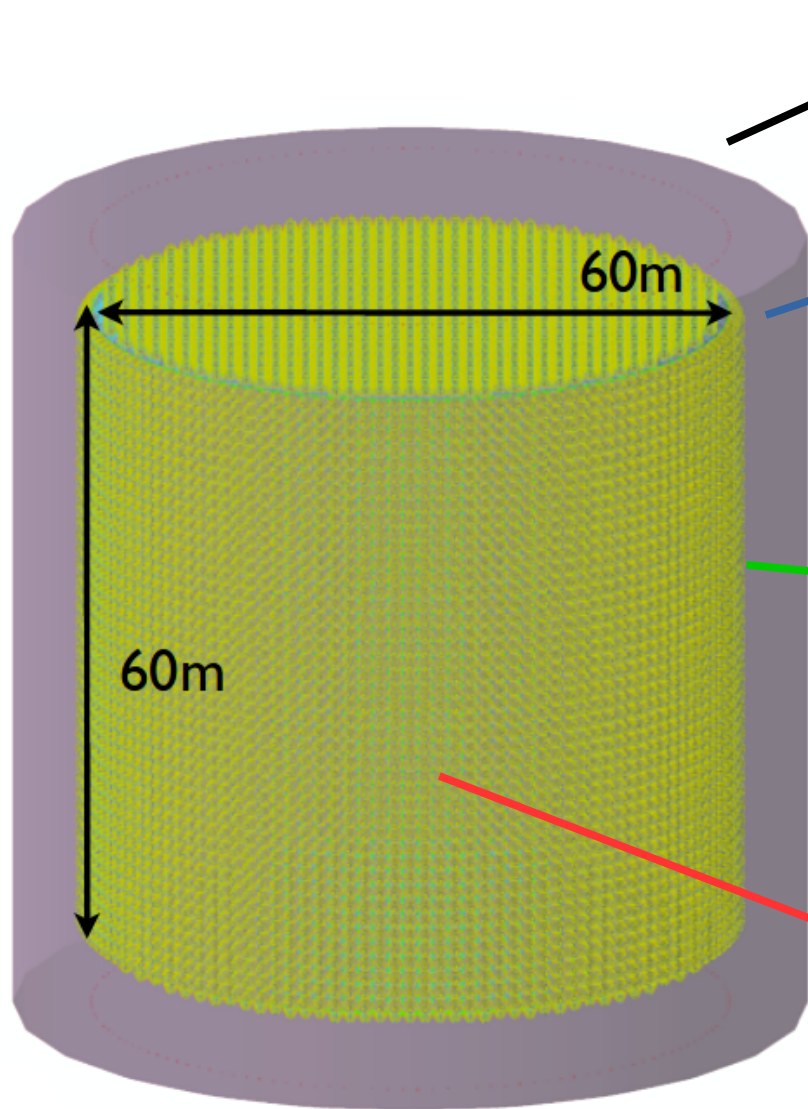
Large Area Picosecond PhotoDetectors - Principles

- Large, flat **MCP-based** photosensors
 - Time resolution: **~60 picoseconds**
 - Spatial resolution: **< 1 cm**
- Design includes an **intrinsic PSEC-4 readout system**
- Manufactured by the **Incom company**, being tested extensively at **Iowa State University** and expected to be deployed in the **ANNIE** detector
- **More information:** A Brief Technical History of the Large-Area Picosecond Photodetector (LAPPD) Collaboration, arXiv:1603.01843





- **LAPPD Tile #25** is the **first one to be deployed in ANNIE**
 - First deployment in an operating neutrino detector
- Being **thoroughly tested** at ISU (gain, time resolution, etc.)
- Preliminary results show a **good behavior** even at lower voltage than expected during nominal operation



- Underground location ($> \sim 4000$ m.w.e)

- **Outer shielding:**

- Pure water
- Instrumented to act as an active veto

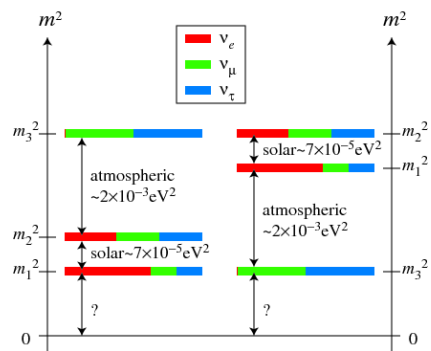
- **Support structure:**

- Diameter/Height ~ 60 meters
- Holds > 100000 PMTs
→ **90% coverage**

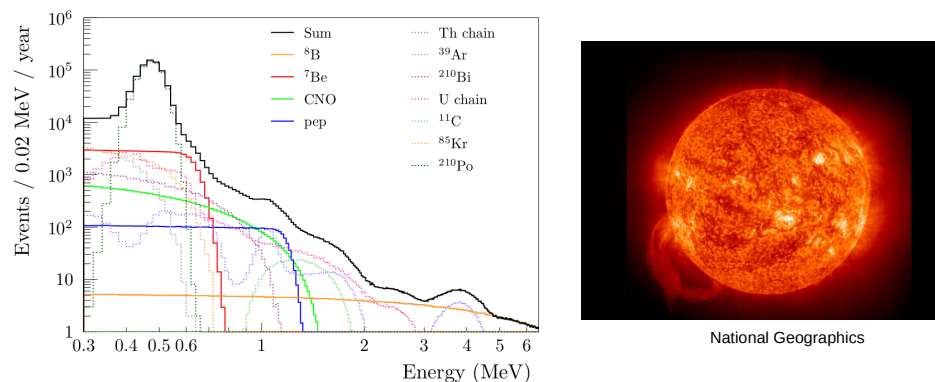
- **Target volume:**

- Reference design: **Water-based Liquid Scintillator**
- Possibility to **load metals**
- Possibility to **deploy inner “bag”**

Long baseline



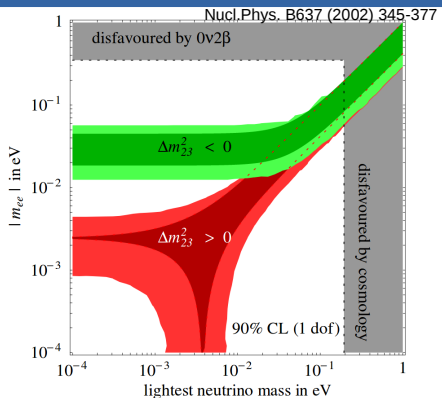
Solar neutrinos



Supernova neutrinos

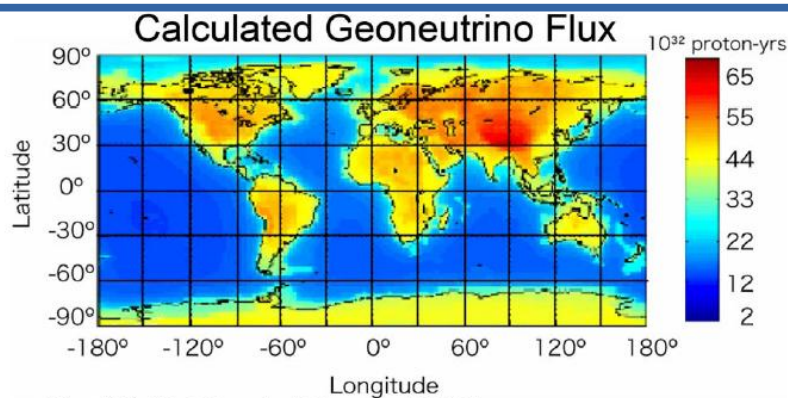


Double-beta decay

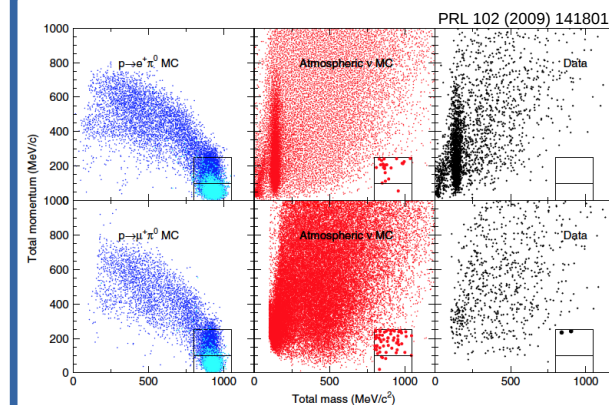


Broad range of neutrino physics!

