



UNIVERSITY OF
SOUTH DAKOTA



Status and Initial Results of the MAJORANA DEMONSTRATOR

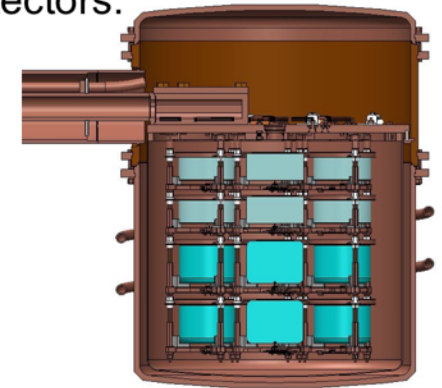
Wenqin Xu for the MAJORANA Collaboration
University of South Dakota
CIPANP 2018
Palm Springs, CA, May 29, 2018

The MAJORANA DEMONSTRATOR



Operating underground at the 4850' Sanford Underground Research Facility

- Demonstrating backgrounds low enough to justify building a tonne scale experiment.
- Goals:
- Establishing feasibility to construct & field modular arrays of Ge detectors.
 - Searching for additional physics beyond the standard model.



Energy resolution of 2.5 keV FWHM @ 2039 keV is the best of any $\beta\beta$ -decay experiment

Background Goal in the $0\nu\beta\beta$ peak after analysis cuts with the achieved resolution:
2.5 counts/(FWHM t yr)

- Projected backgrounds based on assay results ≤ 2.2 counts/(FWHM t yr)

44.1-kg of Ge detectors

- 29.7 kg of 88% enriched ^{76}Ge crystals
- 14.4 kg of $^{\text{nat}}\text{Ge}$
- Detector Technology: P-type, point-contact.

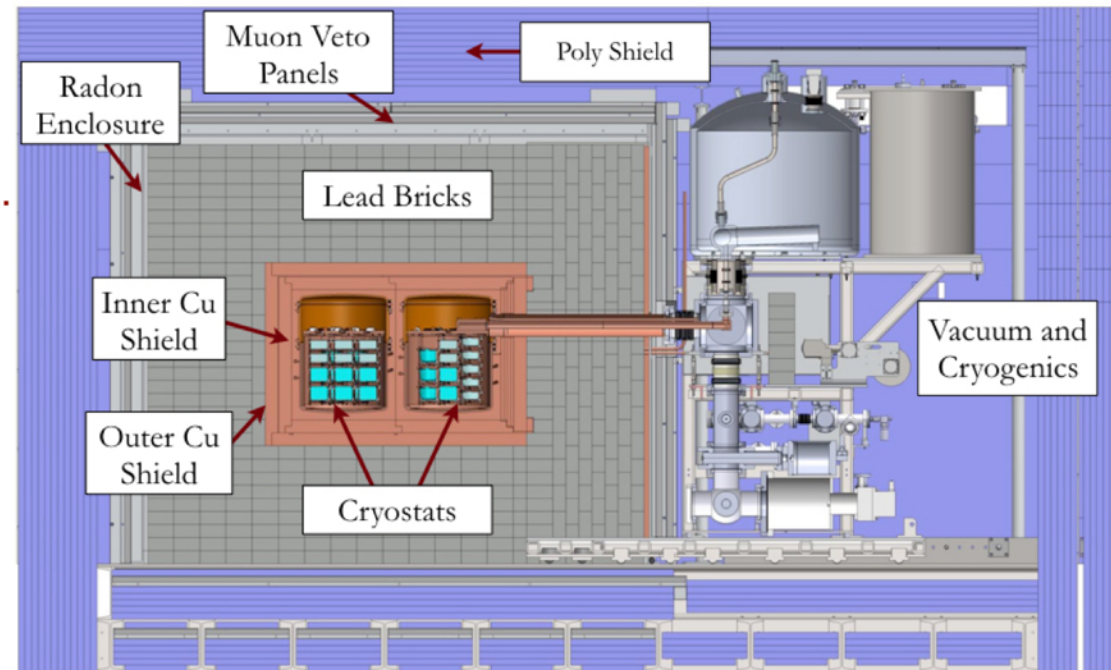
2 independent cryostats

- Ultra-clean, electroformed Cu
- 22 kg of detectors per cryostat
- Naturally scalable

Compact Shield

- Low-background passive Cu and Pb shield with active muon veto

Funded by DOE Office of Nuclear Physics, NSF Particle Astrophysics, NSF Nuclear Physics with additional contributions from international collaborators.

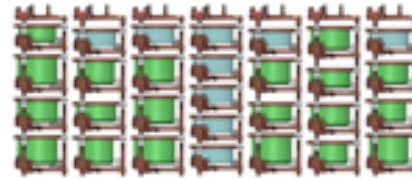


[N. Abgrall et al. Adv. High Energy Phys **2014**, 365432 (2014)]

