



New Formulation of the γW -box correction to neutron and nuclear β -decay

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Outline

- Beta decay in presence of RC
- Dispersive representation of the γW -box
- Physics input to the dispersion integral
- Nuclear effects
- New formulation of RC for V_{ud} extraction
- Can nuclear effects turn the inner correction inside-out?

Neutron β -decay in presence of RC

Beyond RC that enter the Fermi constant:

Sirlin '67, Marciano & Sirlin '86 ...

Coulomb distortion - Fermi fn.

$$F(\beta) \approx 1 + \alpha\pi/\beta$$

"Inner" correction -

depends on hadron structure;
independent of kinematics

$$\frac{d\Gamma}{dE_e d\Omega_e d\Omega_\nu} = \frac{G_\mu^2 V_{ud}^2}{(2\pi)^5} (1 + 3\lambda^2) |\vec{p}_e| E_e (E_m - E_e)^2 F(\beta) \left(1 + \frac{\alpha}{\pi} \text{Re } c\right) \left(1 + \frac{\alpha}{2\pi} \delta^{(1)}\right) \times$$

$$\left[1 + \left(1 + \frac{\alpha}{2\pi} \delta^{(2)}\right) a \vec{\beta} \cdot \hat{p}_\nu + \hat{s} \cdot \left(\left(1 + \frac{\alpha}{2\pi} \delta^{(2)}\right) A \vec{\beta} + B \hat{p}_\nu \right) \right]$$

"Outer" corrections - IR-sensitive;
depend on kinematics;
independent of hadronic structure
Exactly calculable

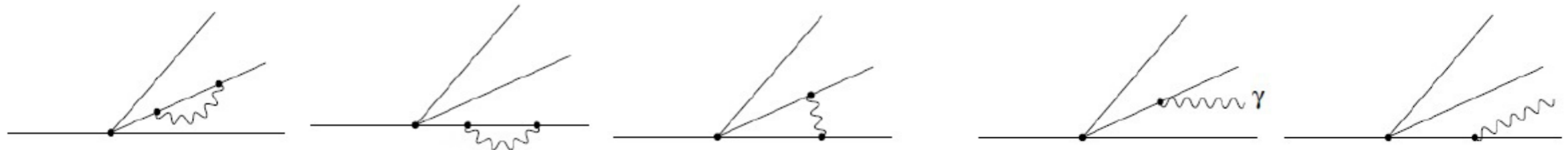
Wilkinson '82, Severijns et al. '17

Separation due to scale hierarchy: Q-values from <1 keV (n) to few MeV (nuclei);
Hadronic scales: at least 140 MeV - on top of $\alpha/2\pi \sim 10^{-3} \rightarrow 10^{-5}$ effect <<