Geoneutrinos

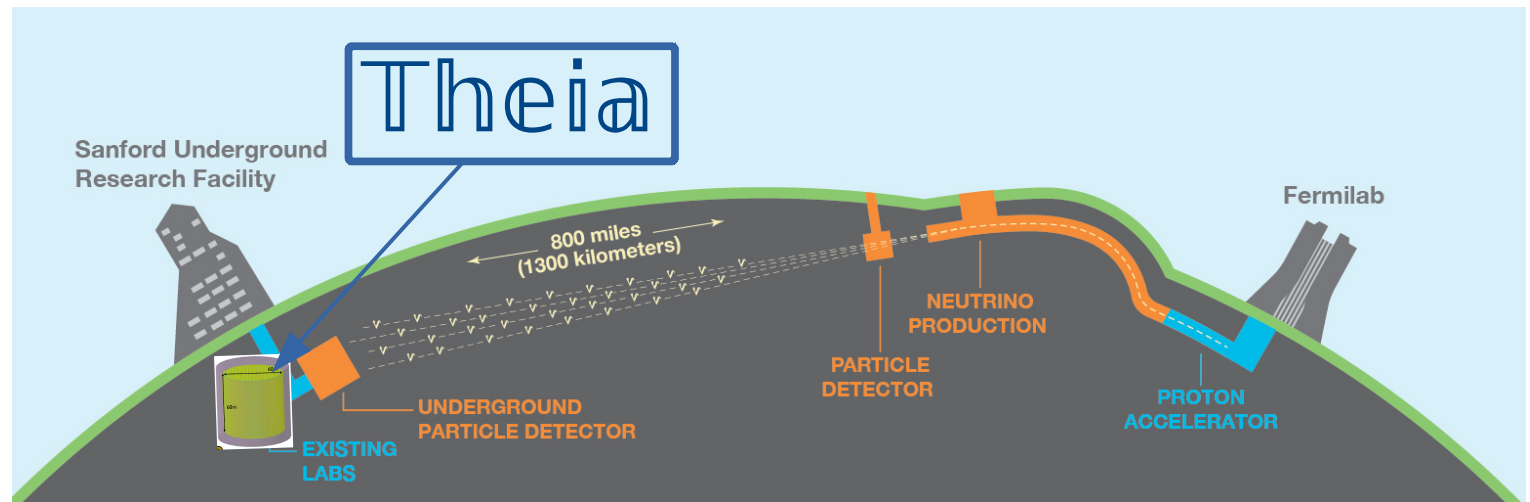


Nucleon decay



... and more !

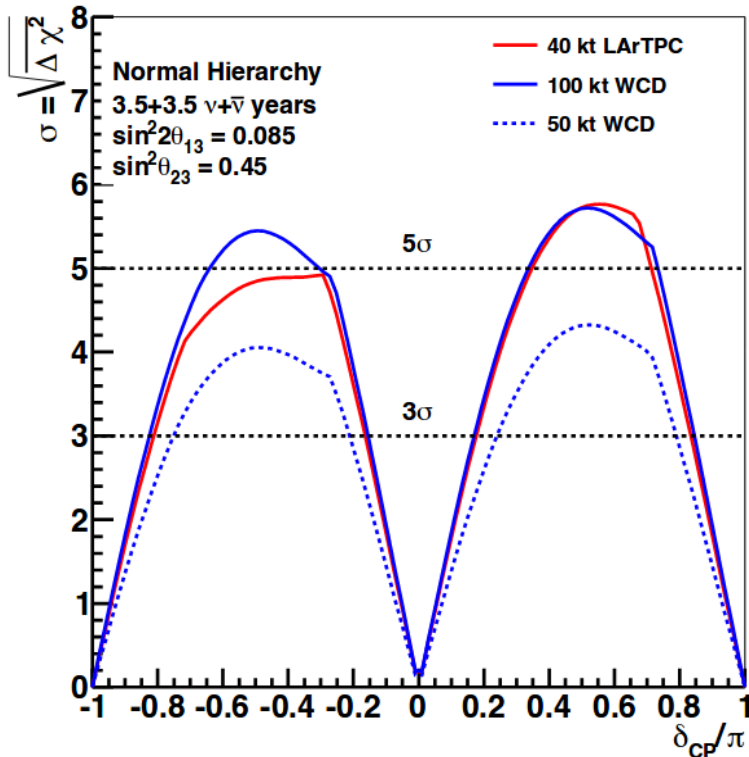
- **DUNE:** Four 10-kt liquid argon detectors placed at SURF
 - 4000 m.w.e. overburden
 - High intensity beam from Fermilab 1300 km away
- Another detector using a different detection medium in the DUNE cavern would **increase the sensitivity to CP violation and mass hierarchy** (if not already discovered)
- Design studies ongoing



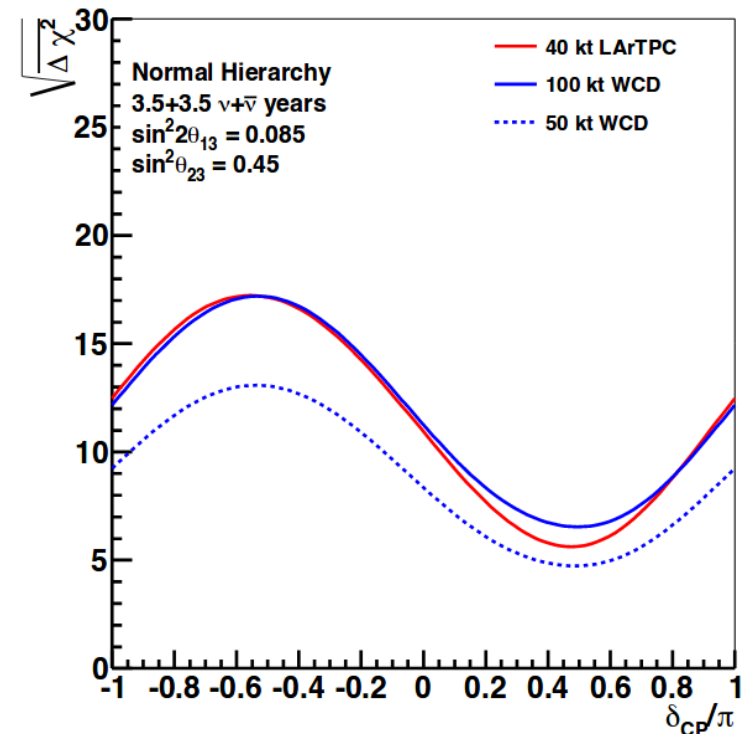
- The addition of a **50-kt water Cherenkov detector next to DUNE** increases the sensitivity to the missing neutrino oscillation parameters:
 - CP violation: **5σ over 50% of the δ_{CP} range**
 - Mass hierarchy: **$> 5\sigma$ the whole δ_{CP} range**