DEMONSTRATOR Implementation

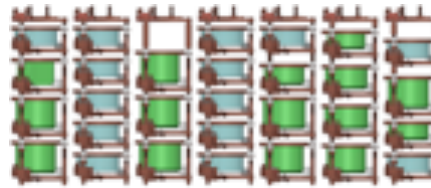


Module 1: **16.8 kg (20) ^{enr}Ge**
 5.6 kg (9) ^{nat}Ge

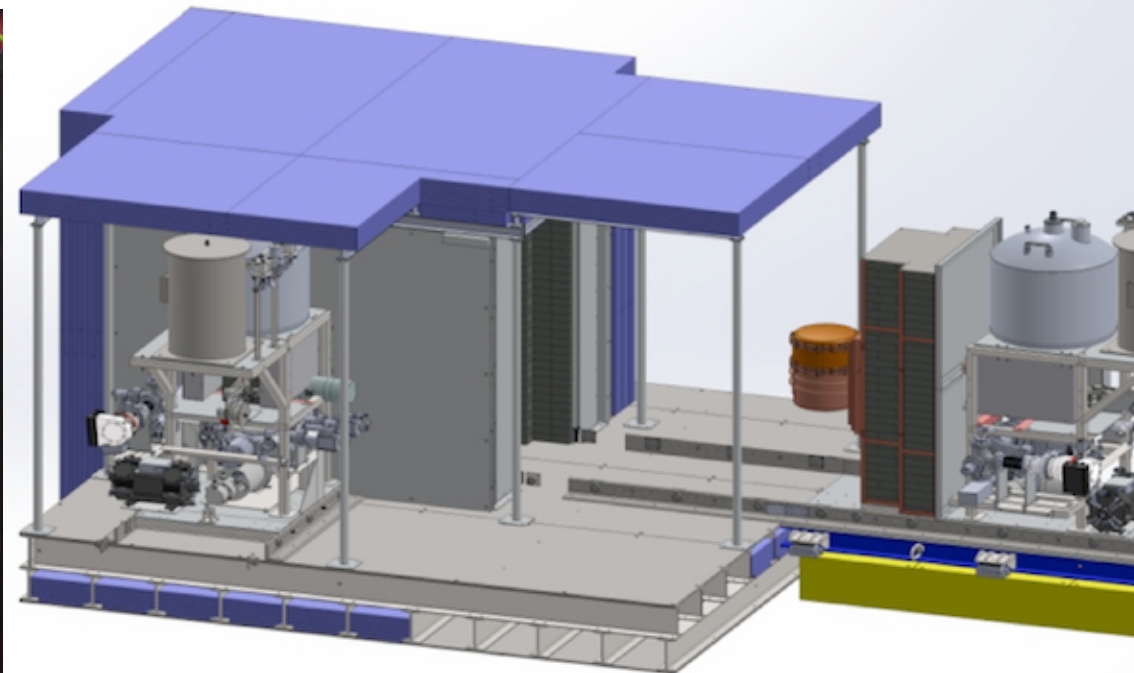


In shield Operation
May – Oct. 2015,
Final Installation,
Dec. 2015 — ongoing

Module 2: **12.9 kg (15) ^{enr}Ge**
 8.8 kg (14) ^{nat}Ge

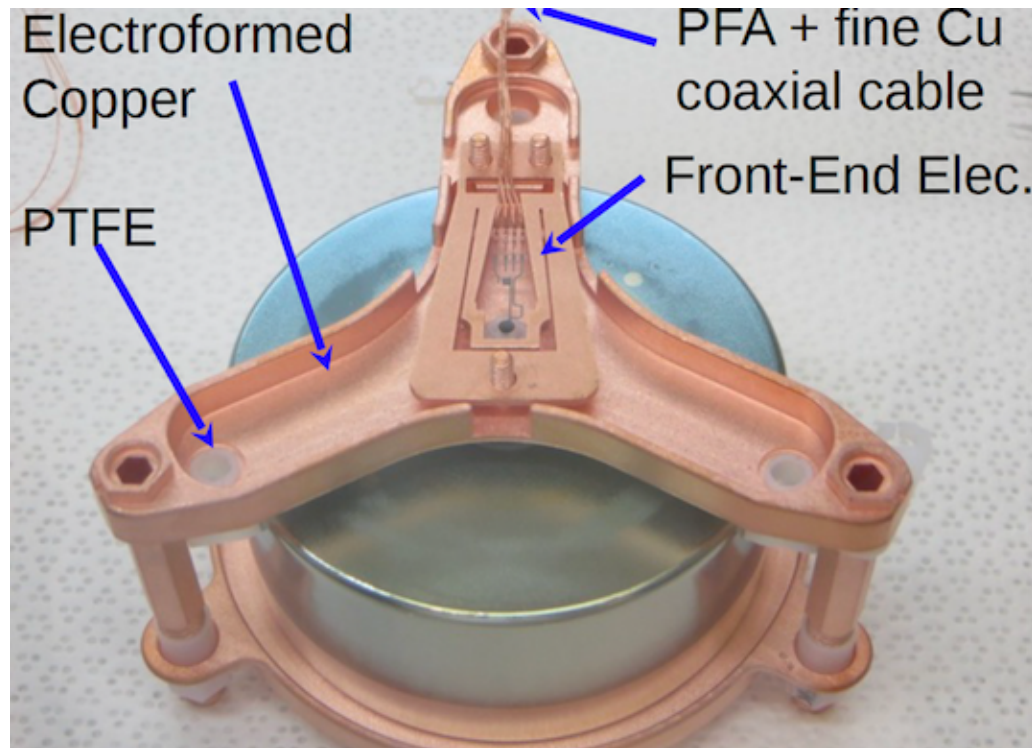


July 2016 — ongoing

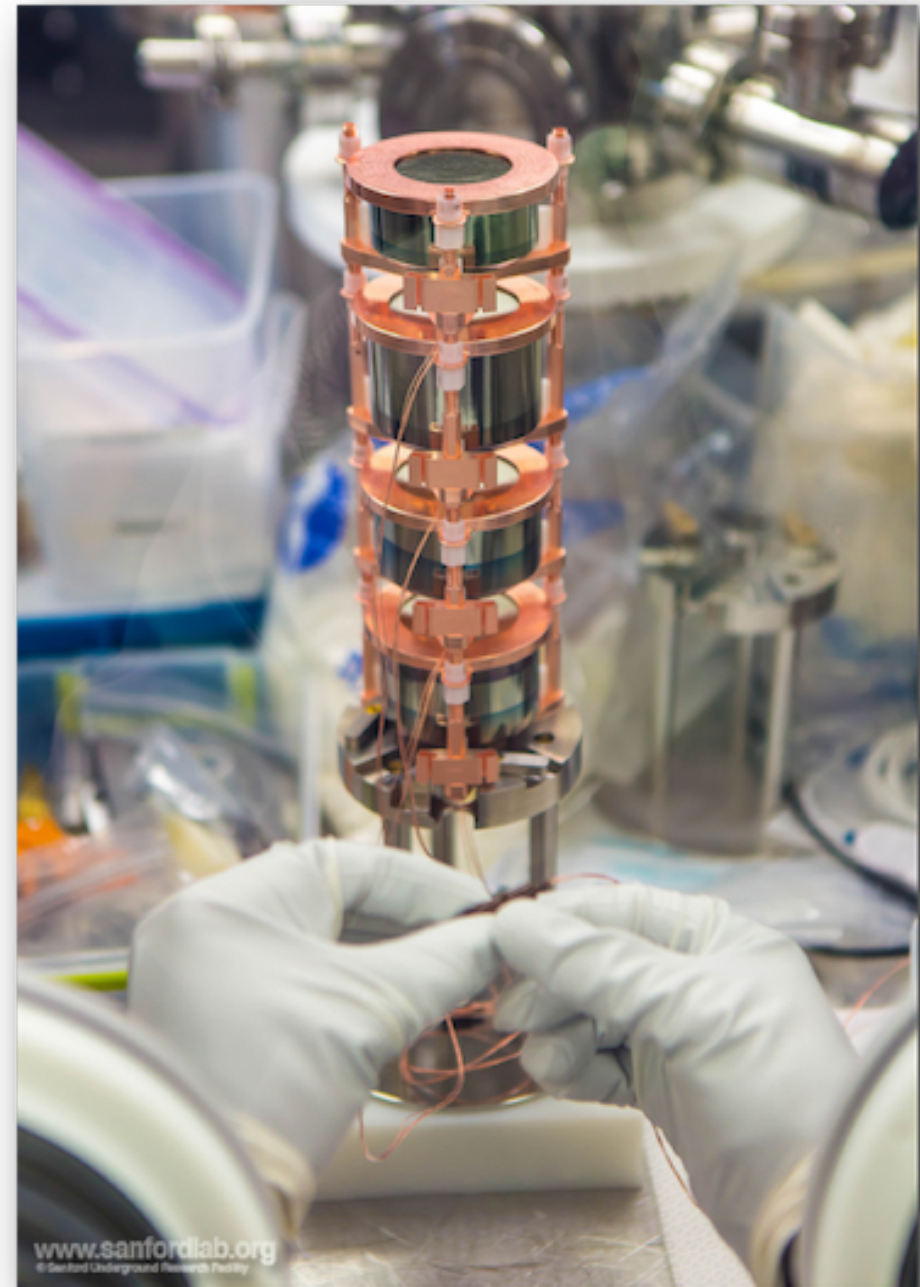




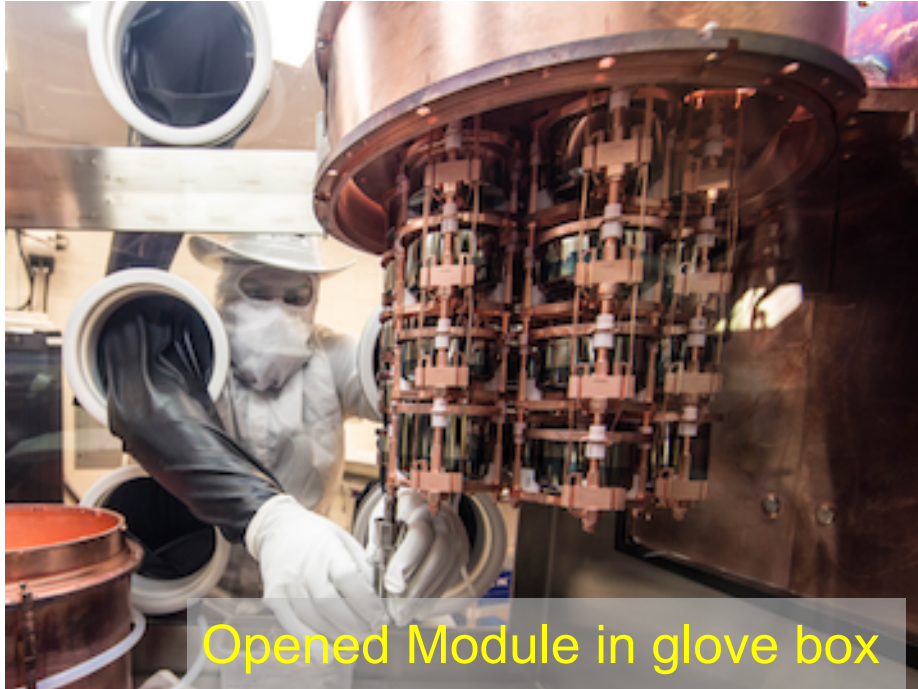
The DEMONSTRATOR Detectors & Strings



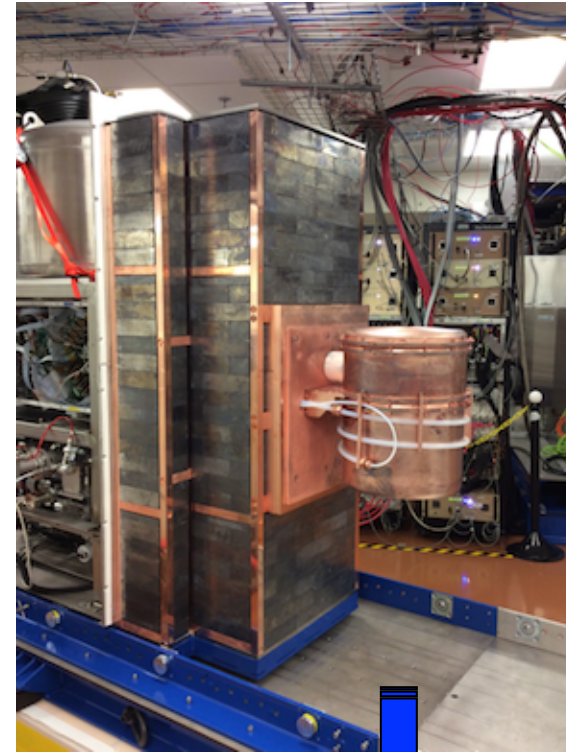
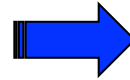
- AMTEK (ORTEC) fabricated enriched detectors
- 35 enriched point contact detectors (29.7 kg), 88% ^{76}Ge
- 33 Canberra modified natural BEGe detectors (20 kg)
- Tracked and minimized surface exposure of enriched material to determine cosmogenic activation



The DEMONSTRATOR Strings & Modules



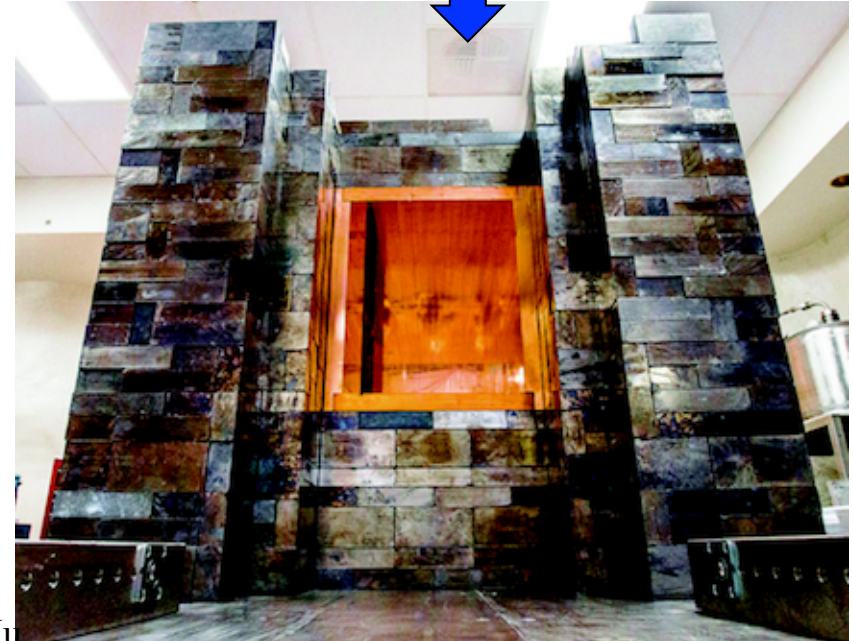
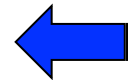
Opened Module in glove box



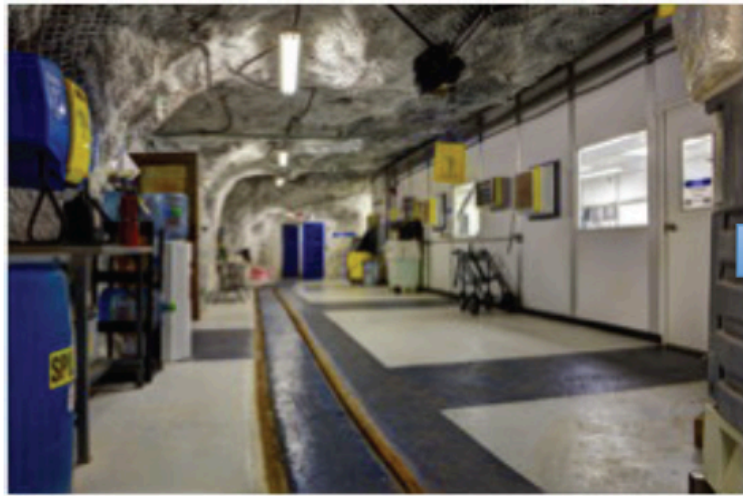
Transporting a closed module from the glove box into shield



Closed Module in shield



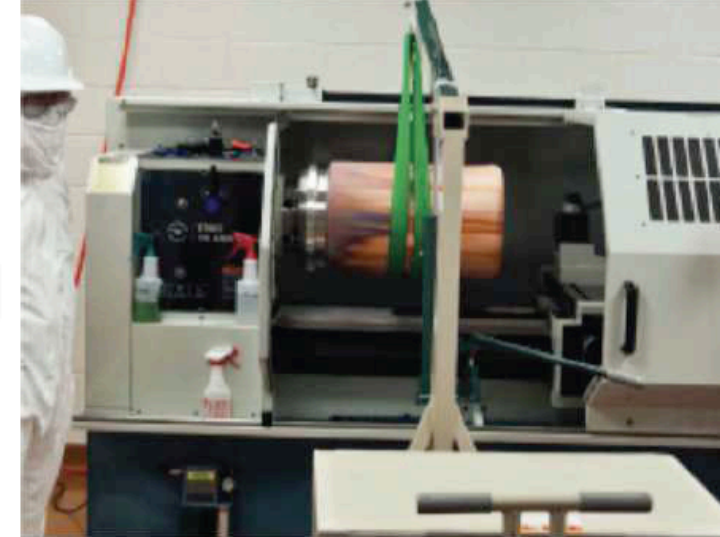
Electroformed Cu



10 baths



Max thickness 1.40cm



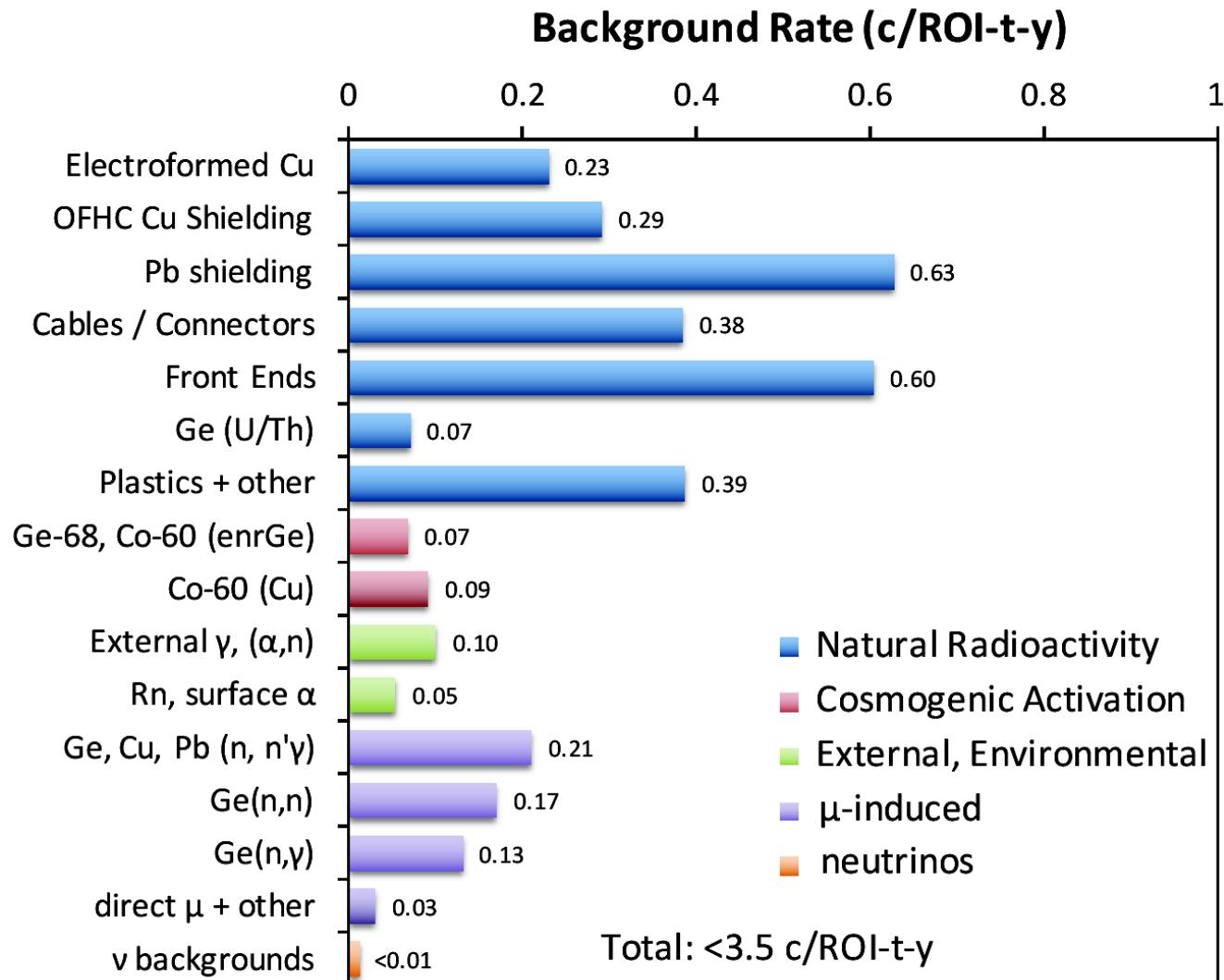
- Electroformed underground
- Average Th decay chain $\leq 0.1 \mu\text{Bq/kg}$, Average U decay chain $\leq 0.1 \mu\text{Bq/kg}$
- ~1.1 tons used in the DEMONSTRATOR
 - String components
 - Cryostats/thermosyphon
 - Inner layers of shielding

DEMONSTRATOR Backgrounds



Based on assays of materials; When upper limit, use upper limit value as contribution

(NIMA 828 (2016) 22)



Data Sets and Duty Cycles



M1 Commissioning (No Inner Shield)
June 27 - Oct. 7, 2015

M1 (With Inner Shield)
Dec. 31, 2015 – May 24, 2016

M1 (DAQ Multisampling)
May 24 – July 14, 2016

Module 1 (M1 & M2 in-shield)
Aug. 25 – Sept. 27, 2016

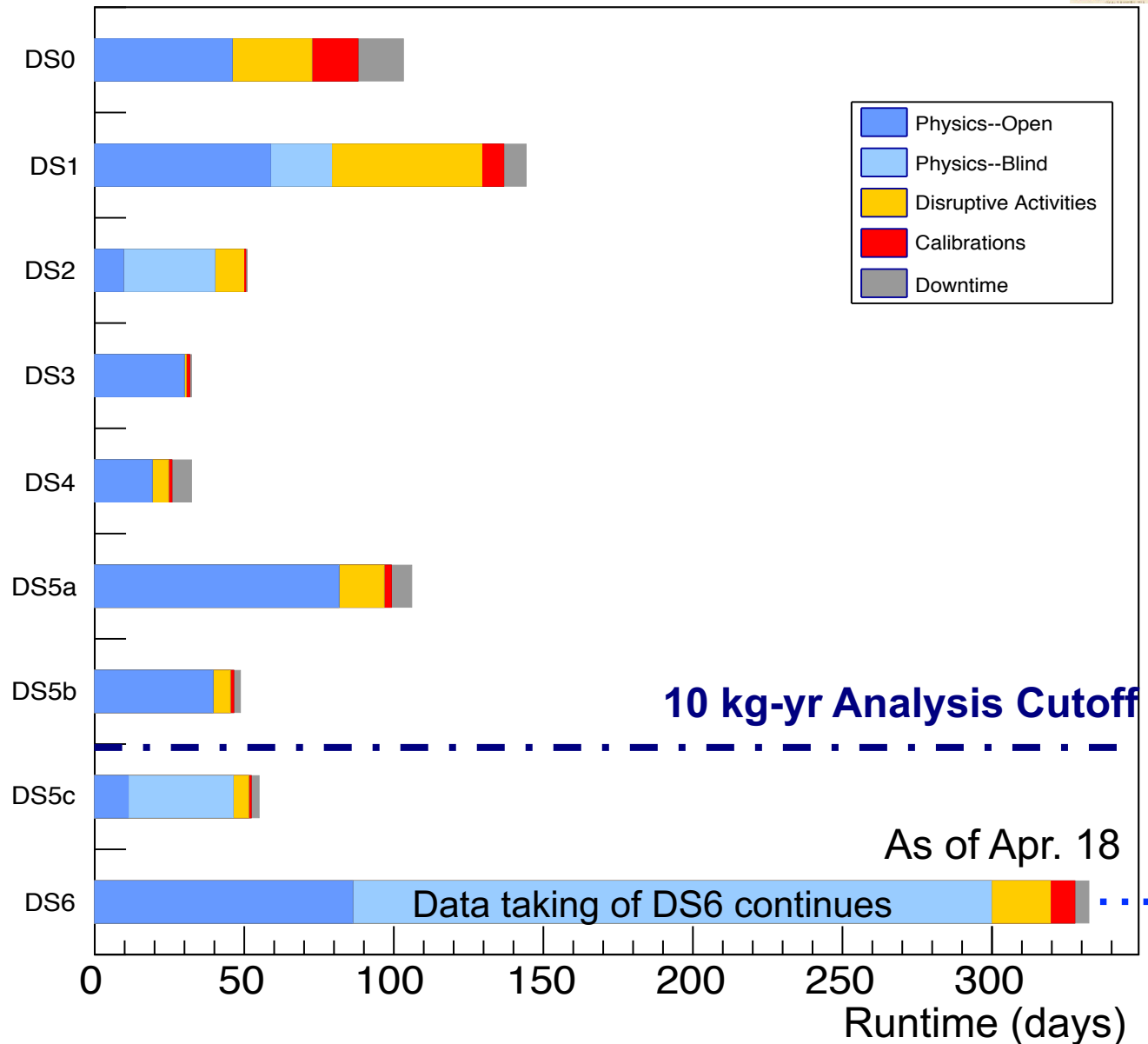
Module 2 (M1 & M2 in-shield)
Aug. 25 – Sept. 27, 2016