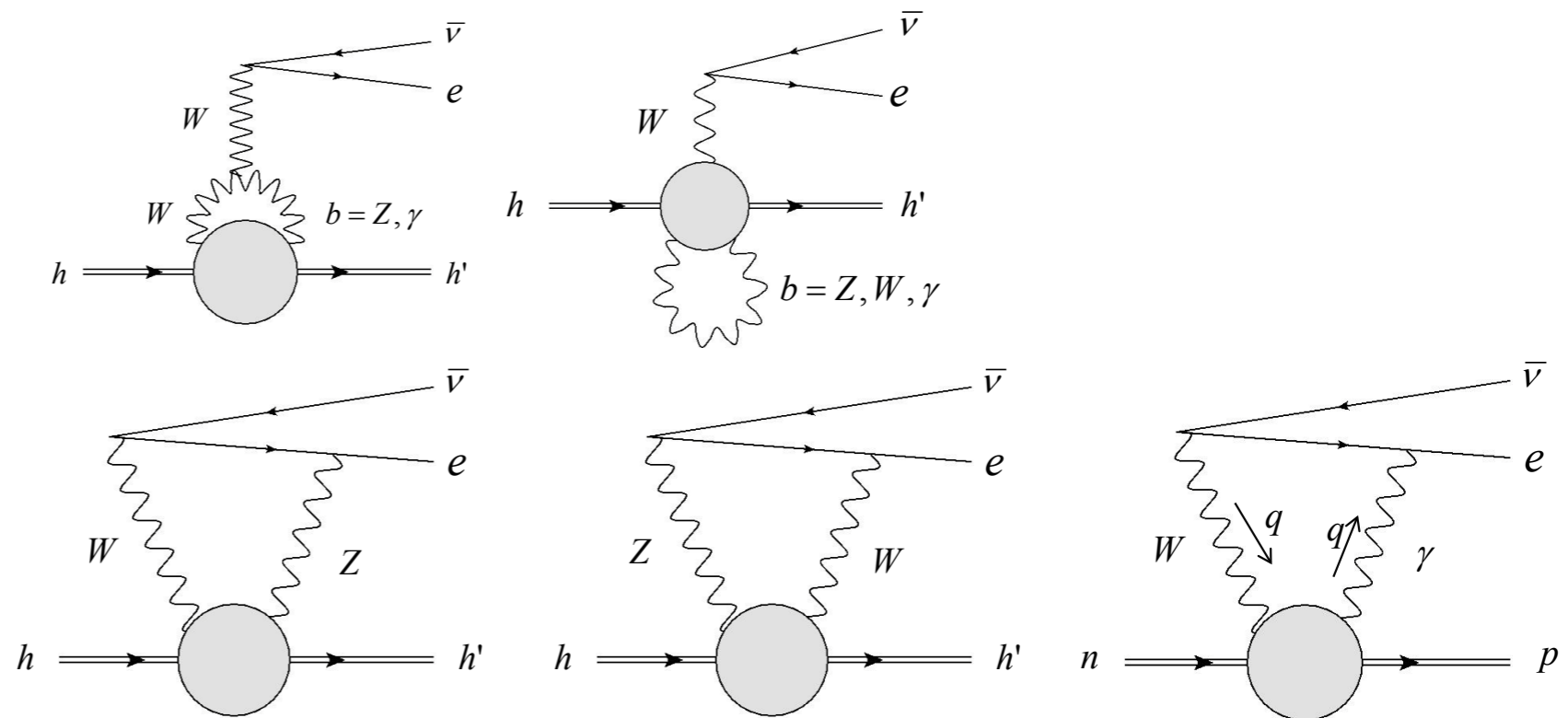
Radiative corrections - In & Out

1-loop RC (specific for a semileptonic process)

Outer: retain only IR divergent pieces



Inner: everything else



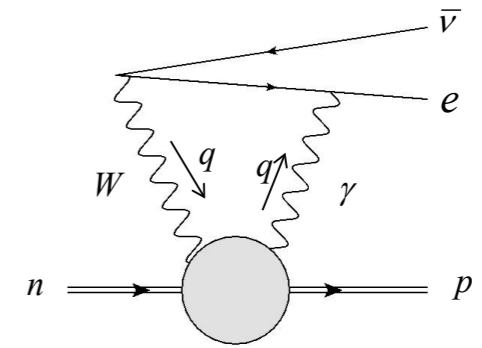
W,Z-exchange: UV-sensitive, pQCD; model-independent

When γ involved - possible sensitivity to long range physics

Model-dependent!

γW -box

Consider the box at zero energy
and zero momentum transfer



$$T_{\gamma W} = \sqrt{2}e^2 G_F V_{ud} \int \frac{d^4 q}{(2\pi)^4} \frac{\bar{u}_e \gamma^\mu (\not{k} - \not{q} + m_e) \gamma^\nu (1 - \gamma_5) v_\nu}{q^2 [(k - q)^2 - m_e^2]} \frac{M_W^2}{q^2 - M_W^2} T_{\mu\nu}^{\gamma W}$$

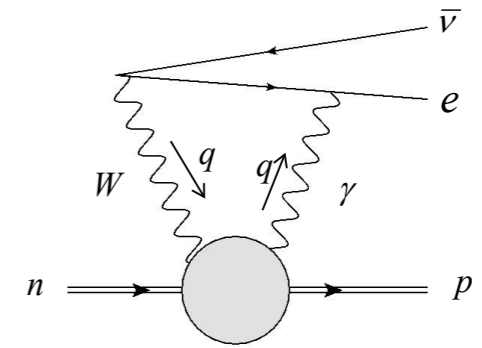
Hadronic tensor: two-current correlator $T_{\gamma W}^{\mu\nu} = \int dx e^{iqx} \langle p | T [J_{em}^\mu(x) J_W^\nu(0)] | n \rangle$

General gauge-invariant decomposition (spin-independent)

$$T_{\gamma W}^{\mu\nu} = \left(-g^{\mu\nu} + \frac{q^\mu q^\nu}{q^2} \right) T_1 + \frac{1}{(p \cdot q)} \left(p - \frac{(p \cdot q)}{q^2} q \right)^\mu \left(p - \frac{(p \cdot q)}{q^2} q \right)^\nu T_2 + \frac{i\epsilon^{\mu\nu\alpha\beta} p_\alpha q_\beta}{2(p \cdot q)} T_3$$

γW -box

Consider the box at zero energy and zero momentum transfer



$$T_{\gamma W} = \sqrt{2}e^2 G_F V_{ud} \int \frac{d^4 q}{(2\pi)^4} \frac{\bar{u}_e \gamma^\mu (\not{k} - \not{q} + m_e) \gamma^\nu (1 - \gamma_5) v_\nu}{q^2 [(k - q)^2 - m_e^2]} \frac{M_W^2}{q^2 - M_W^2} T_{\mu\nu}^{\gamma W}$$

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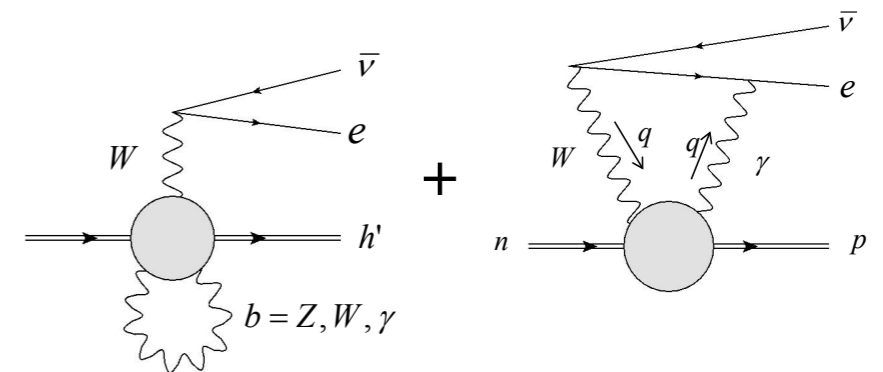
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V-V correlator is cancelled exactly