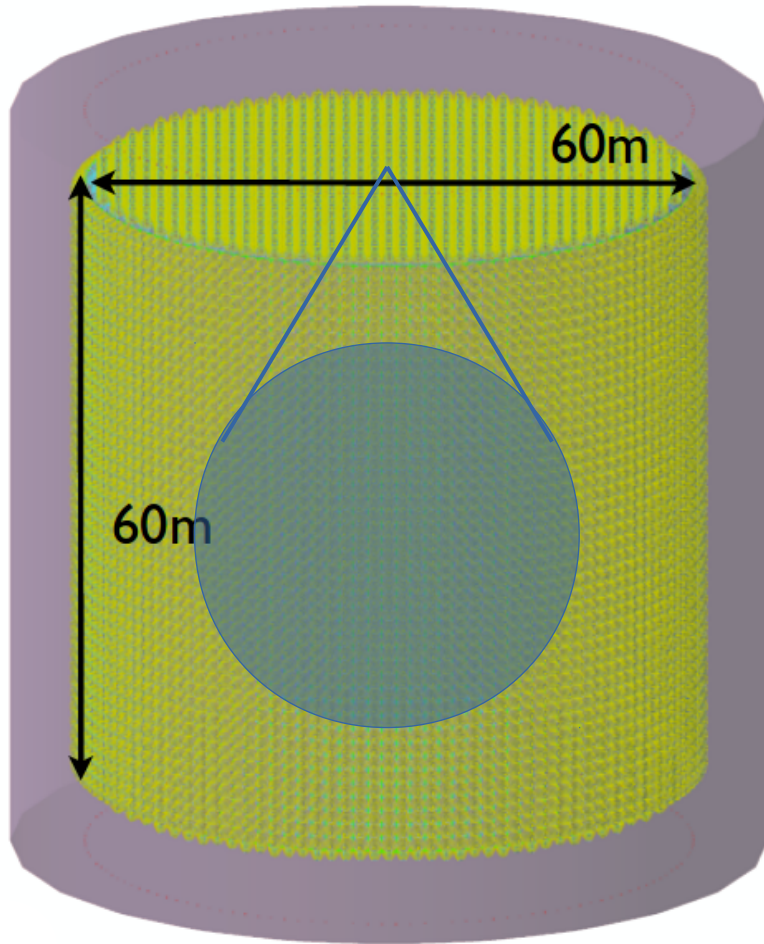
- However, this is a **conservative scenario for Theia+DUNE**:
 - Increased background rejection
 - Reconstruction of missing energy (neutrons)
 - Fast timing sensors

CP Violation Sensitivity



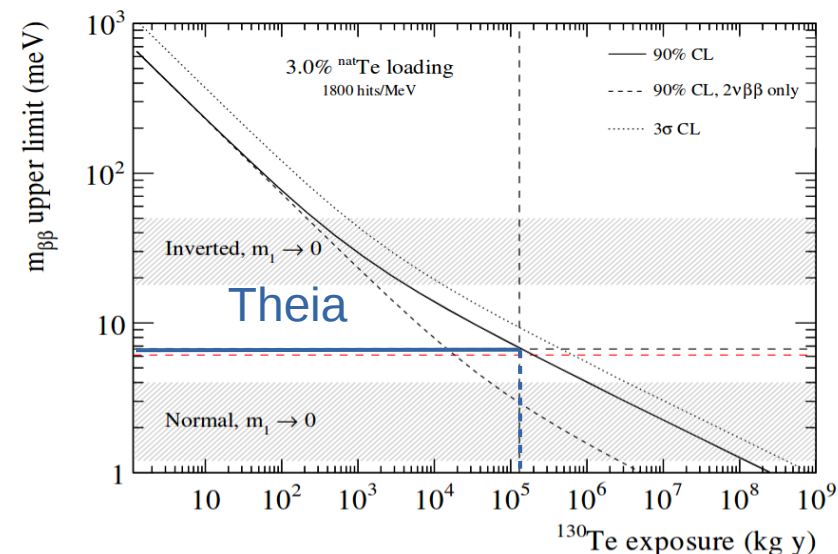
Mass Hierarchy Sensitivity





- $0\nu 2\beta$ approaches:
 - **Entire detector** filled with WBLs loaded with a 2β isotope
→ “SNO+-like”
 - **Only inner balloon** filled but with a higher concentration of 2β isotope
→ “KamLAND-Zen-like”
- **Fiducialization** and **tagging techniques** (triple coincidence, directionality, etc..) greatly reduce backgrounds
- Isotopes in consideration: ^{136}Xe and ^{130}Te
- **Goal: Reach $T_{1/2}^{0\nu 2\beta} \sim 10^{28}$ years**

Assumptions:
- 8 m balloon
- 3% Te loading
(300 tonnes)



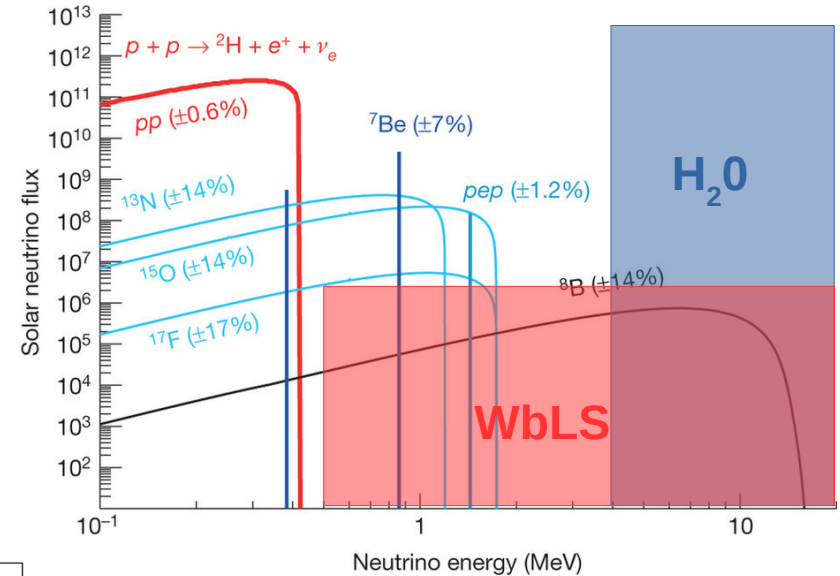
Limit on $T_{1/2}^{0\nu 2\beta} = 1.2 \times 10^{28}$ years

Theia will completely probe the IH part of the mass phase space!

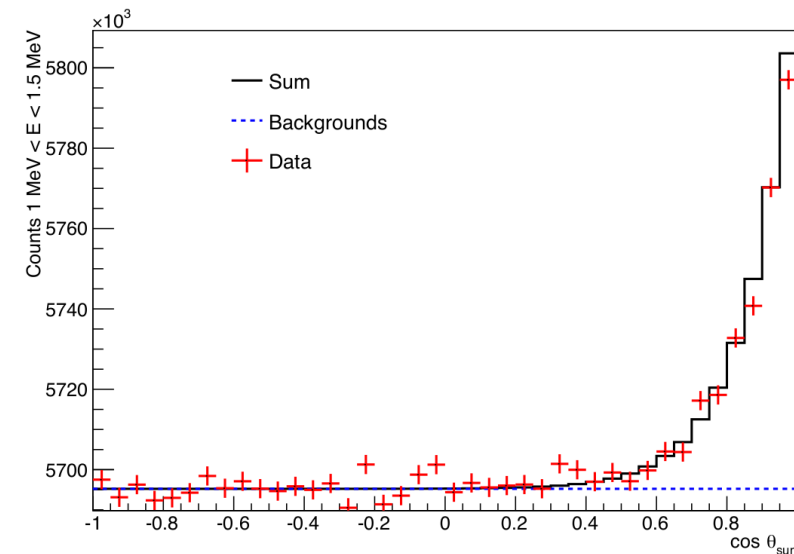
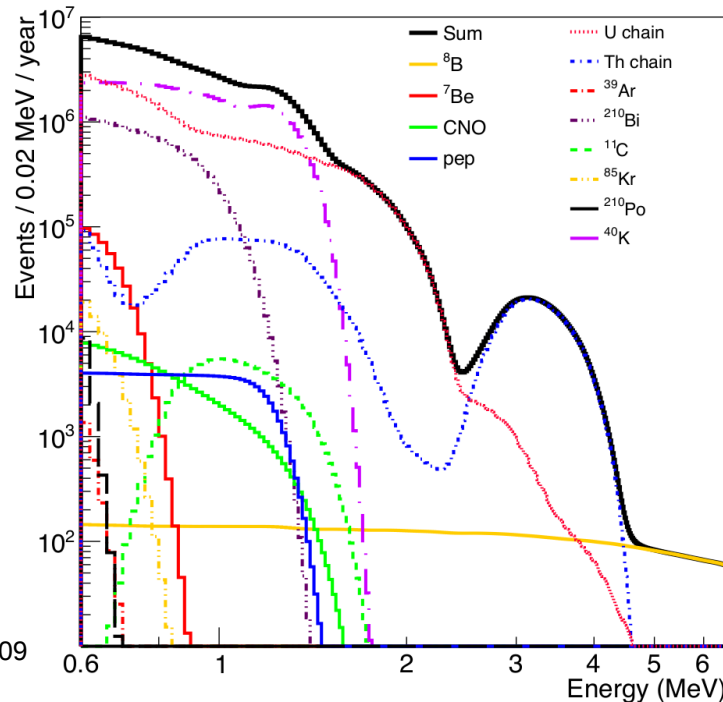
... and start probing the NH part with a higher isotope loading!