M1 & M2 (Integrate DAQ; high noise)
Oct. 13, 2016 – Jan. 27, 2017

M1 & M2 (Optimized Grounding)
Jan. 27 – Mar. 17, 2017

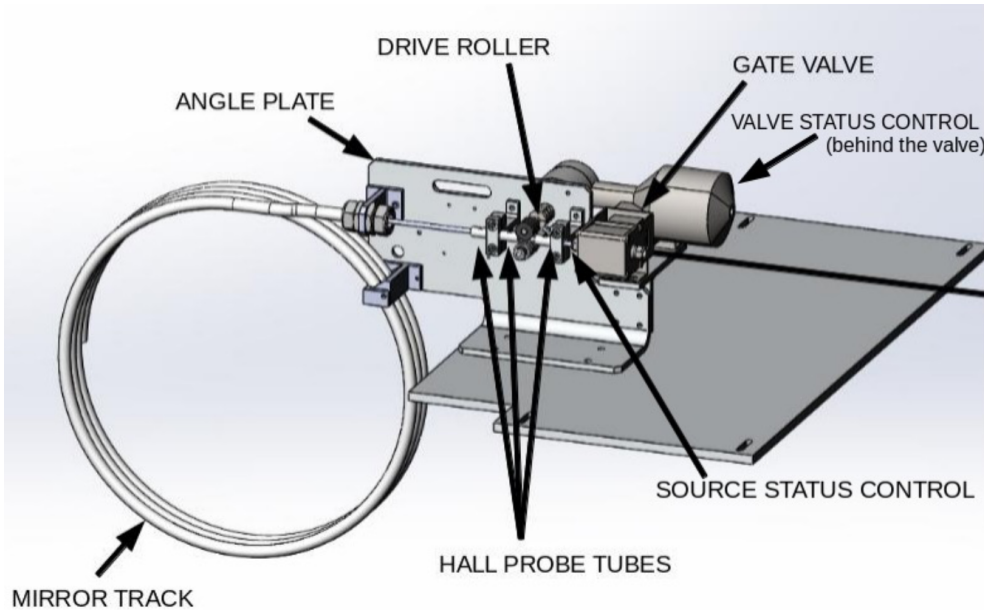
M1 & M2 (Blindness Implemented)
Mar. 17 – May 11, 2017

M1 & M2 (DAQ Multisampling)
May 11, 2017 - Present



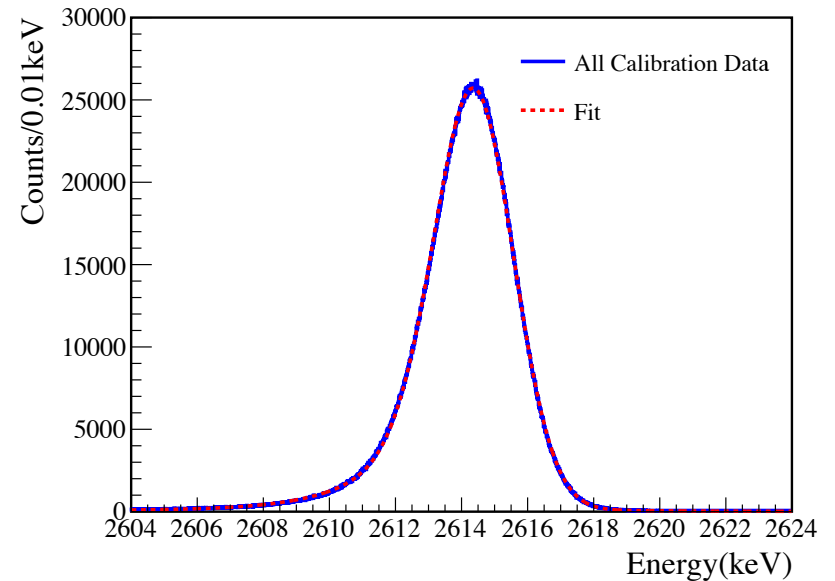
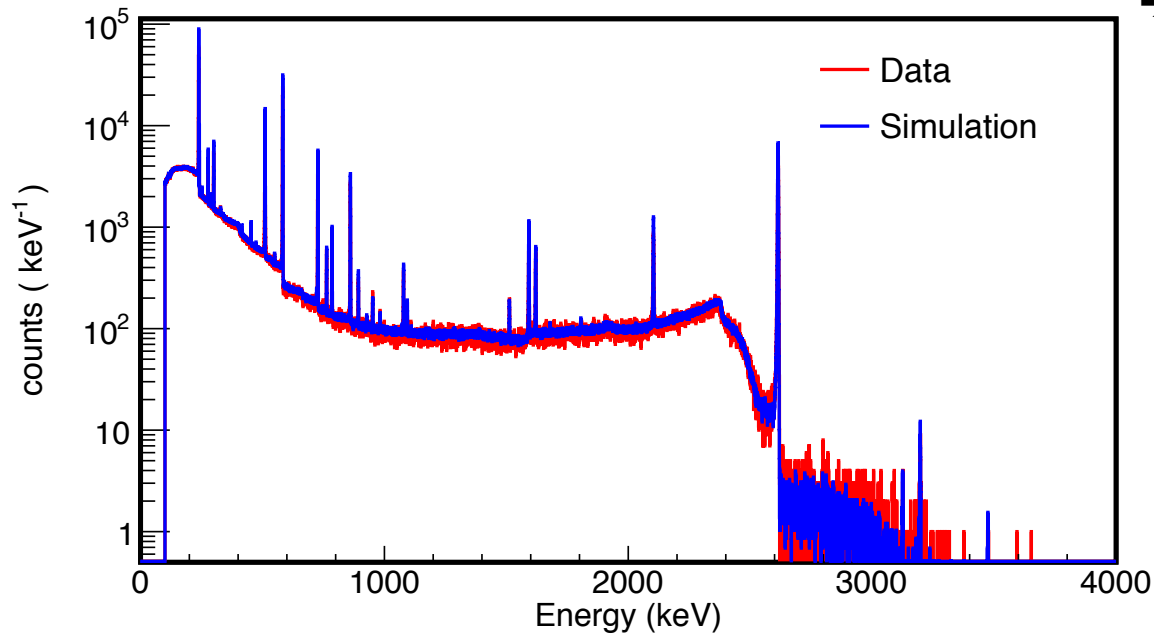
New results based on all data up to Apr 15th 2018 under preparation and to be released soon.
Total blind + open physics data ~ 26 kg-yr

Energy Calibration



Calibration System Paper NIMA, 16-22. **872**. 2017.

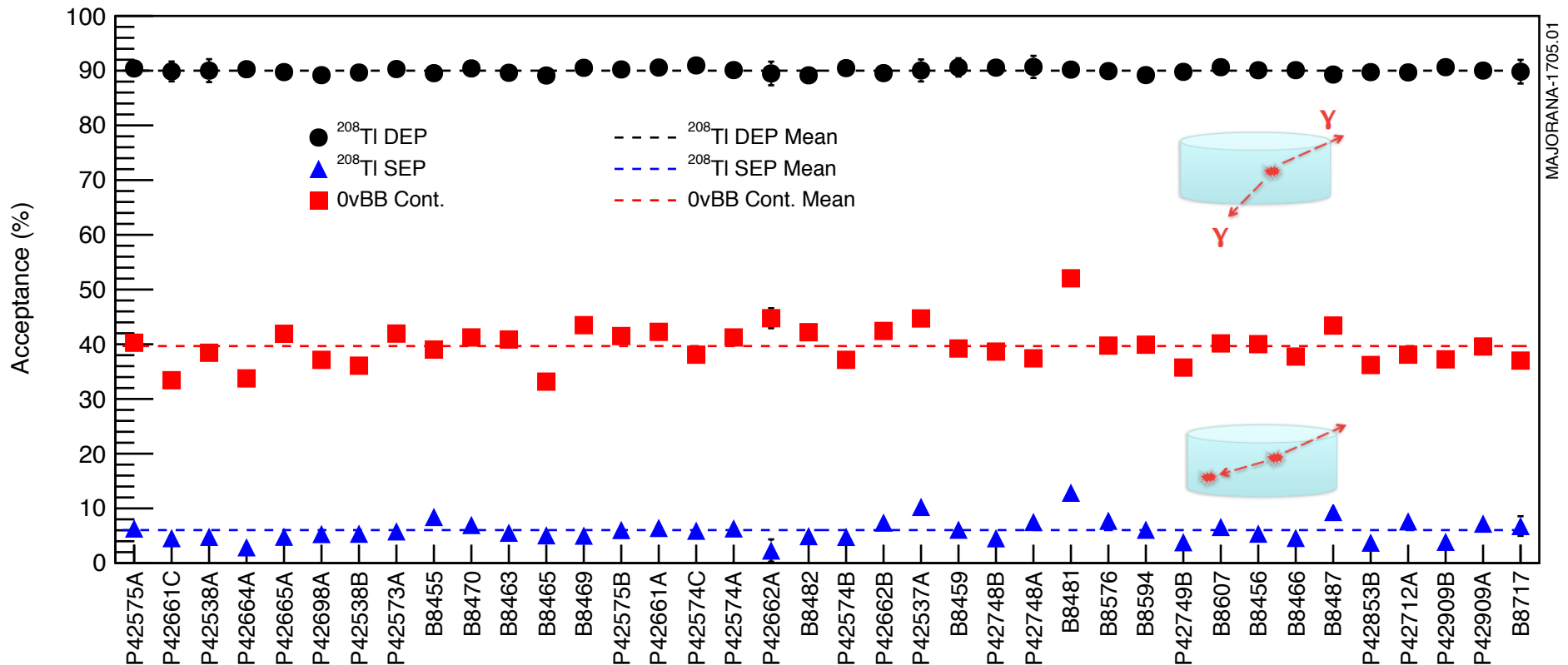
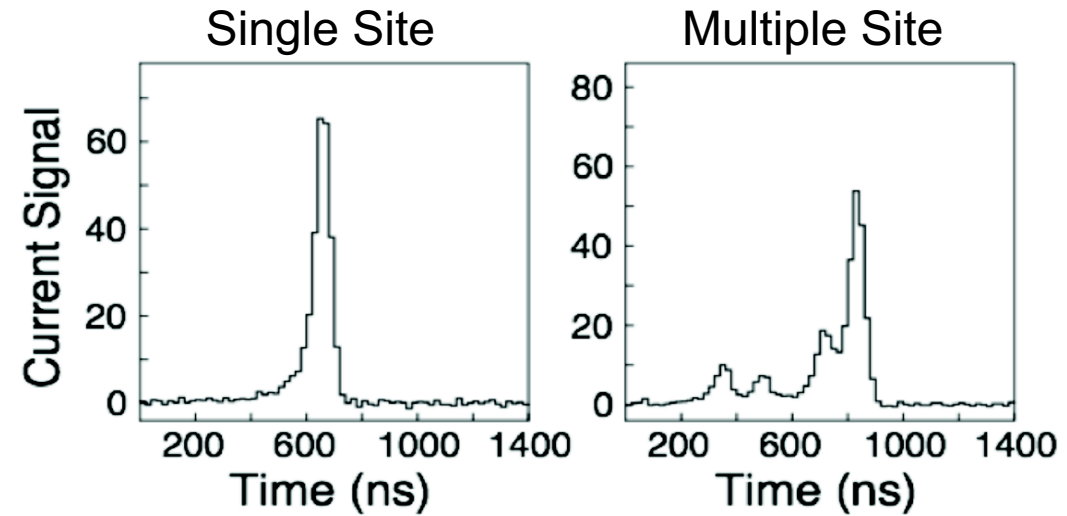
~3 keV FWHM at 2614 keV, approaching 0.1%
→ 2.5 keV FWHM at Q-value