By the 3-current correlator - Sirlin '67

Reason: conserved vector-isovector current



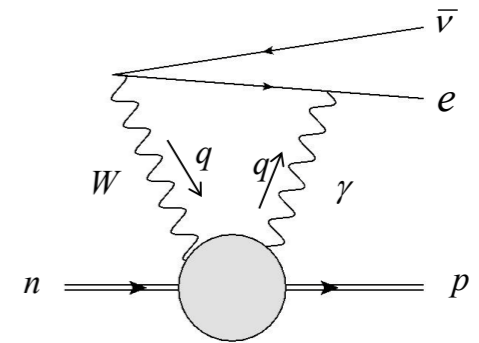
Axial current not conserved \rightarrow only A-V correlator contributes

γW -box

Consider the box at zero energy
and zero momentum transfer

Define the box contribution as

$$T_W + T_{\gamma W}^{VA} = -\sqrt{2}G_F V_{ud} (1 + \square_{\gamma W}^{VA}) \bar{u}_e \not{p} (1 - \gamma_5) v_\nu$$

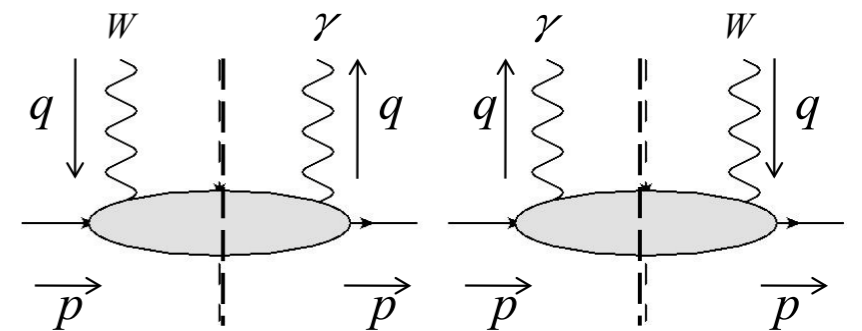


$$\nu = (pq)/M$$

Loop integral with T_3

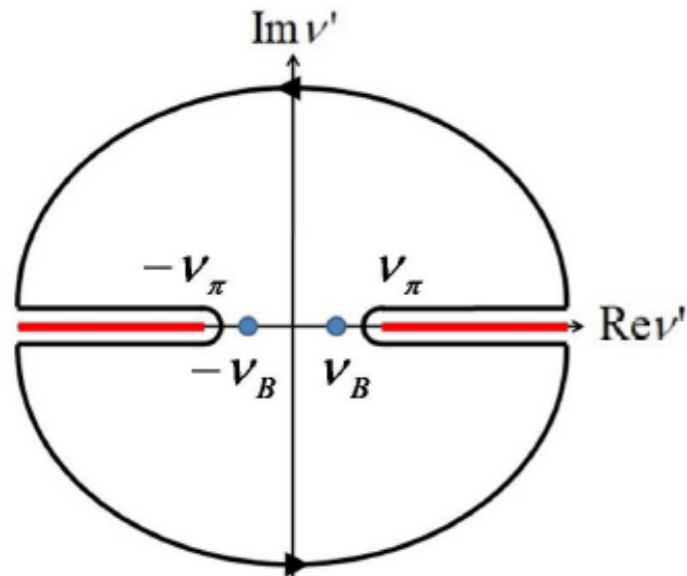
$$\square_{\gamma W}^{VA} = 4\pi\alpha \text{Re} \int \frac{d^4q}{(2\pi)^4} \frac{M_W^2}{M_W^2 + Q^2} \frac{Q^2 + \nu^2}{Q^4} \frac{T_3(\nu, Q^2)}{M\nu}$$

Forward amplitude T_3 - unknown;
Its absorptive part could be related to
production of on-shell intermediate states
a γW -analog of the SF F_3



$$\text{Im} T_3^{\gamma W}(\nu, Q^2) = 2\pi F_3^{\gamma W}(\nu, Q^2)$$

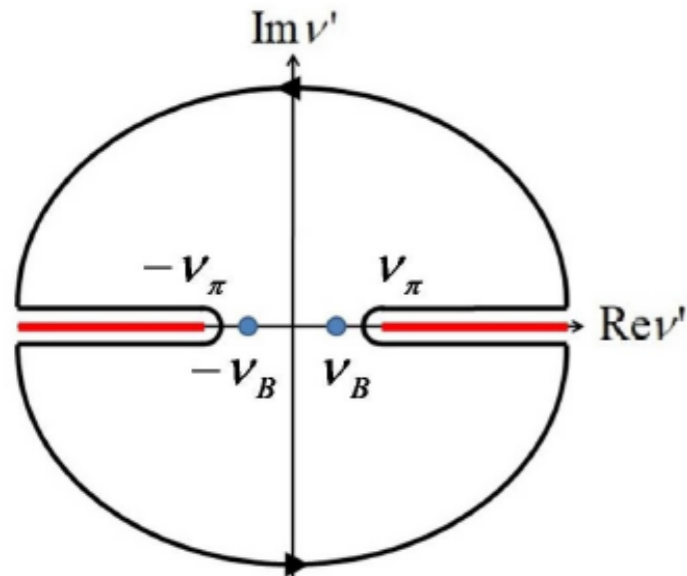
γ W-box in dispersion representation



T_3 - analytic function inside the contour C in the complex ν -plane determined by its singularities on the real axis - poles + cuts