- The addition of scintillation dramatically **lowers the detection threshold**
 - Increases sensitivity to a wider range of solar ν 's
 - Keeps directional information with Cherenkov
- Measuring the CNO and pep components of the solar flux allows to:
 - Study **solar metallicity**
 - Study **neutrino oscillations** and matter effects

Nature, 512, pp. 383 (2014)

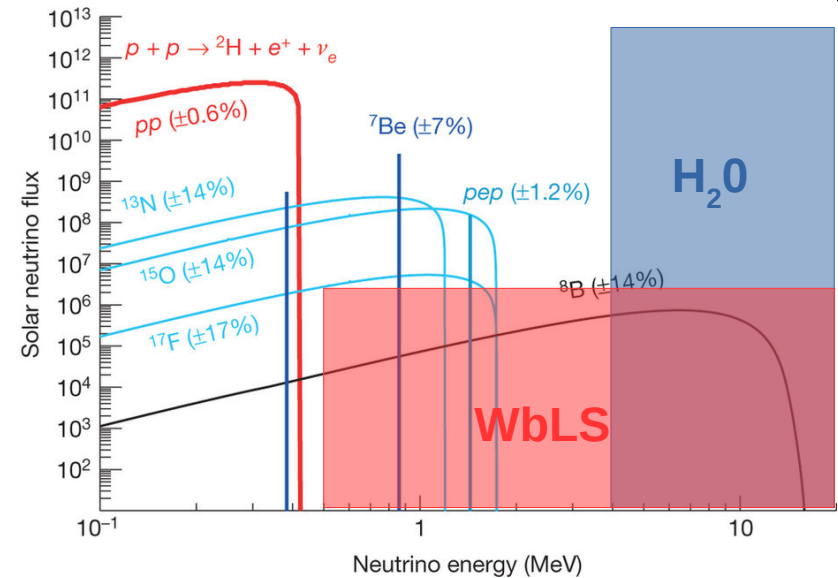


Predicted solar neutrino spectra for:
 - 50 kT Theia with 5% WbLS
 - 5 years
 - 90% coverage
 - 25° resolution
 - Borexino and SNO background levels



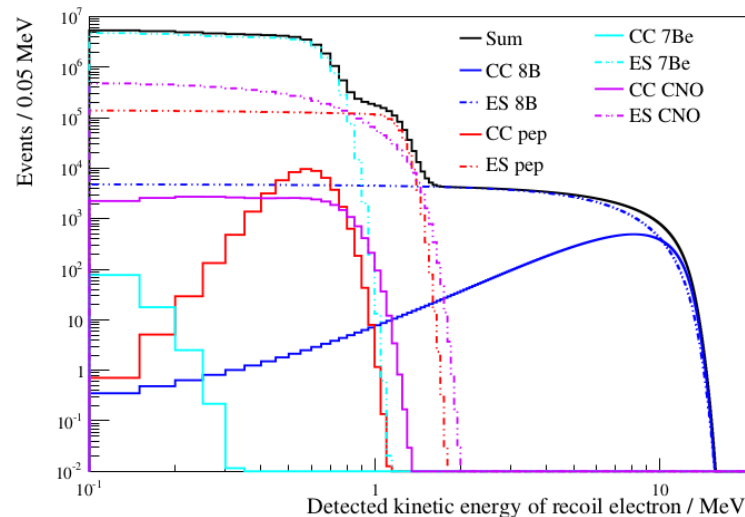
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- Measuring the CNO and pep components of the solar flux allows to:
 - Study **solar metallicity**
 - Study **neutrino oscillations** and matter effects
- Isotope loading (${}^7\text{Li}$, ${}^{71}\text{Ga}$, ${}^{37}\text{Cl}$) allows the detection of **charged current interactions** and not only elastic scatterings

Nature, 512, pp. 383 (2014)

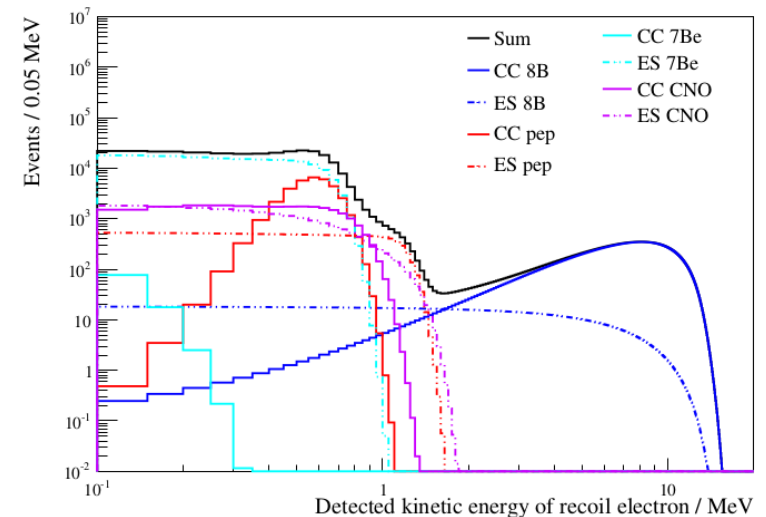


Predicted solar neutrino spectra in 30 kT Theia loaded with 1% ${}^7\text{Li}$

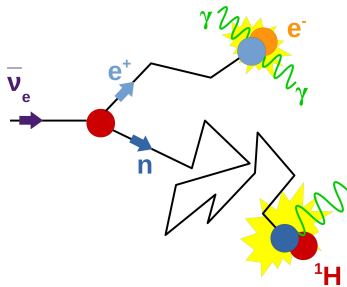
More details:
arXiv:1409.5864



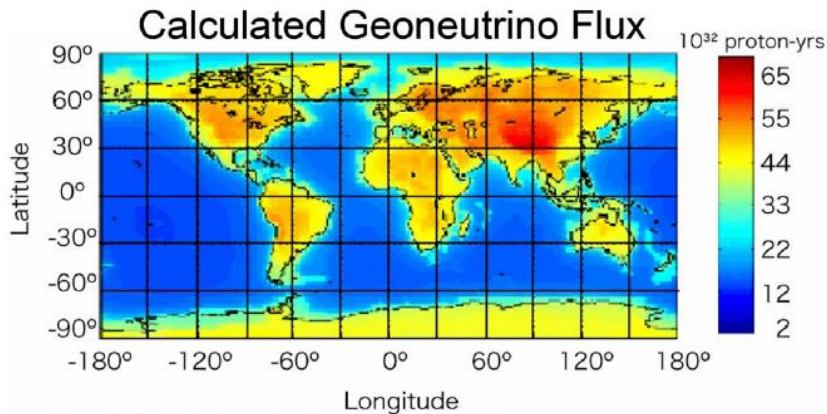
No directionality cut



Directionality cut (to cut ES events)