Multiple Site Event Rejection



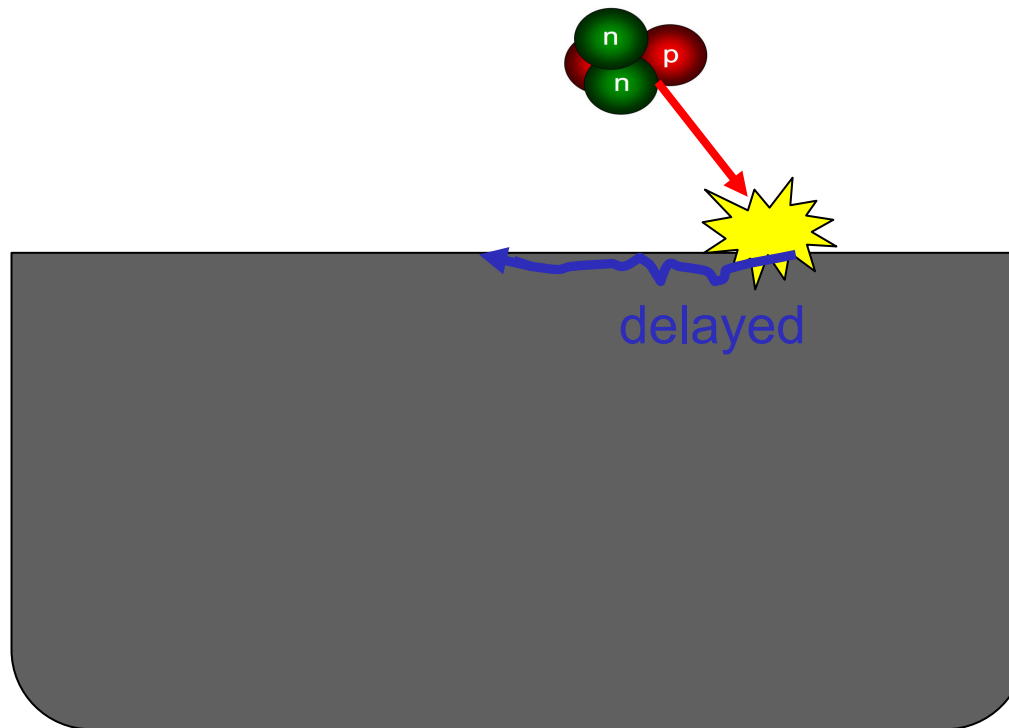
- $0\nu\beta\beta$ decays occur at a single site in the Ge crystal
- Point-contact detectors have sufficient differences in drift times throughout the bulk to identify multiple site interactions
- Tune max current amplitude-to-energy ratio (A_{vsE}) to ^{208}Tl calibrations to accept 90% of single site double escape events



Alpha Backgrounds



- Energy degraded alpha background observed
- Charge from these events drifts along the surface rather than through the bulk

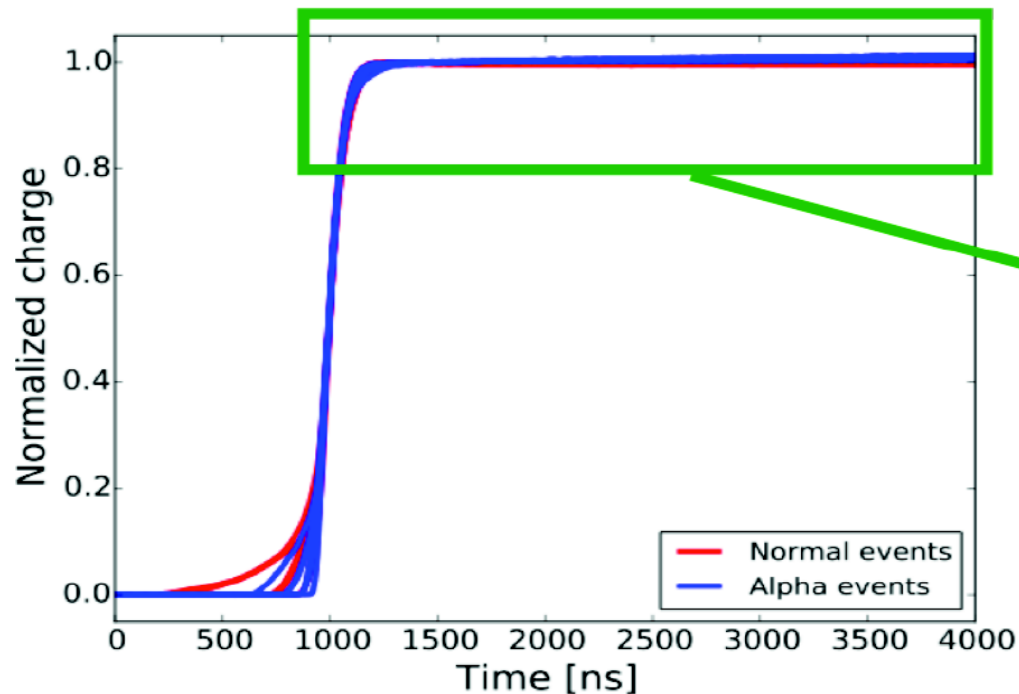


Alpha Backgrounds

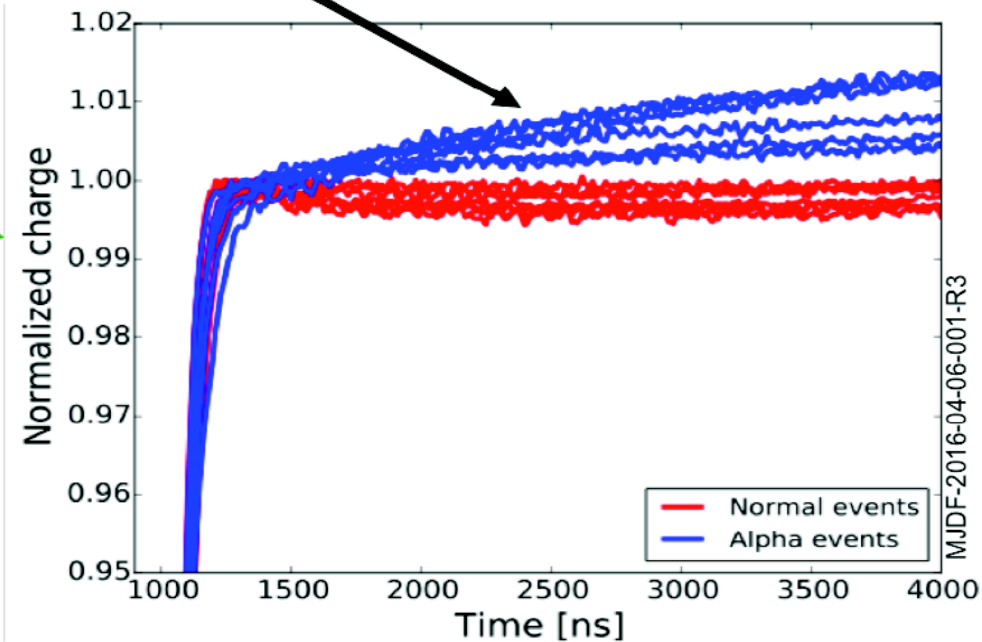


- Energy degraded alpha background observed
- Charge from these events drifts along the surface rather than through the bulk
- Results in a distinctive delayed charge recovery (DCR) signal which is used to efficiently cut alpha events based on the slope past the rising edge

Example pole-zero corrected waveforms



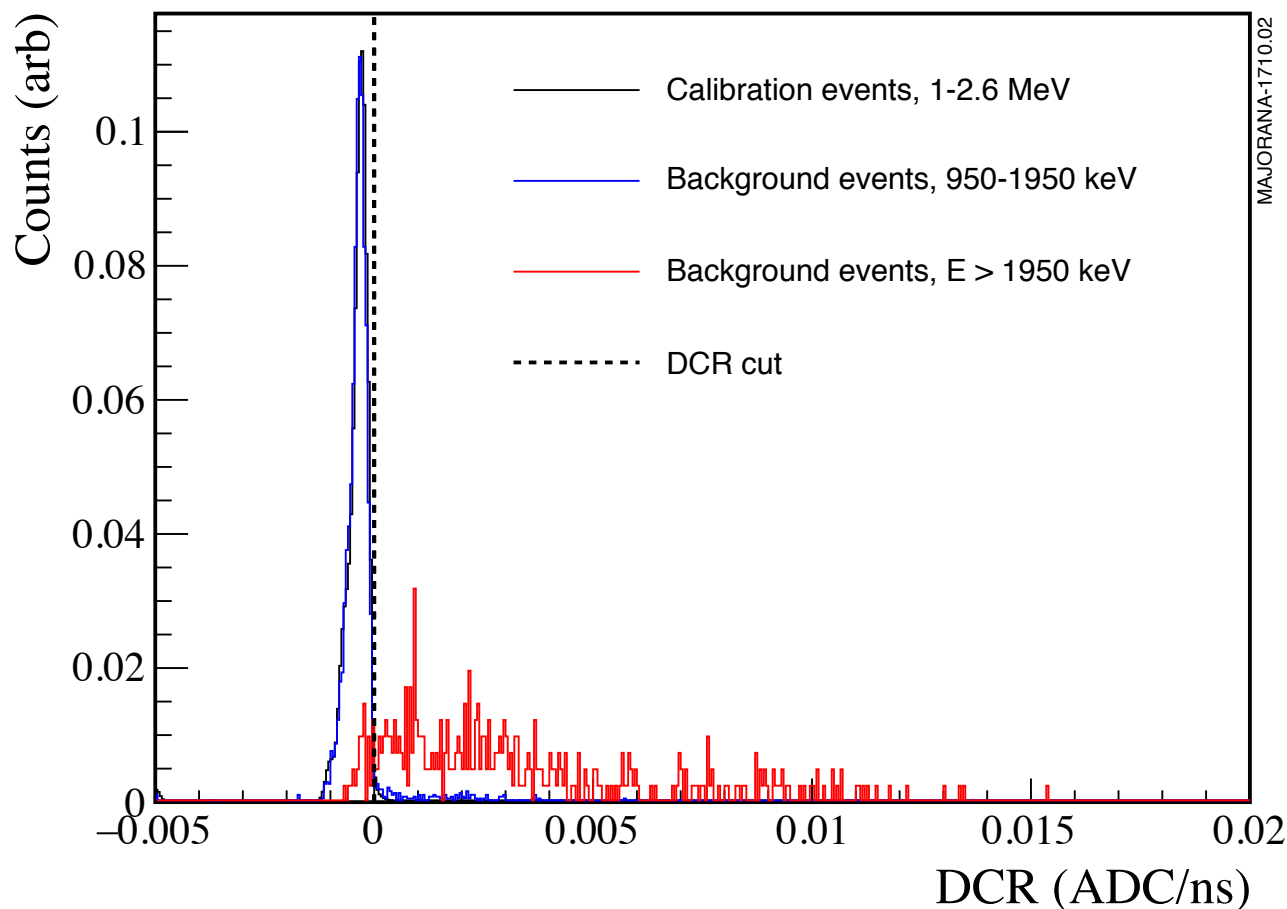
Slow drift of charges along passivated surface results in very slow signal component



Alpha Backgrounds



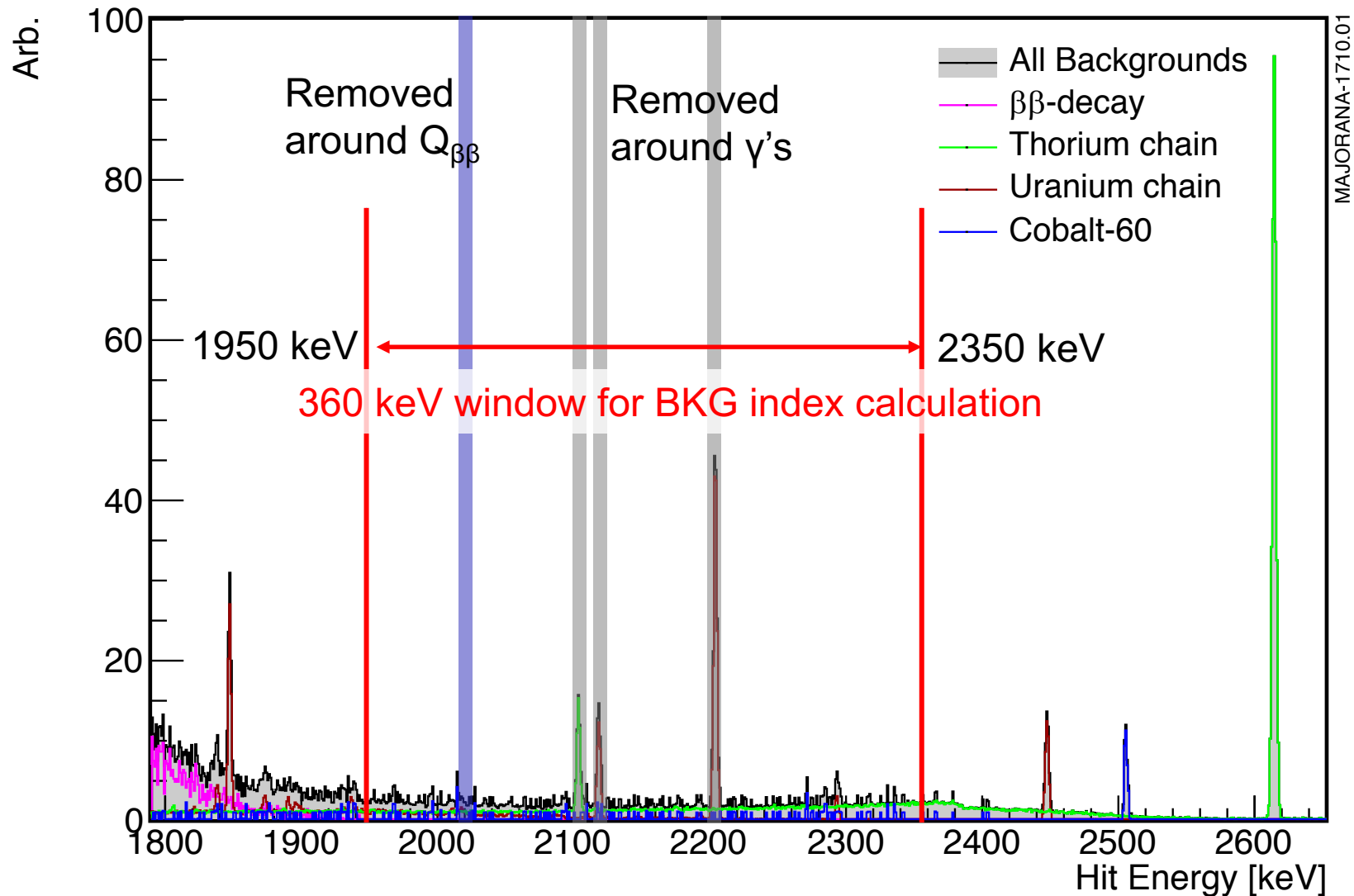
- Energy degraded alpha background observed
- Charge from these events drifts along the surface rather than through the bulk
- Results in a distinctive delayed charge recovery (DCR) signal which is used to efficiently cut alpha events based on the slope past the rising edge.
- DCR cut tuned to keep 99% of the photons in calibration
- Surface alpha response has been characterized using a MAJORANA detector in the TUBE alpha scanner at Technical University of Munich