$$T_3(\nu, Q^2) = \frac{1}{2\pi i} \oint_C \frac{T_3(z, Q^2) dz}{z - \nu} \quad \nu \in C$$

γW -box in dispersion representation



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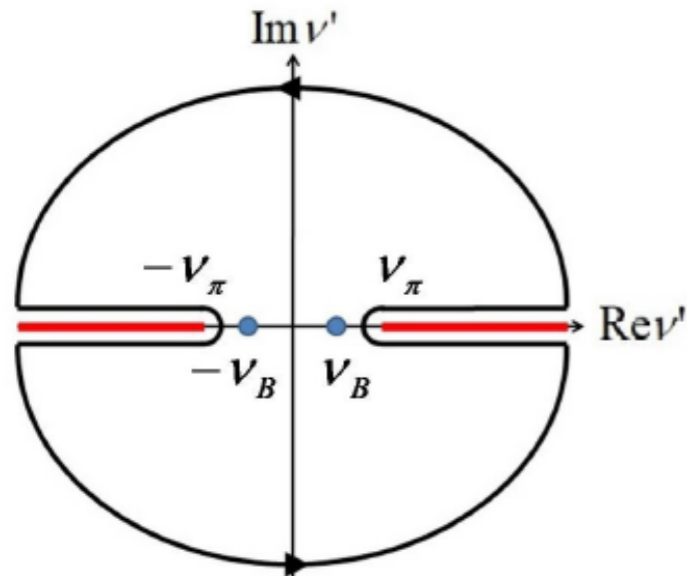
Crossing behavior: photon is isoscalar or isovector

$$T_3^{\gamma W} = T_3^{(0)} + T_3^{(3)}$$

Different isospin channels behave differently under crossing

$$T_3^{(0)}(-\nu, Q^2) = -T_3^{(0)}(\nu, Q^2), \quad T_3^{(3)}(-\nu, Q^2) = +T_3^{(3)}(\nu, Q^2)$$

γW -box in dispersion representation



T_3 - analytic function inside the contour C in the complex ν -plane determined by its singularities on the real axis - poles + cuts

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Dispersion representation of the γW -box correction at zero energy

$$\square_{\gamma W}^{VA(0)} = \frac{\alpha}{\pi M} \int_0^\infty \frac{dQ^2 M_W^2}{M_W^2 + Q^2} \int_0^\infty d\nu \frac{(\nu + 2q)}{\nu(\nu + q)^2} F_3^{(0)}(\nu, Q^2),$$