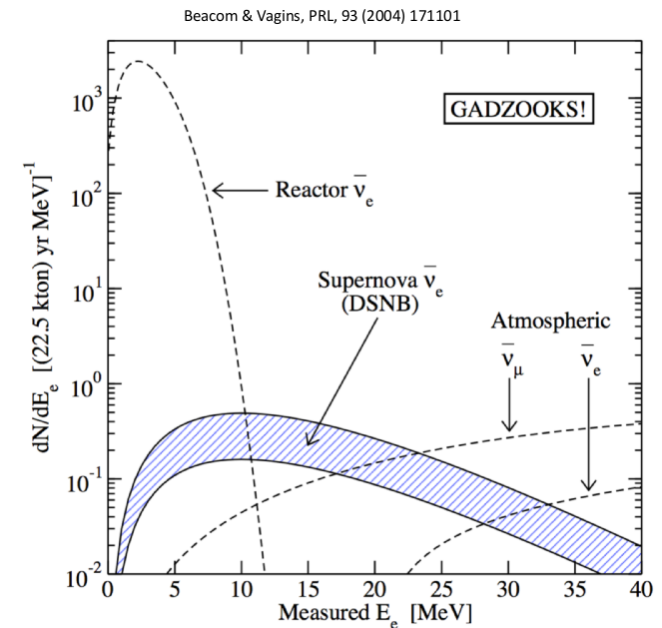
Antineutrinos detected through **Inverse Beta Decay** on H
→ Coincidence (e^+, n)



G. Fiorentini, M. Lissia, F. Mantovani, and R. Vannucci, hep-ph/0401085

- A total of ~100 geoneutrinos have been detected by Borexino and KamLAND so far
- Given its size and low threshold, Theia would detect **the same number in a few months**
- Beneficial isotope loading: **Gd, Li, B**

- ~8000 events for a 10 kpc SN
- Ability to tag IBD neutrons:
 - Selection of ES events and good angular reconstruction (~3°)
 - Crucial input for astronomers



- **Diffuse Supernova Background:**
 - About 1 event/10kt/year expected

- **Theia** is a large water-based liquid scintillator detector with a **impressively broad range of physics goals**
- **Water-based Liquid Scintillator** combines the **advantages of both water and liquid scintillator**: directionality, light yield, optical properties, cost, etc..
- **Long baseline** program is **complementary to DUNE's** and could be happening in parallel
- Theia would provide **world leading measurements** in a wide range of physics goals
- Theia benefits from other existing and/or funded experiments such as **ANNIE** and **WATCHMAN**

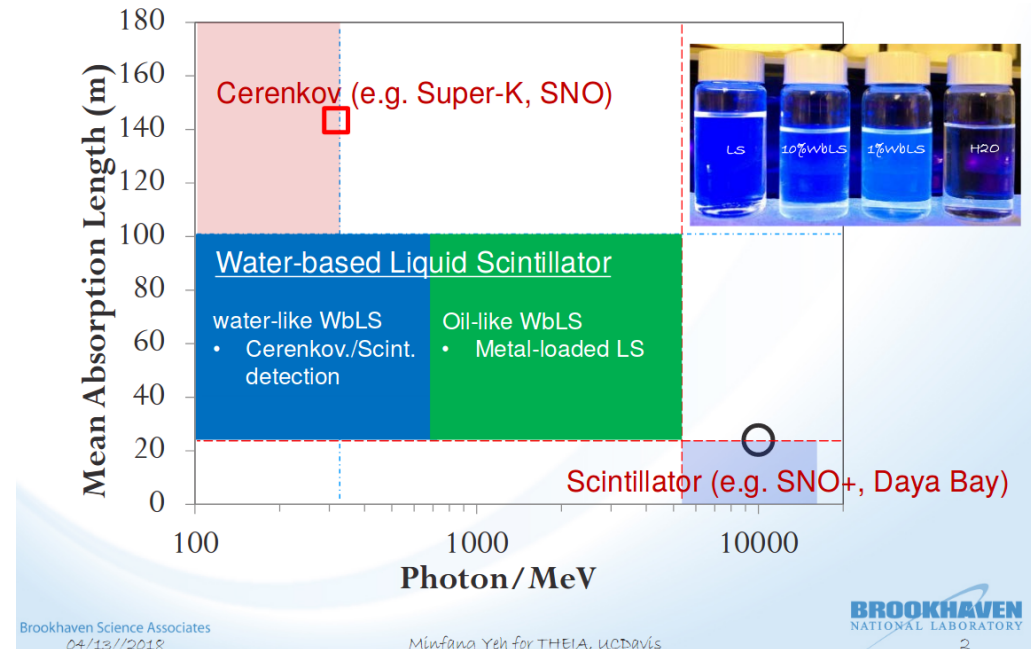


THANK YOU FOR YOUR ATTENTION

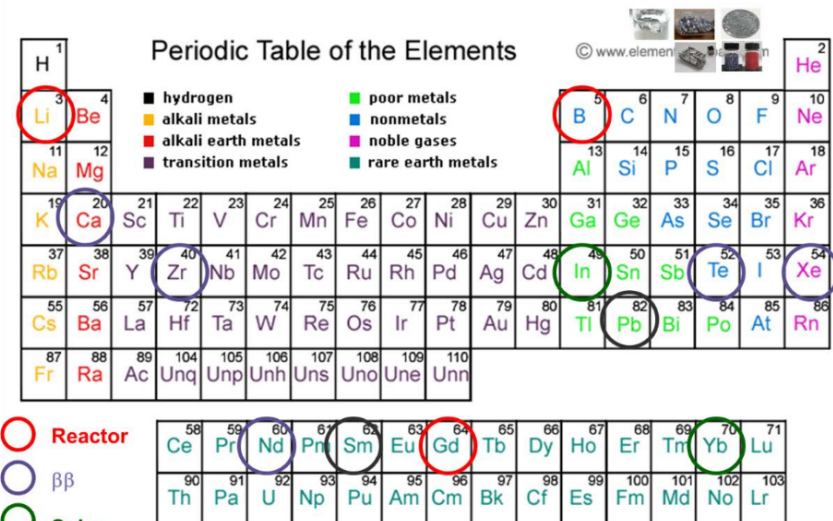
BACK-UP

Water-based Liquid Scintillator - Tunability

- Different applications might require different levels of scintillation and Cherenkov light:
 - **Low energy physics** (~MeV)
 - High LS content (> 20%) and **oil-like WbLS**
 - **Beam physics** (~GeV)
 - Lower LS content (< 20%) and **water-like WbLS**

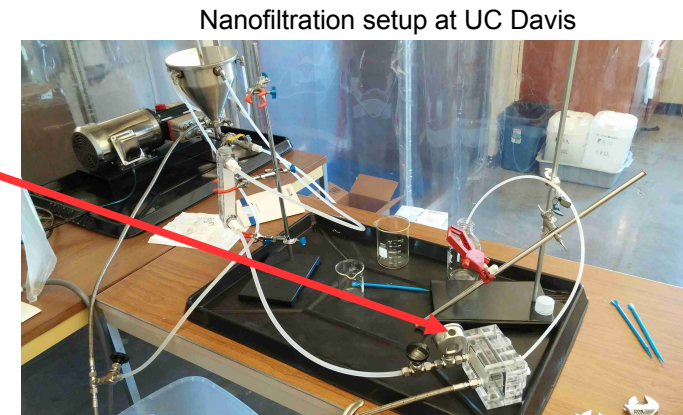
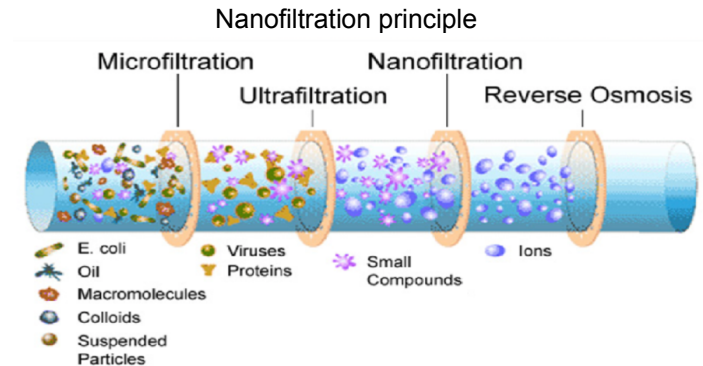