Background Index Calculation



- Simulated background PDFs, relative scaling based on assay results
- Flat between 1950 keV and 2350 keV
- Remove ± 5 keV around $Q_{\beta\beta}$ and 3 prominent γ lines
- The remaining 360 keV window is used for background index calculation



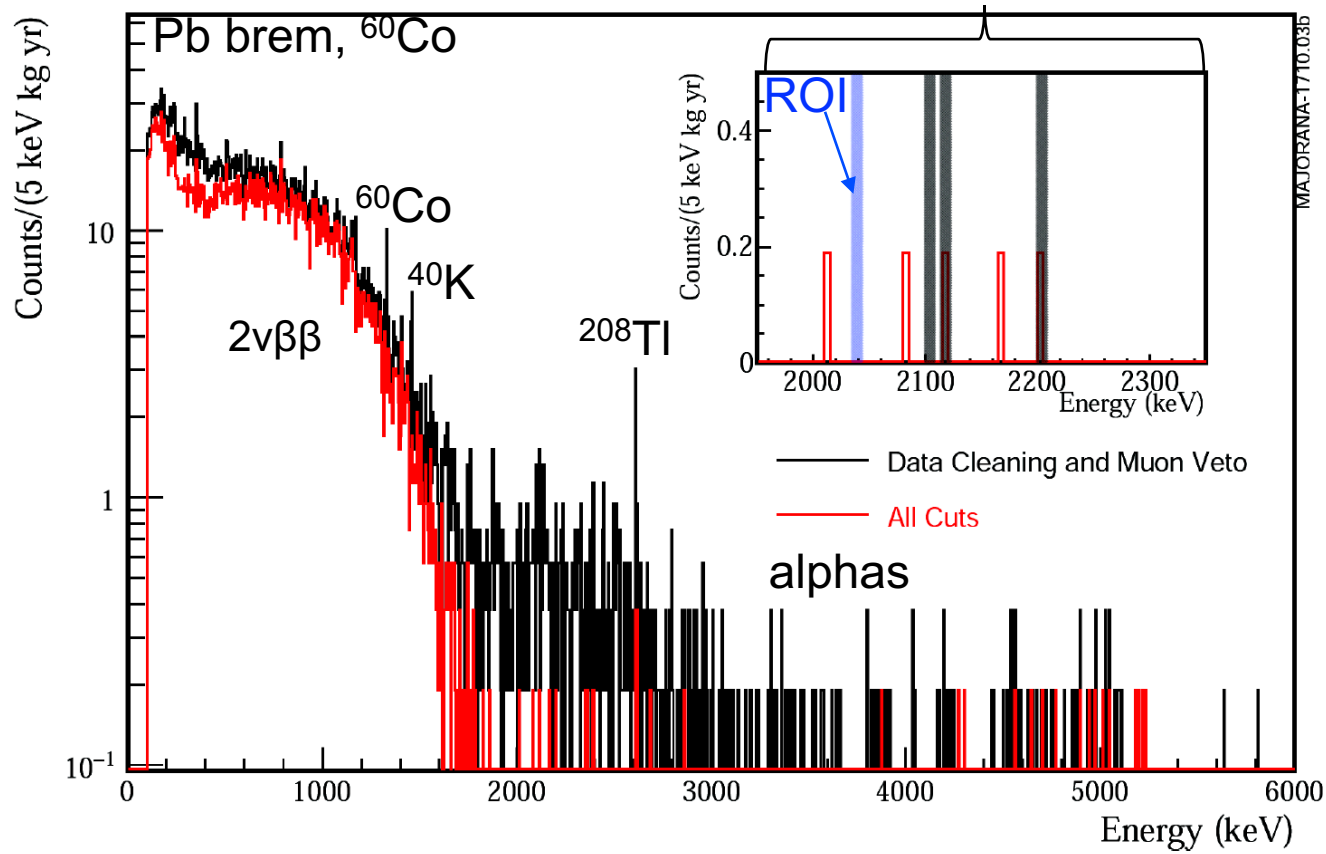
Background Spectrum



- Active exposure: 5.24 kg-y (^{enr}Ge) w/. configurations allowing lowest background
- Background after cuts: 3 counts in 360 keV window
- Background rate: $4.0^{+3.1}_{-2.5}$ c/FWHM/t/y; $1.6^{+1.2}_{-1.0} \times 10^{-3}$ c/keV/kg/y

Phys. Rev. Lett. **120.132502** (2018)

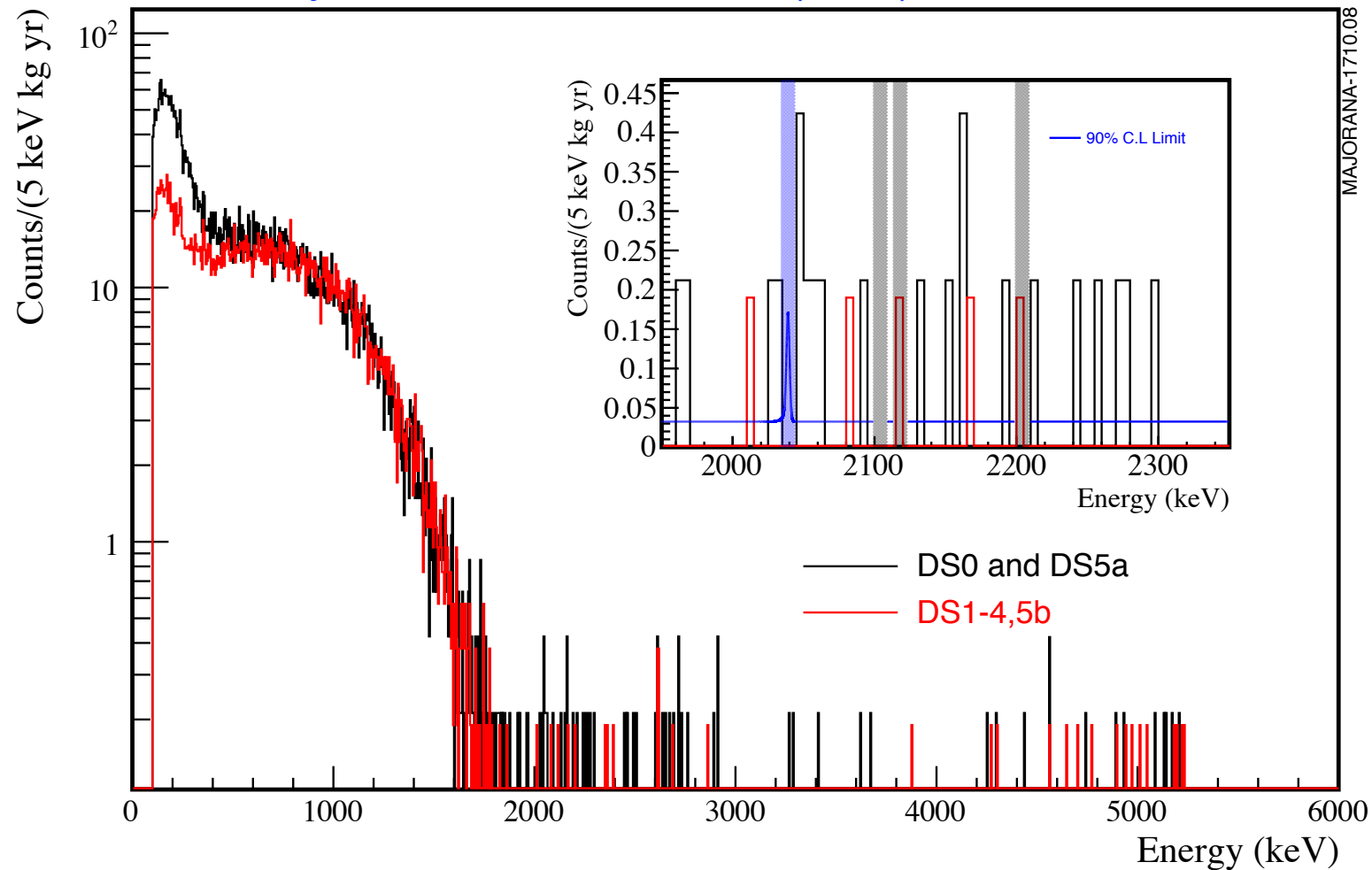
360 keV window for BKG calculation



Background Spectrum



Phys. Rev. Lett. **120**.132502 (2018)

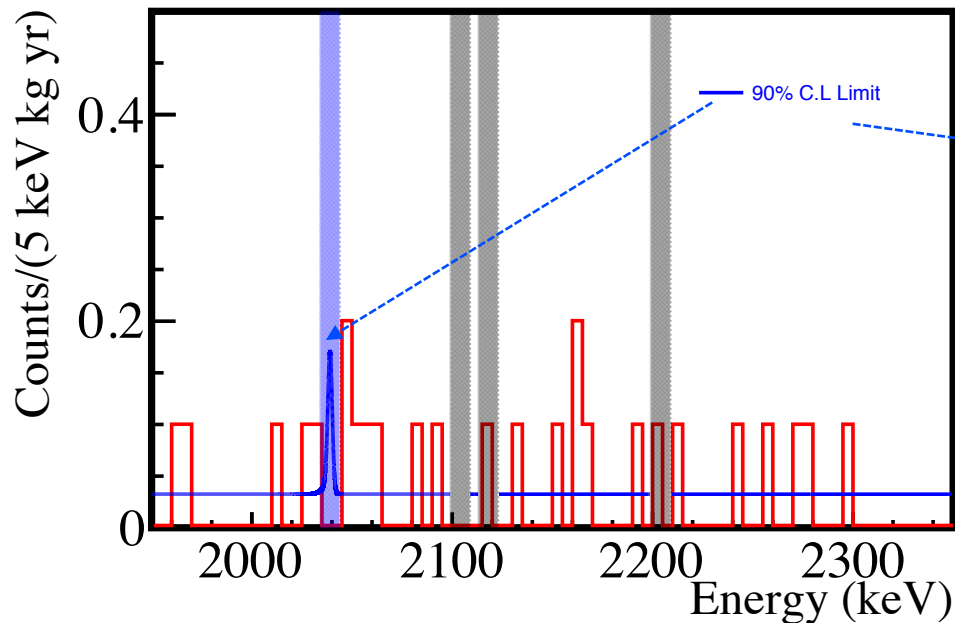


- DS0 and DS5a
 - not lowest background configurations due to incomplete shield and high noise during shield construction.
 - still can be included in decay half limit calculation.
- Total exposure is 9.95 kg-yr of enriched Ge.

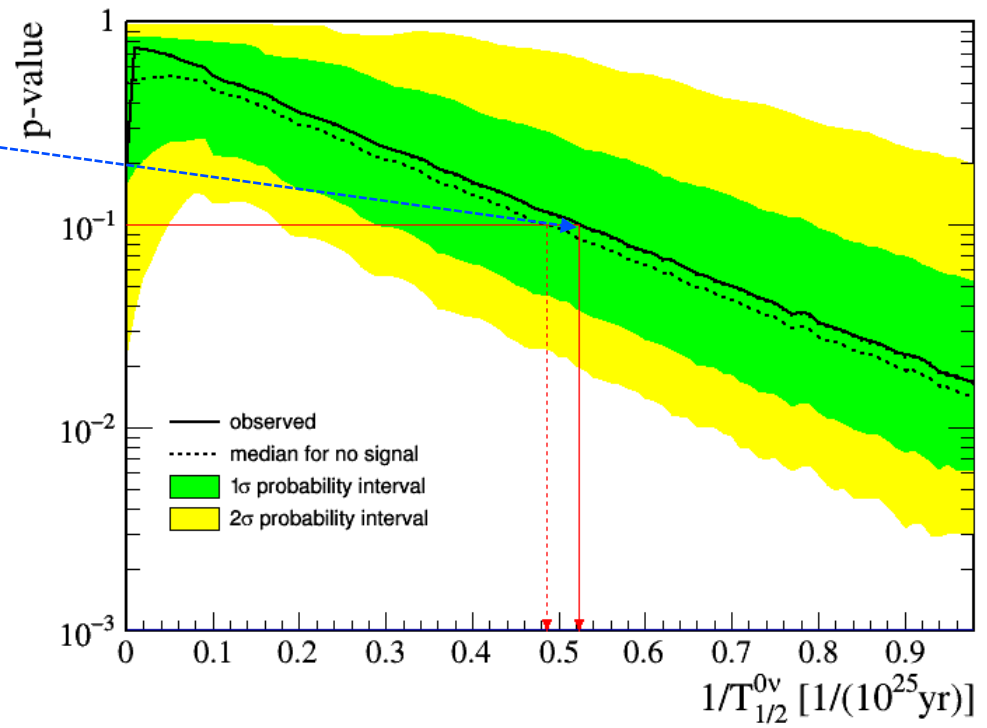
Half Life Limits



total exposure: 9.95 kg-yr



unbinned, extended profile likelihood



Phys. Rev. Lett. **120**.132502 (2018)

Also reports limits from Feldman-Cousins, CLs, and Bayesian methods: $1.5 - 1.6 \times 10^{25}$ yr.

Improved limits with ~ 26 kg-yr exposure will be released soon.