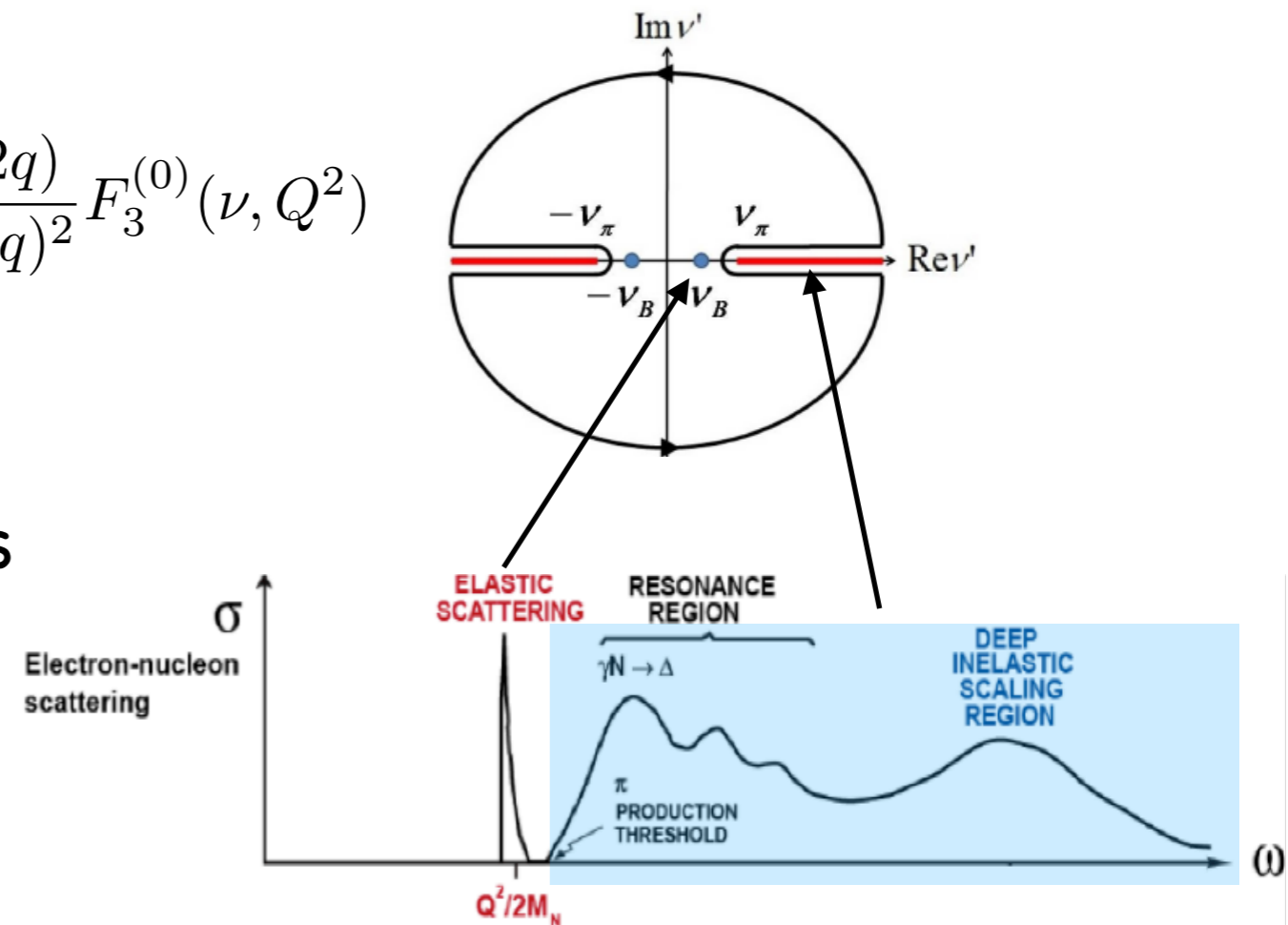
$$\square_{\gamma W}^{VA(3)} = 0,$$

$$q = \sqrt{\nu^2 + Q^2}$$

Input into dispersion integral

$$\square_{\gamma W}^{VA(0)} = \frac{\alpha}{\pi M} \int_0^\infty \frac{dQ^2 M_W^2}{M_W^2 + Q^2} \int_0^\infty d\nu \frac{(\nu + 2q)}{\nu(\nu + q)^2} F_3^{(0)}(\nu, Q^2)$$

Dispersion in energy:
scanning hadronic intermediate states



Relation to the old Marciano & Sirlin's result:

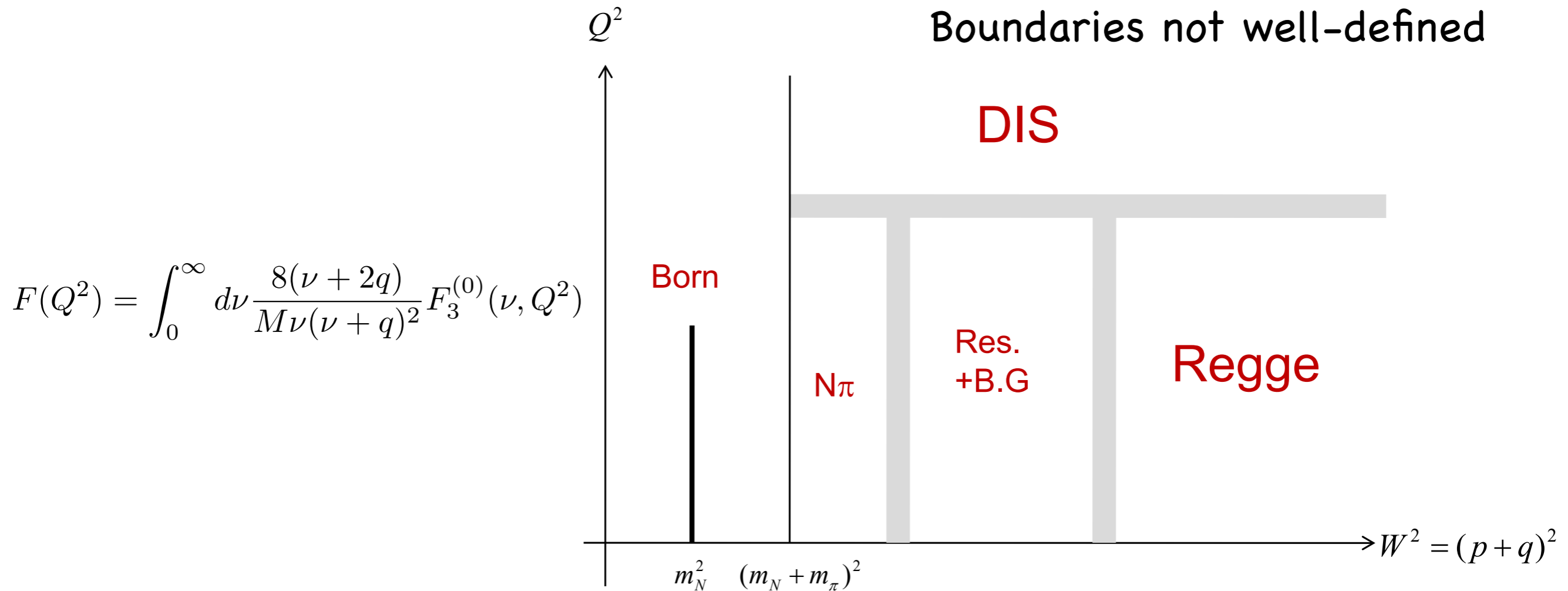
$$\square_{\gamma W}^{VA} = \frac{\alpha}{8\pi} \int_0^\infty \frac{dQ^2 M_W^2}{M_W^2 + Q^2} F(Q^2)$$