- **Isotope loading** (Li, Gd, Te, etc..) has been performed on both water and LS
- **Hydrophilic elements** would be the easiest to load (e.g. Gadolinium compounds)
- Metal-loaded WbLS would help a broad range of neutrino experiments:
 - **Reactor** (neutron sensitivity): Gd, Li, B
 - **Double-beta decay**: Te, Nd, Xe



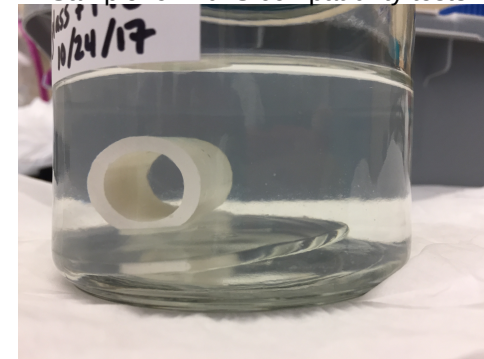
i.e. Daya Bay, PROSPECT, SNO+

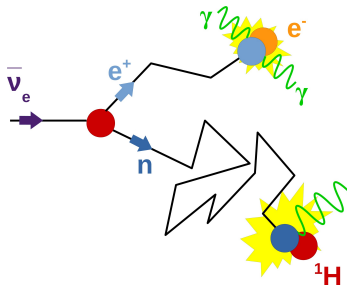
- **Nanofiltration:** Filtration process using membranes to separate water and organic components (micelles)
 - Widely used in the food industry
 - Allows water to go through the membrane **without damaging the micelles**
 - **Scalable** with reasonable flow rate
- Water will then be purified using a regular de-ionizing process
- **Work at UC Davis:**
 - Small-scale nanofiltration apparatus
 - Post-filtration absorbance and light yield measurements to ensure wbLS hasn't been damaged
- Since surfactant might be aggressive on some materials, all materials must undergo a **compatibility test beforehand**



Credit: **Leon Pickard** and **Julie He**
(UC Davis)

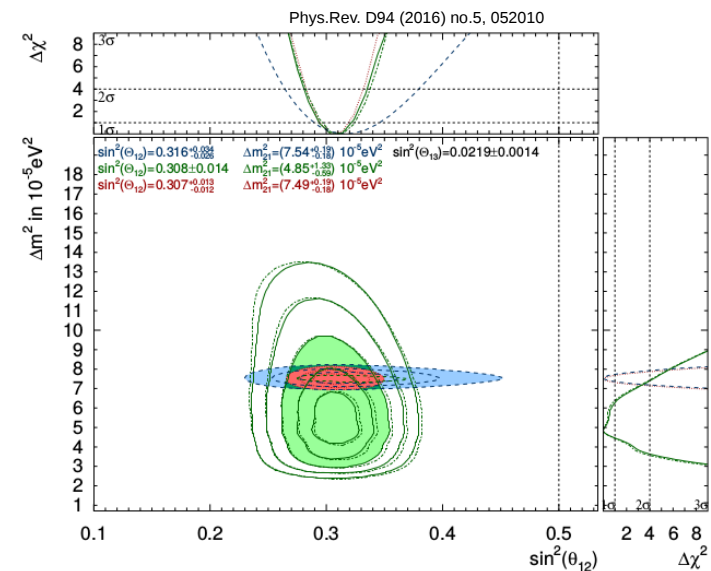
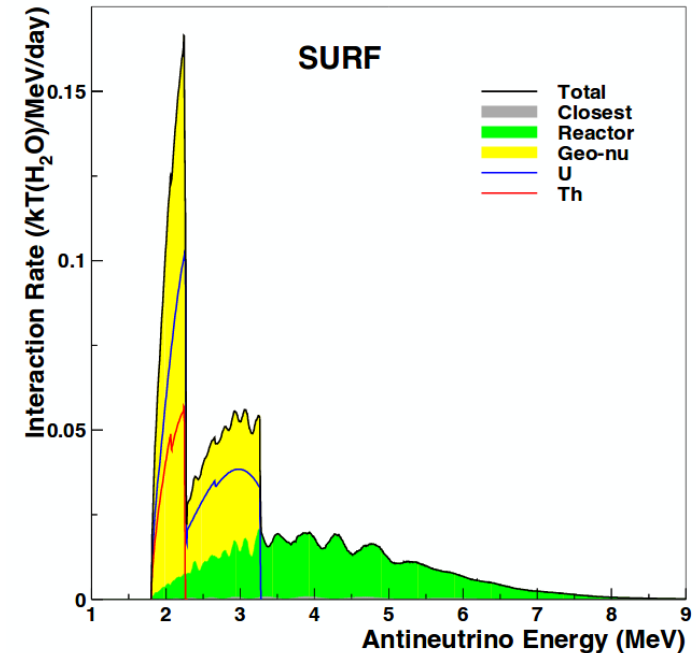
Sample for wbLS compatibility tests

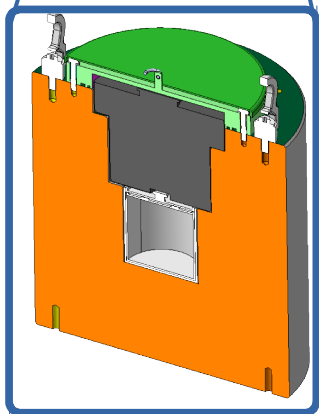
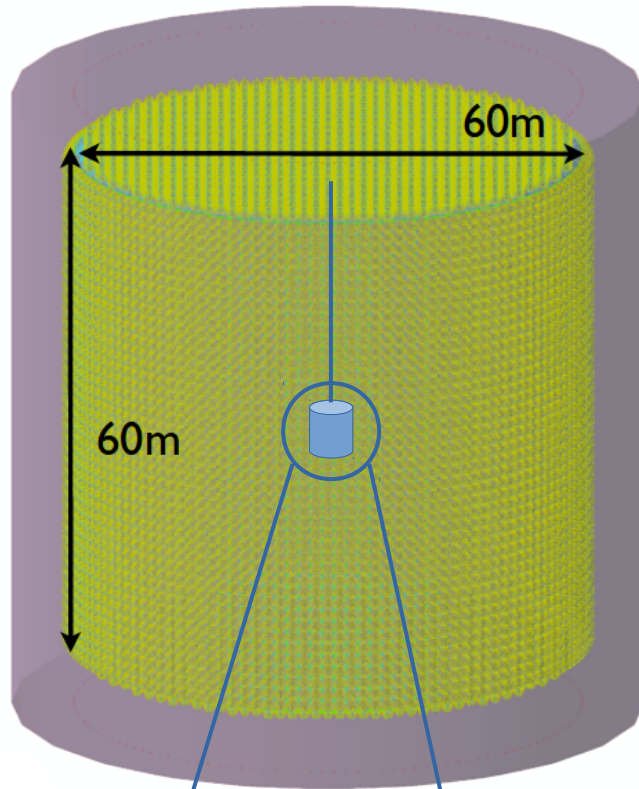




Antineutrinos detected through **Inverse Beta Decay** on H
→ Coincidence (e^+, n)

- About 22 reactor neutrinos (and 30 geoneutrinos) expected per kt per year at SURF
- 30 kt Theia → ~650 IBD reactor events per year
- Possibility to **measure Δm_{12}^2** using reactor neutrinos
- Comparison with Δm_{12}^2 measured using solar neutrinos **in the same detector**
- Theia could provide an answer to the **existing solar/reactor tension** (1.5σ) on Δm_{12}^2