Projected sensitivity is $\sim 5 \times 10^{25}$ yr.

$$T_{1/2}^{0\nu} > 1.9 \times 10^{25} \text{ yr (90\% CL)}$$

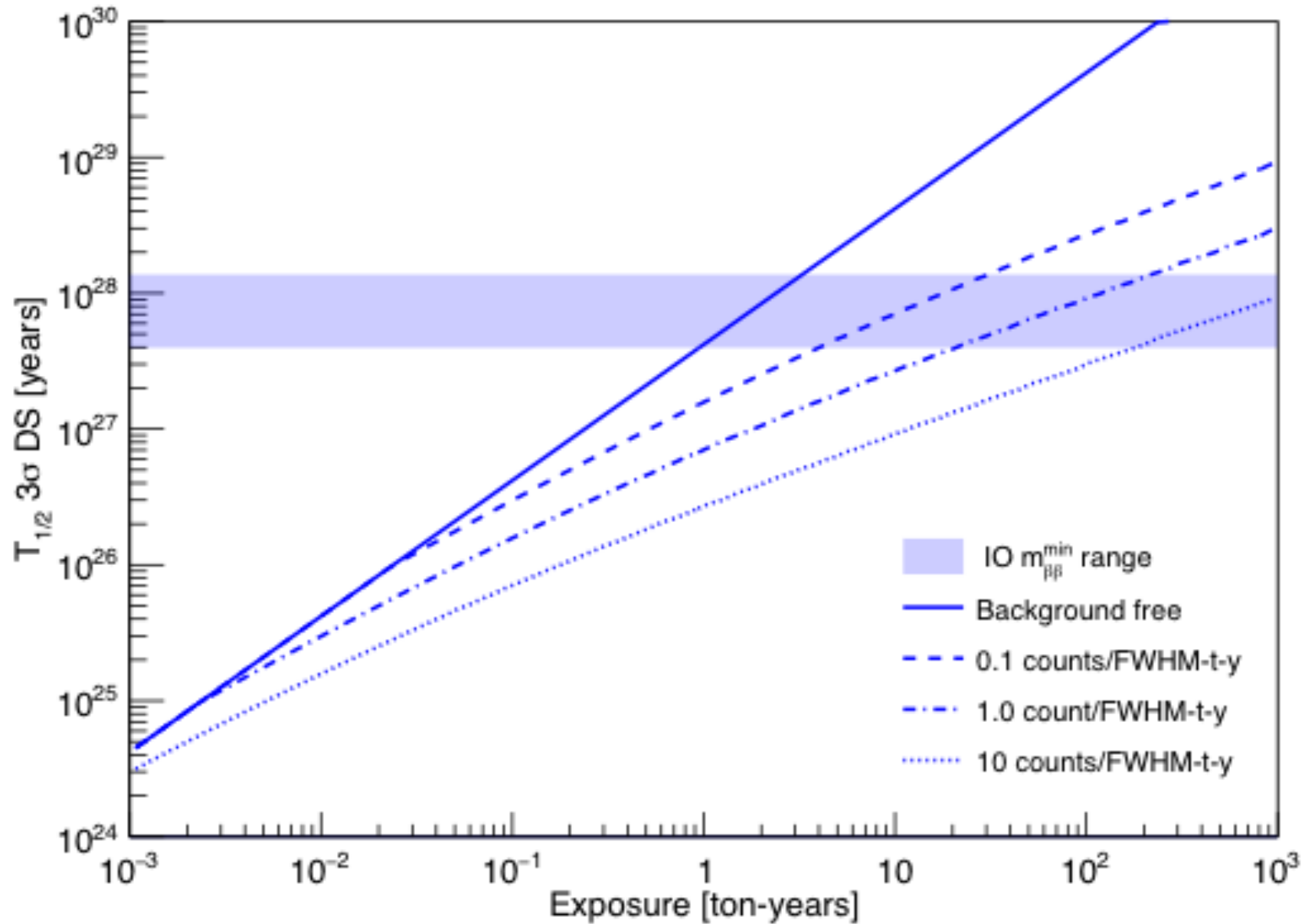
Median sensitivity (90% CL)

$$T_{1/2}^{0\nu} > 2.1 \times 10^{25} \text{ yr}$$

Discovery, Background and Exposure



^{76}Ge (88% enr.)



Advantages of Ge

- Best energy resolution: FWHM 0.12% at Q-value 2039 keV
- Demonstrated ability to enrich from 7.44% to $\geq 87\%$
- Powerful background rejection: multiplicity, timing, pulse-shape discrimination
- Intrinsic high-purity Ge detectors = source



MAJORANA DEMONSTRATOR
Compact configuration:
Vacuum cryostats in a
passive graded shield
with ultra-clean materials

GERDA
Direct immersion in
active LAr shield

- Both experiments are presently operating “background free” and benefiting from excellent energy resolution.
- Excellent limits with modest exposure.
- **Combine the best of Majorana & GERDA !**

Large Enriched Germanium Experiment for Neutrinoless $\beta\beta$ Decay



47 institutions,
237 Scientists



International
collaboration

arXiv:1709.01980. Also see the next talk by Jordan Myslik on LEGEND 17:10 - 17:30

More Physics with the DEMONSTRATOR

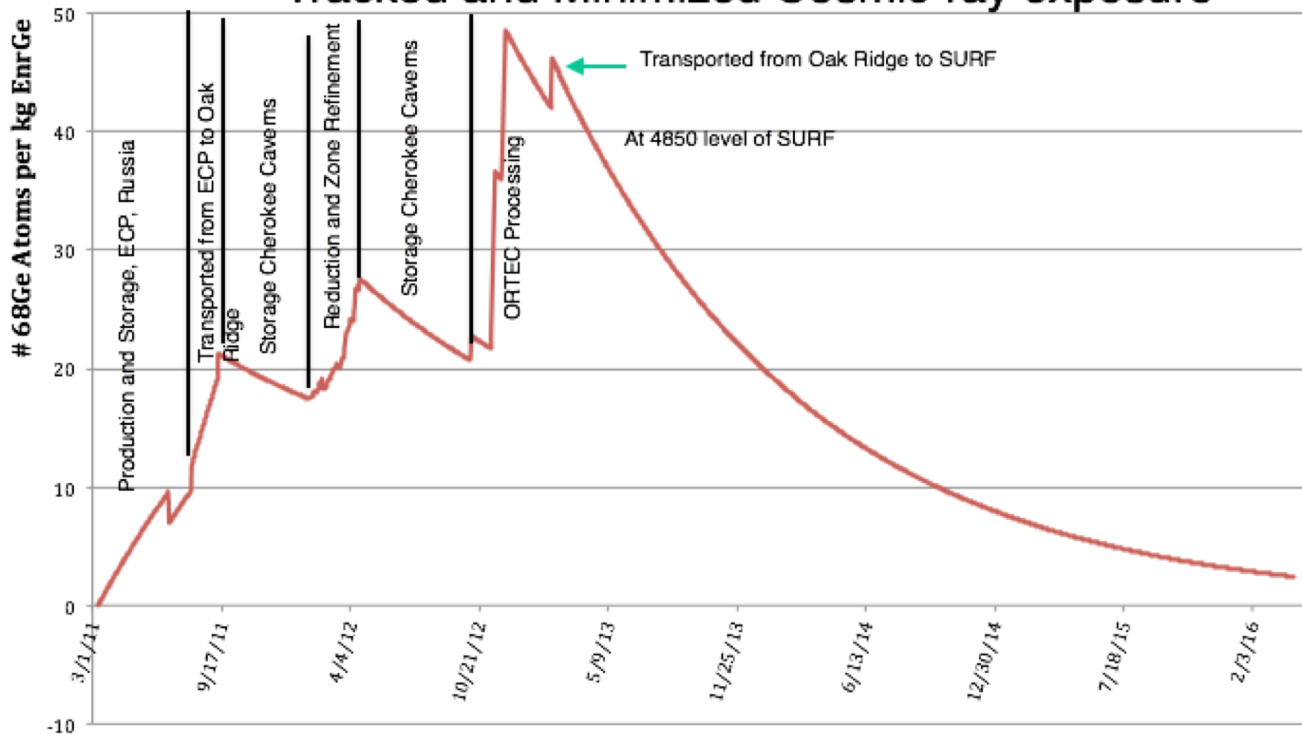


Iron shielding during transportation of the enriched ^{76}Ge from Russia



Underground cave storage during detector manufacture at vendor

Tracked and Minimized Cosmic-ray exposure



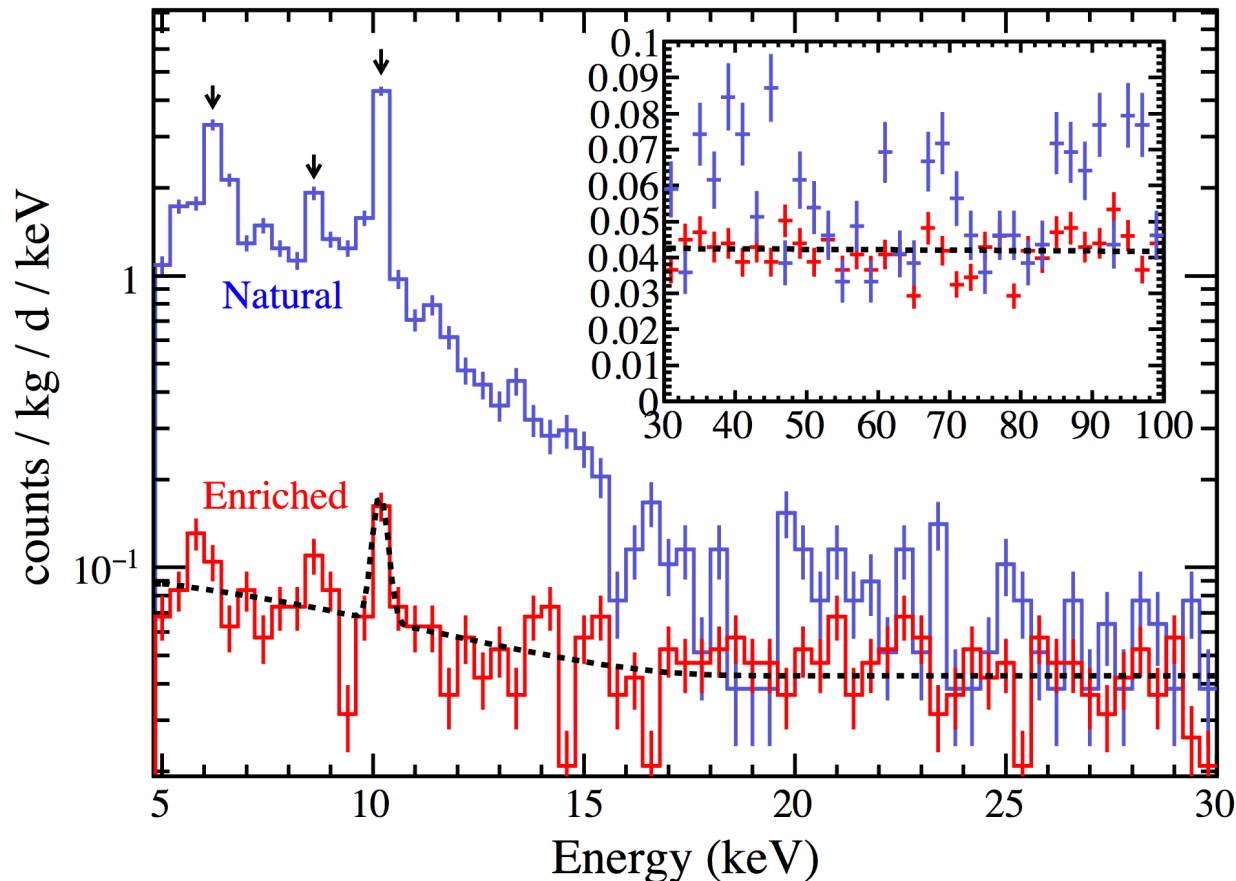
Limited exposure of enriched material to cosmic rays to ~31 days of total sea-level equivalent

Sophisticated parts-tracking database
NIMA 779(2015)52–62

Low-Energy Physics Searches



- Limited exposure of enriched material to cosmic rays
For the DEMONSTRATOR, the enriched detector ^{68}Ge rate is low enough that an X-ray delayed coincidence cut is not necessary
- Tritium is obvious and dominates in natural detectors below 20 keV
- Hardware thresholds below 1 keV, analysis below 5 keV is ongoing
- Commissioning data below. Factor of several reduction in low-energy background in later datasets.
- Updated results and new searches are coming soon.



Low-Energy Searches for Physics Beyond SM

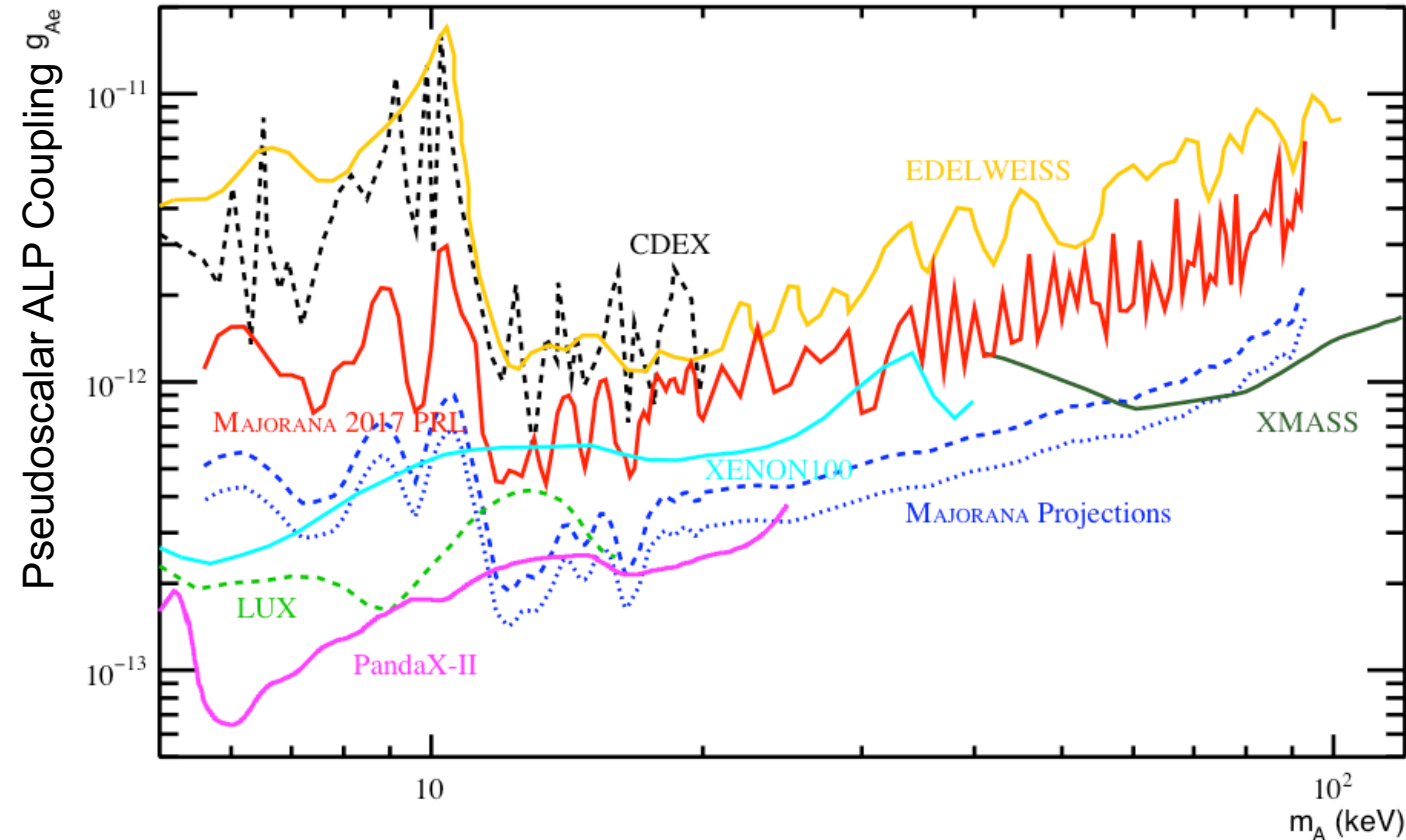
- Pseudoscalar dark matter
- Vector dark matter
- 14.4 keV solar axion
- $e^- \rightarrow 3\nu$
- Pauli Exclusion Principle violation

Phys. Rev. Lett. 118, 161801

Low-Energy Physics Searches



- Limited exposure of enriched material to cosmic rays
For the DEMONSTRATOR, the enriched detector ^{68}Ge rate is low enough that an X-ray delayed coincidence cut is not necessary
- Tritium is obvious and dominates in natural detectors below 20 keV
- Hardware thresholds below 1 keV, analysis below 5 keV is ongoing
- Commissioning data below. Factor of several reduction in low-energy background in later datasets
- Updated results and new searches are coming soon.