Dispersion representation of M&S loop function F:

$$F(Q^2) = \int_0^\infty d\nu \frac{8(\nu + 2q)}{M\nu(\nu + q)^2} F_3^{(0)}(\nu, Q^2)$$

Input into dispersion integral

Dispersion in Q^2 : scanning dominant physics pictures

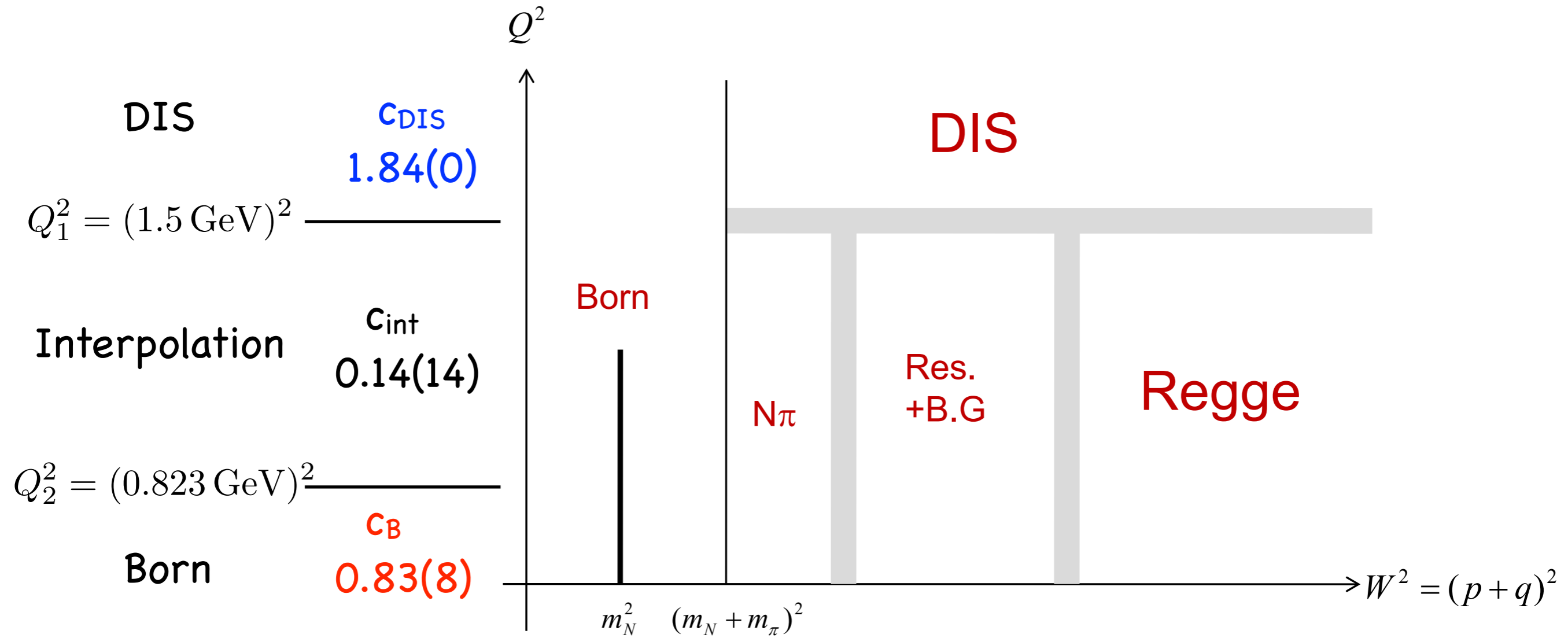


For each value of Q^2 we can relate F to particular hadronic processes
 F can be related to exp. accessible cross sections

$$\text{Im} [W^{+*} + n \rightarrow \gamma^* + p] \leftrightarrow \begin{cases} \sigma(\gamma^* + p \rightarrow X) \\ \sigma(W^{+*} + n \rightarrow X) \end{cases}$$

M&S treatment

M&S notation of the γW -box correction: $\square_{\gamma W}^{VA} = \frac{\alpha}{2\pi} [c_B + c_{int} + c_{DIS}]$



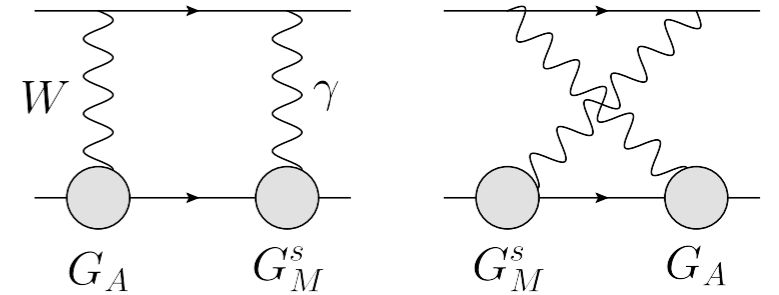
What can be improved?