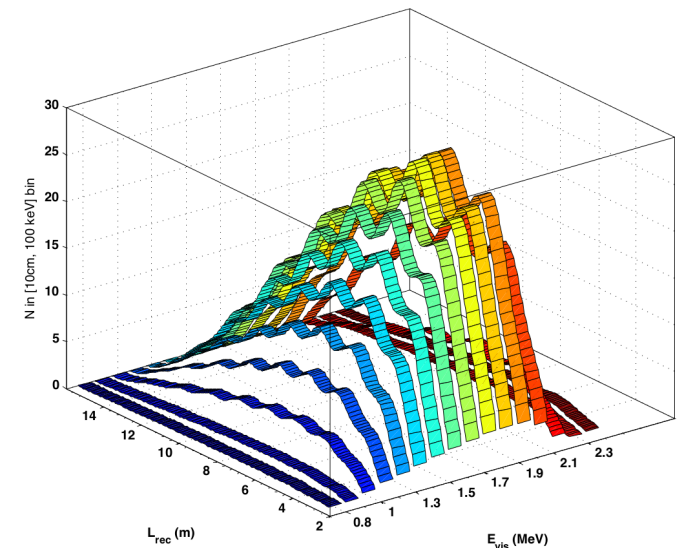




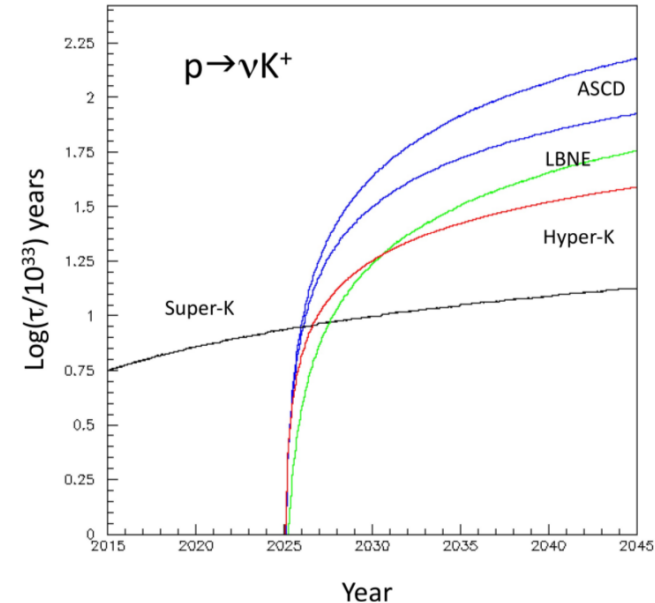
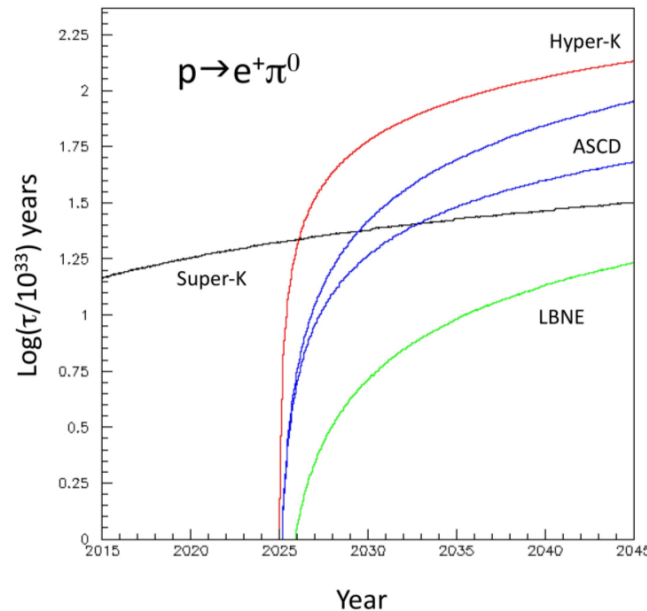
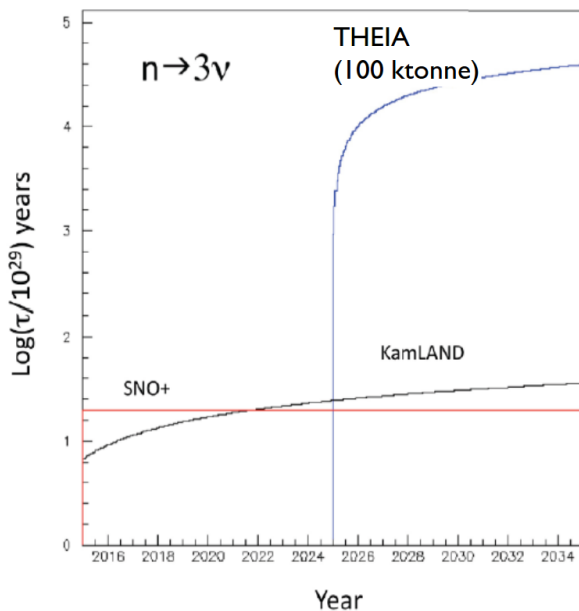
CeSOX source design:
-100 kCi ^{144}Ce - ^{144}Pr
-16 cm tungsten shield

- Hypothetical **eV-scale sterile neutrino** whose existence is accessible through **short baseline oscillations**
- Concept from **CeSOX/CeLAND** (Phys.Rev.Lett. 107 (2011) 201801):
 - Intense radioactive source in or next to Theia
 - Observation of a deficit and a spectral distortion as a function of distance
 - Requires good energy and position reconstruction
- **IsoDAR/DAEδALUS**: High intensity ^8Li decay at rest

Energy spectrum distortion as a function of distance in CeSOX



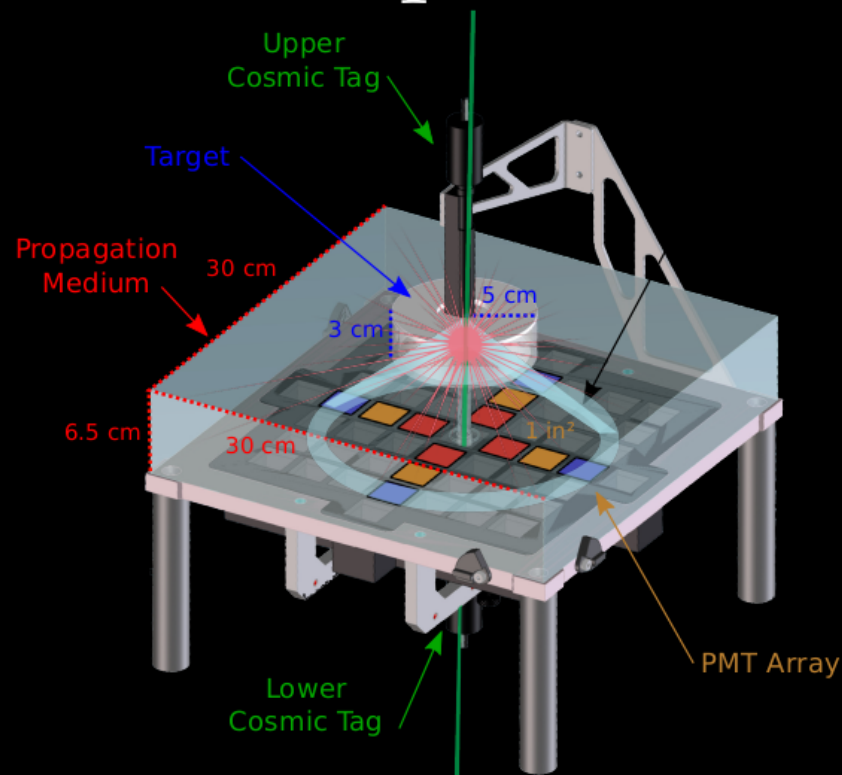
- **Huge size, deep location and scintillation light** make Theia an impressive nucleon decay detector
- Scintillation light allows the **observation of K^+** created upon a proton decay as well as the gammas emitted upon an **invisible neutron or proton decay** (~ 6 MeV)
- **Neutron tagging** enhances Theia's sensitivity for proton decay and can be further improved by **isotope loading**



CHESS: CHerenkov-Scintillation Separation

Berkeley group, PRC 95 055801 (2017)