Low-Energy Searches for Physics Beyond SM

- Pseudoscalar dark matter
- Vector dark matter
- 14.4 keV solar axion
- $e^- \rightarrow 3\nu$
- Pauli Exclusion Principle violation

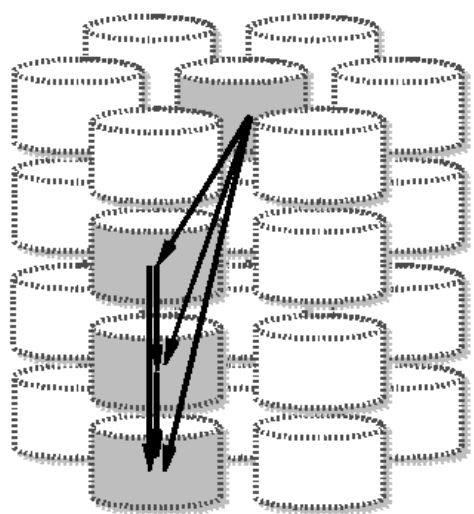
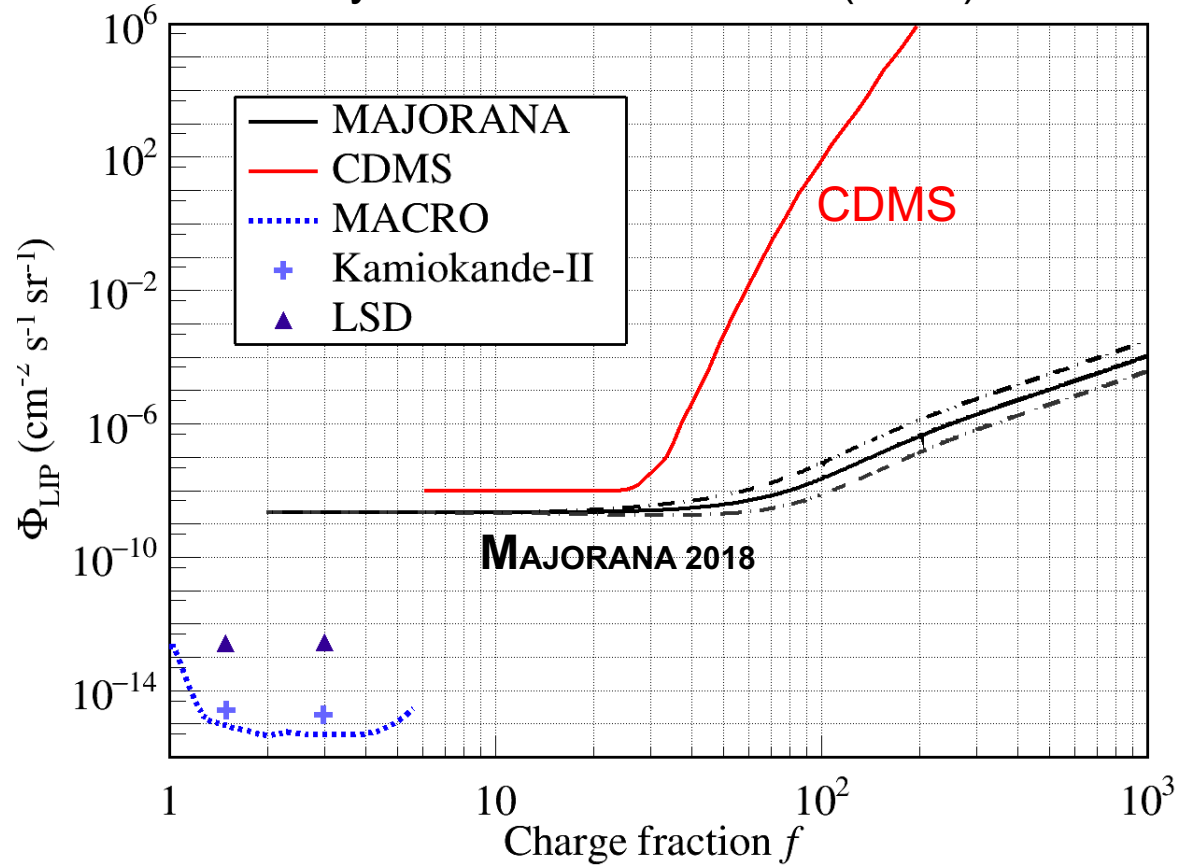
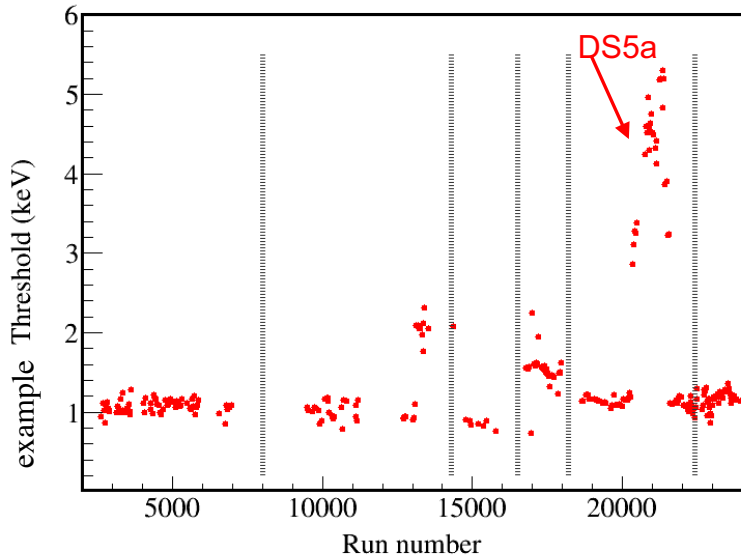
Phys. Rev. Lett. 118, 161801

Low-Energy Physics Searches



First Limit on the Direct Detection of Lightly Ionizing Particles for Electric Charge as Low as $e/1000$

PhysRevLett.120.211804 (2018)



- Thanks to low backgrounds and low energy thresholds
- First results for a non-accelerator based experiment on the natural flux of lightly ionizing particles with charges less than $e/200$
- Improvement of the existing limits between $e/6$ and $e/200$.

Summary



- Using ^{76}Ge enriched PPC detectors, MAJORANA have
 - ultra-low activity materials and low-mass designs
 - the best energy resolution (2.5 keV FWHM at 2039 keV) of any $0\nu\beta\beta$ -decay experiment
 - excellent pulse shape discrimination for reduction of backgrounds.
- The observed background index of the DEMONSTRATOR in the lowest background configuration is $4.0^{+3.1}_{-2.5}$ cts/(FWHM t yr). [Phys. Rev. Lett.120.132502 \(2018\)](#)
- Established $0\nu\beta\beta$ -decay half life $> 1.9 \times 10^{29}$ yr at 90% CL based on ~ 10 kg-yr of exposure.
- Unblind in Summer 2018 with a total exposure of ~ 26 kg-yr and a projected sensitivity of $\sim 5 \times 10^{25}$ yr.
- Combining the strengths of GERDA and the MAJORANA DEMONSTRATOR, the LEGEND collaboration is moving forward towards a ton-scale ^{76}Ge based experiment
- Low energy thresholds, excellent energy resolution and reduced cosmology activations in enriched detectors allowing the DEMONSTRATOR to perform sensitive tests at low-energy for rich physics beyond the standard model.



Black Hills State University, Spearfish, SD
Kara Keeter

Duke University, Durham, North Carolina, and TUNL
Matthew Busch

Joint Institute for Nuclear Research, Dubna, Russia
Viktor Brudanin, M. Shirchenko, Sergey Vasilyev, E. Yakushev, I. Zhitnikov

Lawrence Berkeley National Laboratory, Berkeley, California and
the University of California - Berkeley
Nicolas Abgrall, Yuen-Dat Chan, Lukas Hehn, Jordan Myslik,
Alan Poon, Kai Vetter

Los Alamos National Laboratory, Los Alamos, New Mexico
Pinghan Chu, Steven Elliott, Ralph Massarczyk, Keith Rielage,
Brandon White, Brian Zhu

Massachusetts Institute of Technology, Cambridge, MA
Julieta Gruszko

National Research Center 'Kurchatov Institute' Institute of Theoretical and
Experimental Physics, Moscow, Russia
Alexander Barabash, Sergey Konovalov, Vladimir Yumatov

North Carolina State University, and TUNL
Matthew P. Green

Oak Ridge National Laboratory
Fred Bertrand, Charlie Havener, David Radford, Benjamin Shanks,
Robert Varner, Chang-Hong Yu

Osaka University, Osaka, Japan
Hiroyasu Ejiri

Pacific Northwest National Laboratory, Richland, Washington
Isaac Arnquist, Eric Hoppe, Richard T. Kouzes

Princeton University, Princeton, New Jersey
Graham K. Giovanetti

Queen's University, Kingston, Canada
Ryan Martin, Alex Piliounis, Vasundhara

South Dakota School of Mines and Technology, Rapid City, South Dakota
Brady Bos, Colter Dunagan, Cabot-Ann Christofferson, Jared Thompson

Tennessee Tech University, Cookeville, Tennessee
Mary Kidd

Technische Universität München, and Max Planck Institute, Munich, Germany
Tobias Bode, Susanne Mertens

University of North Carolina, Chapel Hill, North Carolina, and TUNL
Thomas Caldwell, Thomas Gilliss, Chris Haufe, Reyco Henning, Mark Howe,
Samuel J. Meijer, Gulden Othman, Jamin Rager, Anna Reine, John F. Wilkerson

University of South Carolina, Columbia, South Carolina
Frank Avignone, David Edwins, Vincente Guisepppe, David Tedeschi, Clint Wiseman

University of South Dakota, Vermillion, South Dakota
Clay J. Barton, Wenqin Xu

University of Tennessee, Knoxville, Tennessee
Yuri Efremenko, Andrew Lopez

University of Washington, Seattle, Washington
Sebastian Alvis, Micah Buuck, Clara Cuesta, Jason Detwiler,
Ian Guinn, Walter Pettus, Nick Ruof