- * what is the physics content of the interpolating function?
- * are the M&S constraints on F_{int} justified?

Physics input

Elastic (Born) contribution

$$\square_{\gamma W}^{VA, \text{Born}} = -\frac{\alpha}{\pi} \int_0^\infty dQ \frac{2\sqrt{4M^2 + Q^2} + Q}{(\sqrt{4M^2 + Q^2} + Q)^2} G_A(Q^2) G_M^S(Q^2)$$



M & S

$$\square_{\gamma W}^{VA, \text{Born}} \Big|_{\text{MS}} = \frac{\alpha}{2\pi} (0.829 \pm 0.083)$$

Central value:

two dipoles integrated to $(0.823 \text{ GeV})^2$

Uncertainty: vary axial dipole masses between 1 and 1.4 GeV

New evaluation

$$\square_{\gamma W}^{VA, \text{Born}} = \frac{\alpha}{2\pi} (0.908 \pm 0.049)$$

Central value: full Q^2 integral

with most recent FF parametrization

Uncertainty: from recent analyses

Magnetic FF: Ye, Arrington, Hill, Lee '18

Axial FF: Bhattacharya, Hill, Paz '11

Physics input

Inelastic contributions $\square_{\gamma W}^{Inel.} = \frac{\alpha}{\pi} \int_0^\infty \frac{dQ^2}{1 + \frac{Q^2}{M_W^2}} \int_{\nu_\pi}^\infty \frac{d\nu}{(\nu + q)^2} \frac{\nu + 2q}{M\nu} F_3^{(0),inel.}$

Split the Q^2 integral: above $(1.5 \text{ GeV})^2$ - DIS; below - hadronic stuff

Physics input

Inelastic contributions $\square_{\gamma W}^{Inel.} = \frac{\alpha}{\pi} \int_0^\infty \frac{dQ^2}{1 + \frac{Q^2}{M_W^2}} \int_{\nu_\pi}^\infty \frac{d\nu}{(\nu + q)^2} \frac{\nu + 2q}{M\nu} F_3^{(0),inel.}$

Split the Q^2 integral: above $(1.5 \text{ GeV})^2$ - DIS; below - hadronic stuff

DIS contribution $\square_{\gamma W}^{DIS} = \frac{2\alpha}{\pi} \int_{\Lambda^2}^\infty \frac{dQ^2 M_W^2}{Q^2(Q^2 + M_W^2)} \int_0^{x_\pi} dx \frac{1 + 2\sqrt{1 + 4M^2 x^2 / Q^2}}{(1 + \sqrt{1 + 4M^2 x^2 / Q^2})^2} F_3^{(0)}(x, Q^2)$

Physics input

Inelastic contributions $\square_{\gamma W}^{Inel.} = \frac{\alpha}{\pi} \int_0^\infty \frac{dQ^2}{1 + \frac{Q^2}{M_W^2}} \int_{\nu_\pi}^\infty \frac{d\nu}{(\nu + q)^2} \frac{\nu + 2q}{M\nu} F_3^{(0),inel.}$

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Parton model: $F_3^{(0)}(x) = \frac{e_u + e_d}{8} (d(x) - \bar{u}(x))$ $\int_0^1 dx d_v(x) = 2$

$M/Q \rightarrow 0$; loop function becomes $F^{DIS}(Q^2) = \frac{1}{Q^2}$

Large log: $\square_{\gamma W}^{DIS} = \frac{\alpha}{8\pi} \int_{\Lambda^2}^\infty \frac{dQ^2 M_W^2}{M_W^2 + Q^2} F^{DIS}(Q^2) = \frac{\alpha}{4\pi} \ln \frac{M_W}{\Lambda}$

Physics input

Inelastic contributions $\square_{\gamma W}^{Inel.} = \frac{\alpha}{\pi} \int_0^\infty \frac{dQ^2}{1 + \frac{Q^2}{M_W^2}} \int_{\nu_\pi}^\infty \frac{d\nu}{(\nu + q)^2} \frac{\nu + 2q}{M\nu} F_3^{(0),inel.}$

Split the Q^2 integral: above $(1.5 \text{ GeV})^2$ - DIS; below - hadronic stuff

DIS contribution $\square_{\gamma W}^{DIS} = \frac{2\alpha}{\pi} \int_{\Lambda^2}^\infty \frac{dQ^2 M_W^2}{Q^2(Q^2 + M_W^2)} \int_0^{x_\pi} dx \frac{1 + 2\sqrt{1 + 4M^2 x^2/Q^2}}{(1 + \sqrt{1 + 4M^2 x^2/Q^2})^2} F_3^{(0)}(x, Q^2)$