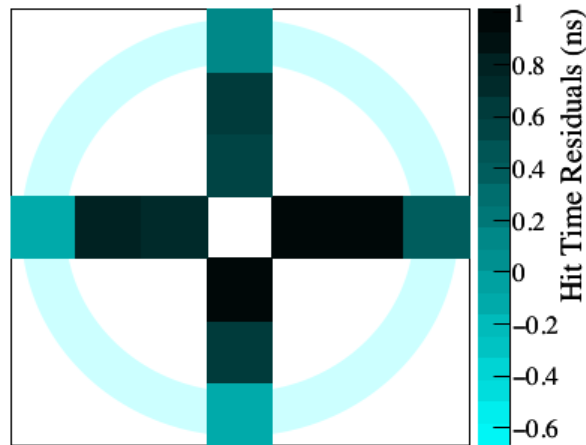
- Select vertical cosmic muon events
- Image Cherenkov ring in Q and T on fast-PMT array
- Allows charge- and time-based separation



12 1-inch H11934 PMTs (300ps FWHM, 42% QE)
CAEN V1742 (5GHz)
675 samples (135ns window)
CAEN V1730 (500MHz)

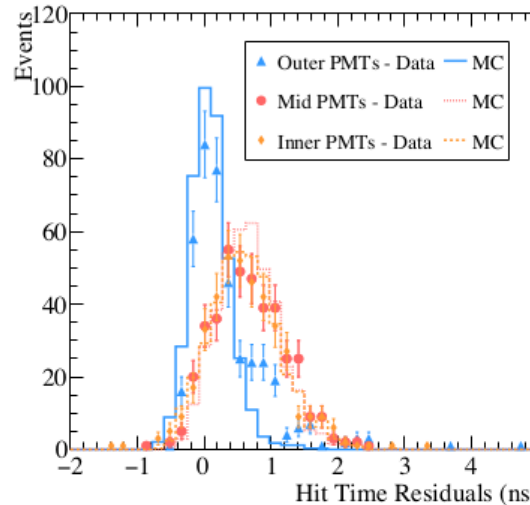
Berkeley group, *Eur. Phys. J. C* (2017) 77: 811

CHESS Results: LAB / PPO

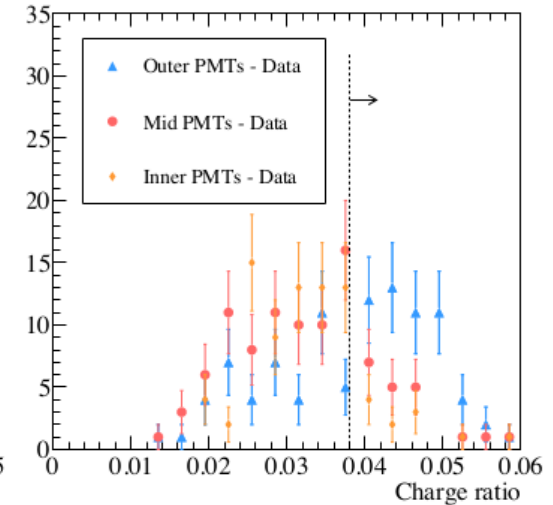


Typical ring candidate event

NOTE: Rise time = 0.75 ± 0.25 ns



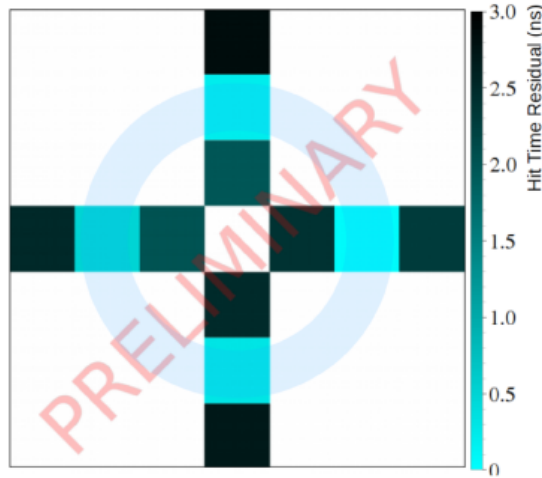
Time at fixed threshold
Corrected by ToF,
channel delays



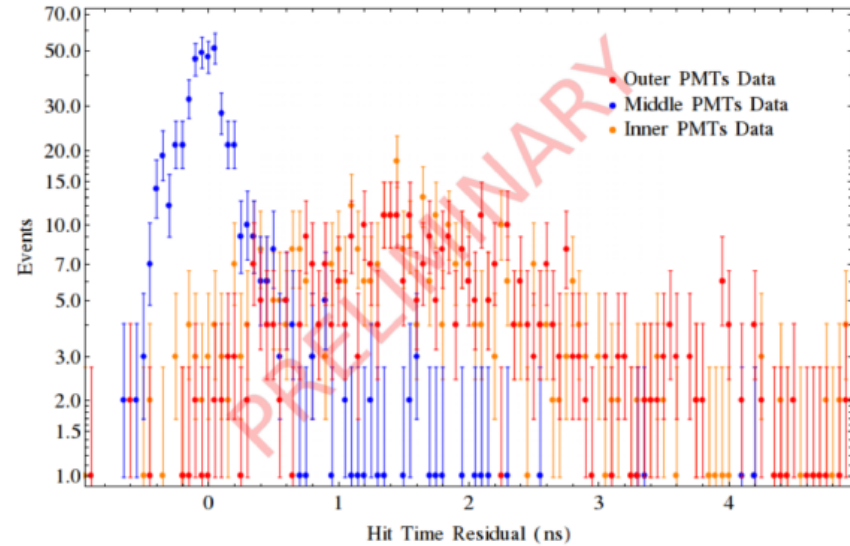
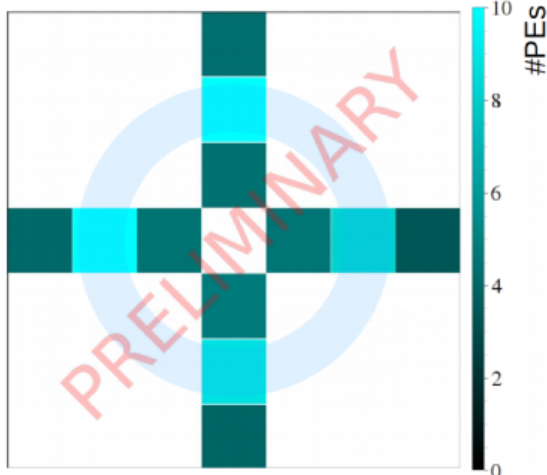
Ratio of charge in
prompt, 5ns window to
charge in total (135ns)
window

	LAB (time)	LAB (charge)	LAB/PPO (time)	LAB/PPO (charge)
Cherenkov detection efficiency	83 ± 3 %	96 ± 2 %	70 ± 3 %	63 ± 8 %
Scintillation contamination	11 ± 1 %	6 ± 3 %	36 ± 5 %	38 ± 4 %

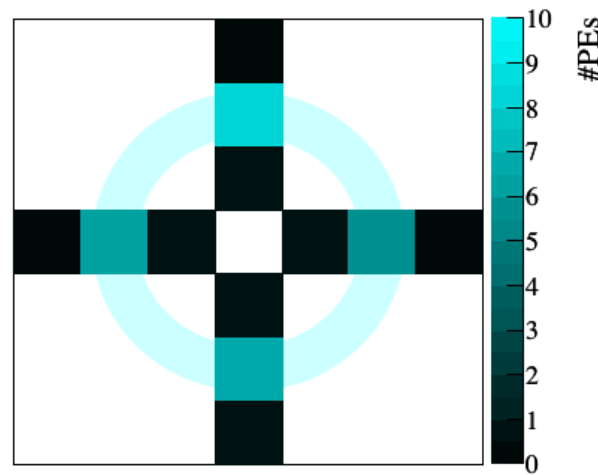
CHES Results: 1% WbLS



Average of WbLS data set



Average of water data set



Charge rings:
Clearly seeing
scintillation light
even at 1% LS fraction