U.S. DEPARTMENT OF ENERGY

Office of Science

The MAJORANA Collaboration



MAX-PLANCK-GESELLSCHAFT



Massachusetts Institute of Technology



大阪大学
OSAKA UNIVERSITY



Proudly Operated by **Battelle** Since 1965



SOUTH DAKOTA



Technische Universität München



THE UNIVERSITY OF NORTH CAROLINA
at CHAPEL HILL



UNIVERSITY OF SOUTH CAROLINA



UNIVERSITY OF SOUTH DAKOTA



THE UNIVERSITY OF TENNESSEE
KNOXVILLE



Backup



Module and Shield Details



Calibration System

Thermosyphon

Upper Veto

Poly Shield

Cryostat

Keyed Pb Stacks

Air Bearing Transport

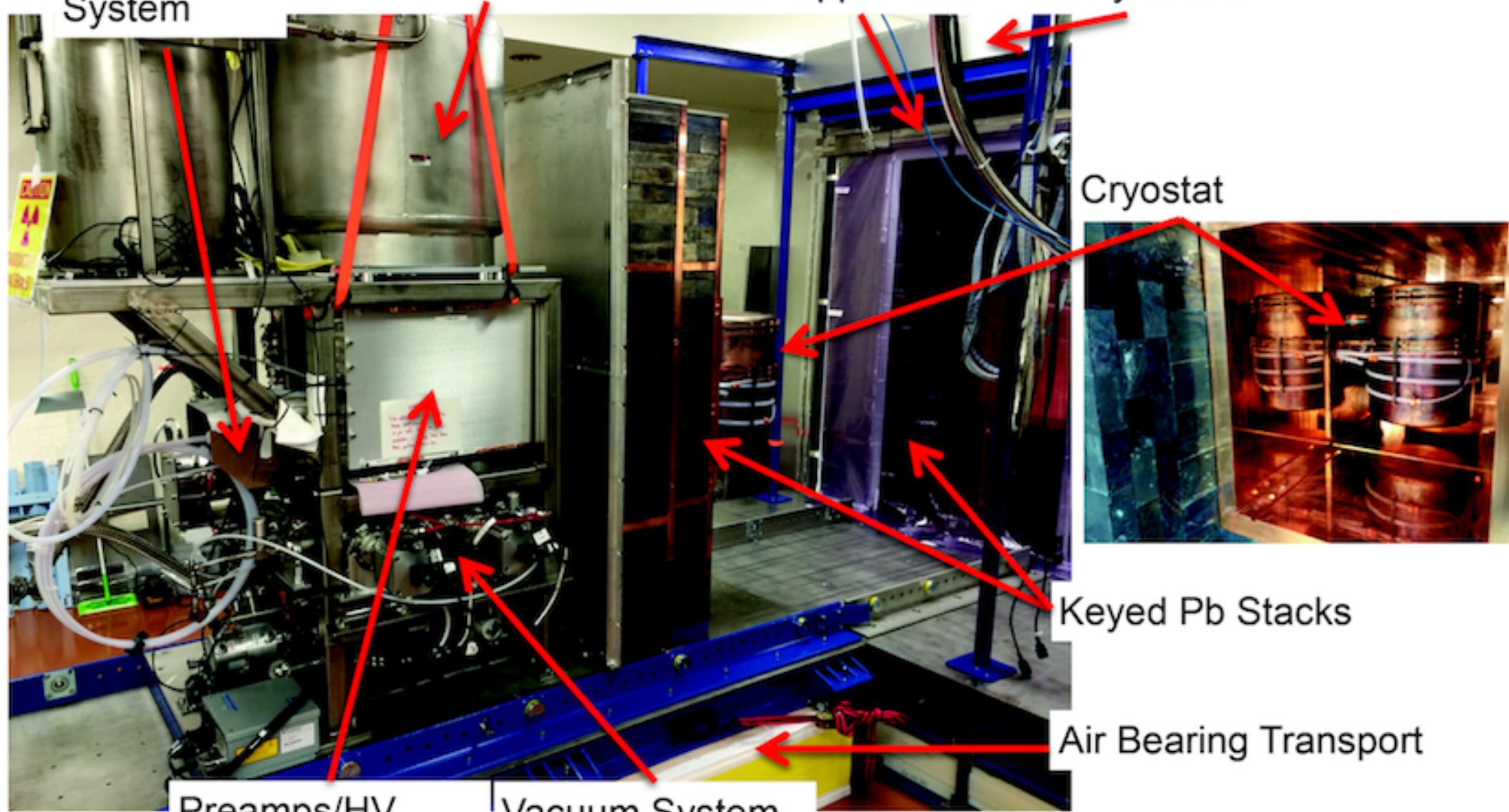
Preamps/HV Distribution

Vacuum System

Wenqin Xu

5/29/18

29



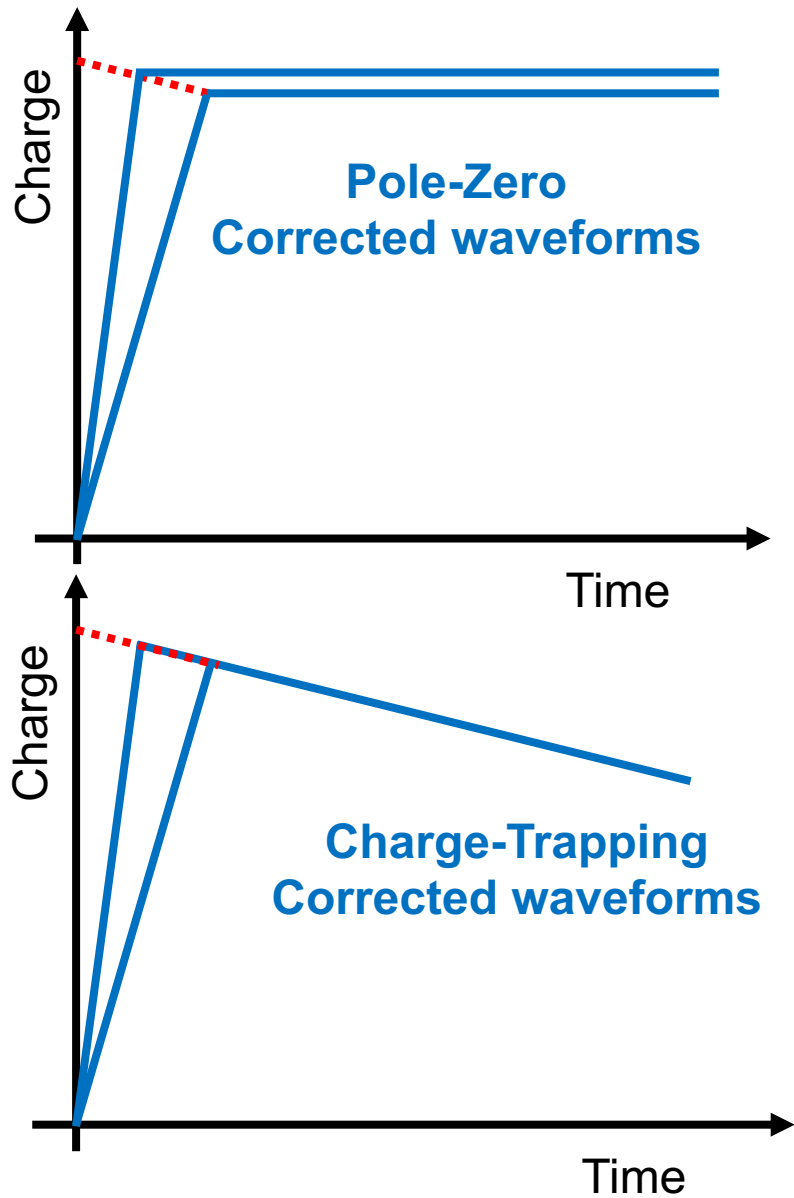
MAJORANA Electroformed Cu



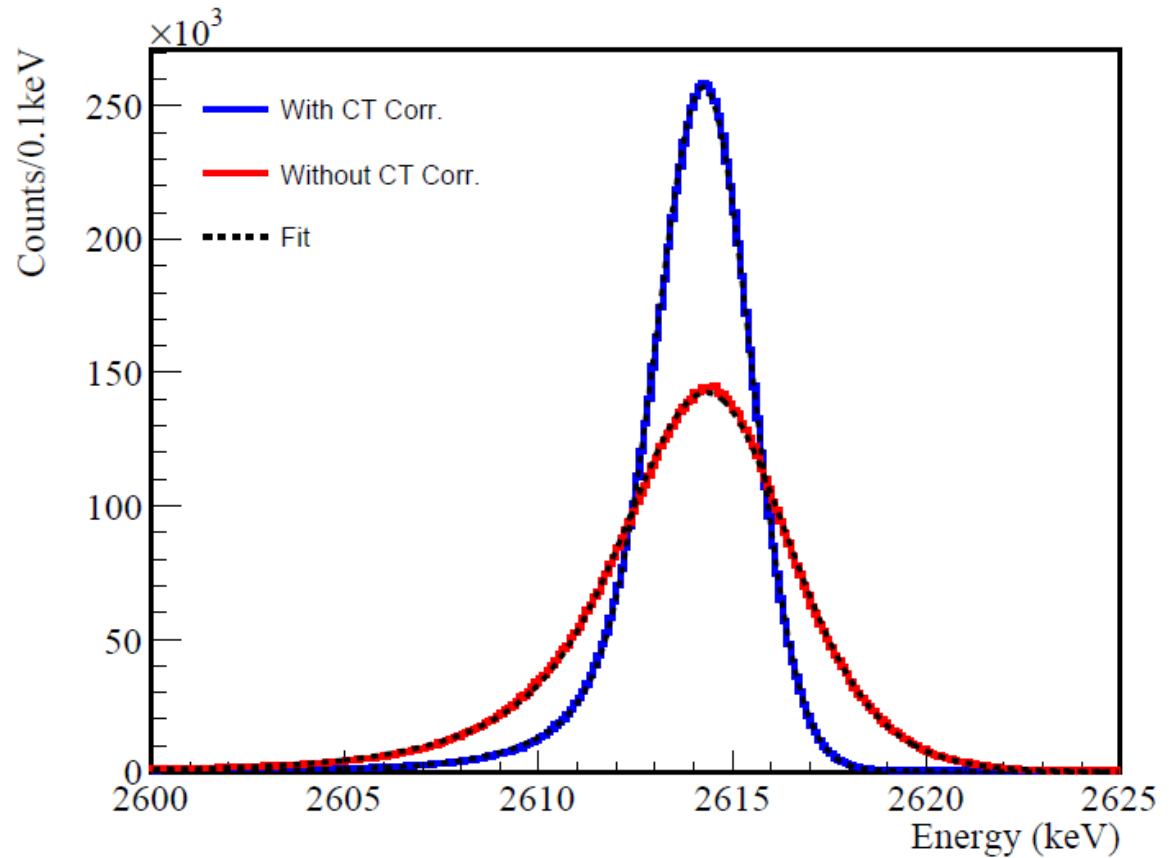
- Majorana operated 10 baths at the Temporary Clean Room (TCR) facility at the 4850' level and 6 baths at a shallow UG site at PNNL. All copper was machined at the Davis campus.
- The electroforming of copper for the Demonstrator successfully completed in May 2015.
- 2474 kg of electroformed copper on the mandrels,
- 2104 kg after initial machining,
- 1196 kg installed in the DEMONSTRATOR.
- Underground machining completed April 2016.

- Th decay chain (ave) $\leq 0.1 \mu\text{Bq/kg}$
- U decay chain (ave) $\leq 0.1 \mu\text{Bq/kg}$

Charge Trapping Correction

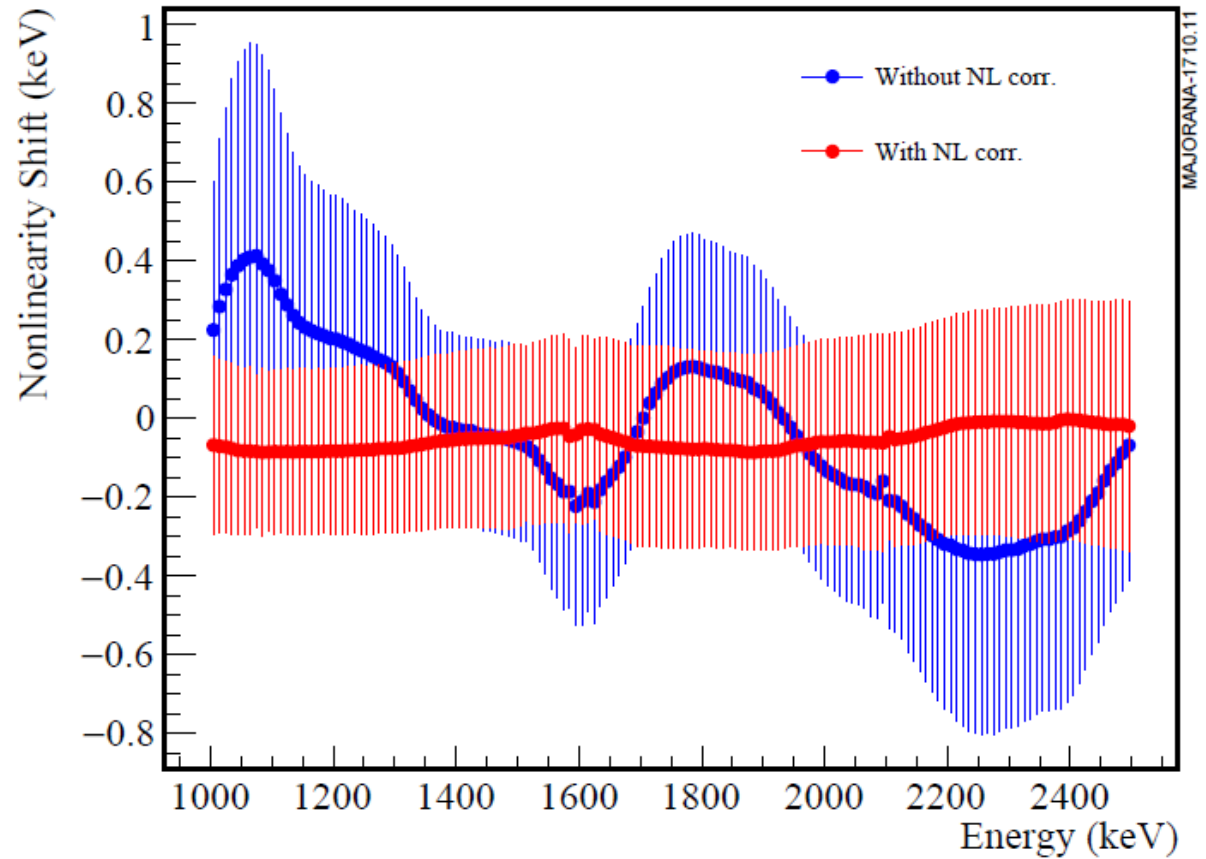
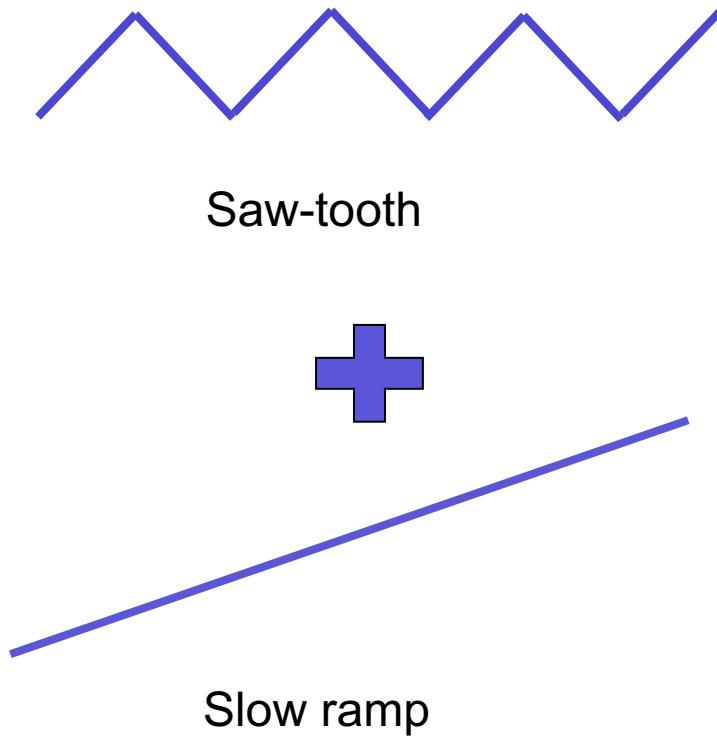


Total charge collected depends on rise time of current pulse



Adjust Pole-Zero correction so that decay time of waveforms matches charge trapping time constant. Provides ~25% improvement in energy resolution

Energy Nonlinearity Correction



Use saw tooth + slow ramp input voltage to measure ΔV of each ADC bin.
Use this to correct ADC nonlinearity