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pQCD corrections:
cf. GLS and Bjorken SR

$$F^{DIS} = \frac{1}{Q^2} \rightarrow \frac{1}{Q^2} \left[1 - \frac{\alpha_s^{\overline{MS}}}{\pi} - C_2 \left(\frac{\alpha_s^{\overline{MS}}}{\pi} \right)^2 - C_3 \left(\frac{\alpha_s^{\overline{MS}}}{\pi} \right)^3 \right]$$

M&S '06; Larin, Vermaseren '97

$$\square_{\gamma W}^{DIS} = \frac{\alpha}{4\pi} [4.11 - 0.34] = \frac{\alpha}{2\pi} 1.84(0)$$

Uncertainty: virtually zero

Physics input

The DIS contribution can be validated by data

Use the Gross-Llewellyn-Smith sum rule (GLS SR) in nu/anti-nu scattering

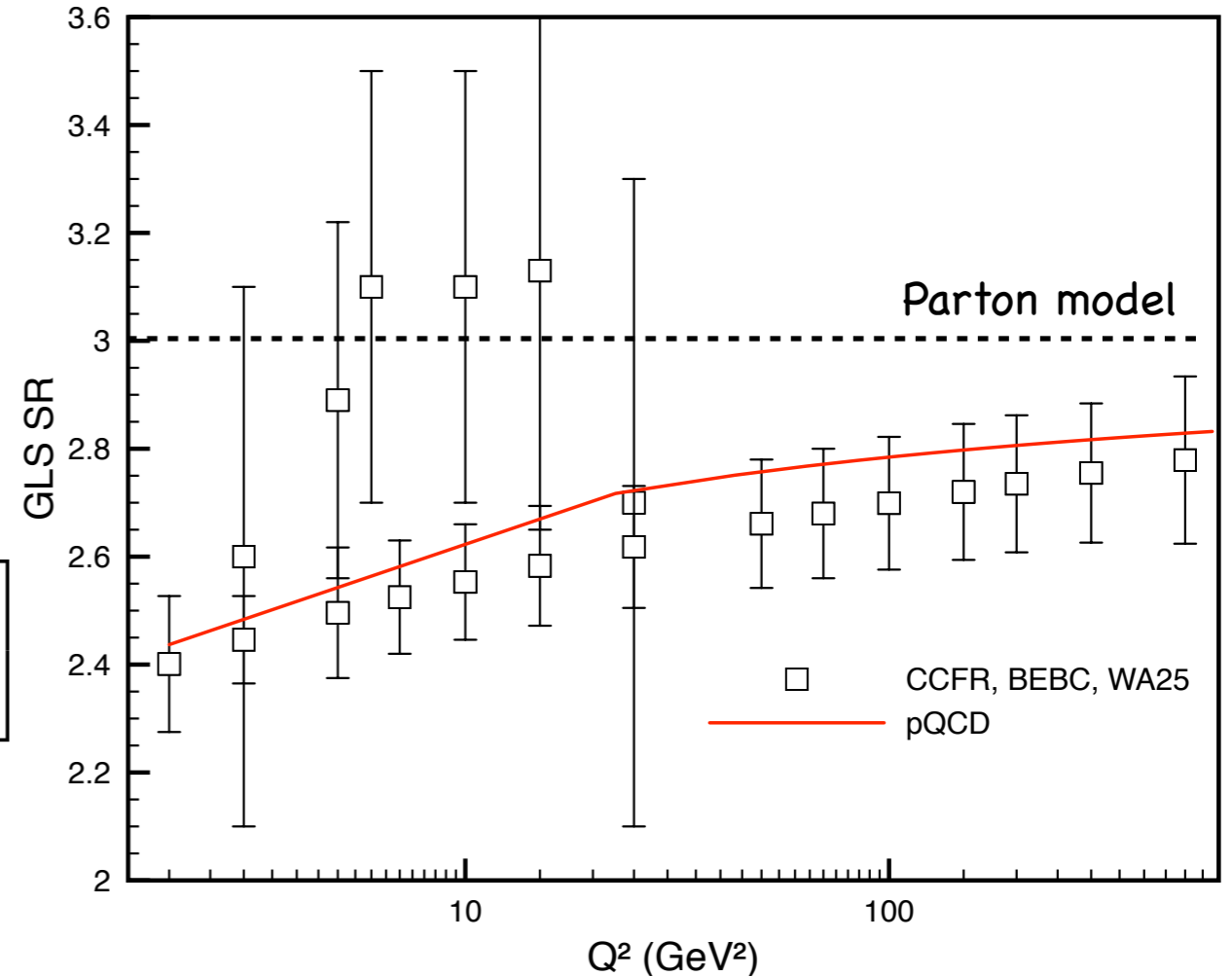
$$\frac{d^2\sigma^{\nu(\bar{\nu})}}{dxdy} = \frac{G_F^2 ME}{\pi} \left[xy^2 F_1 + \left(1 - y - \frac{Mxy}{2E}\right) F_2 \pm x \left(y - \frac{y^2}{2}\right) F_3 \right]$$

$$\sigma^{\nu p} - \sigma^{\bar{\nu} p} \sim F_3^{\nu p} + F_3^{\bar{\nu} p} = u_v^p(x) + d_v^p(x) \qquad \int_0^1 dx (u_v^p(x) + d_v^p(x)) = 3$$

Plot vs. data derived from nu-DIS

Including pQCD corrections

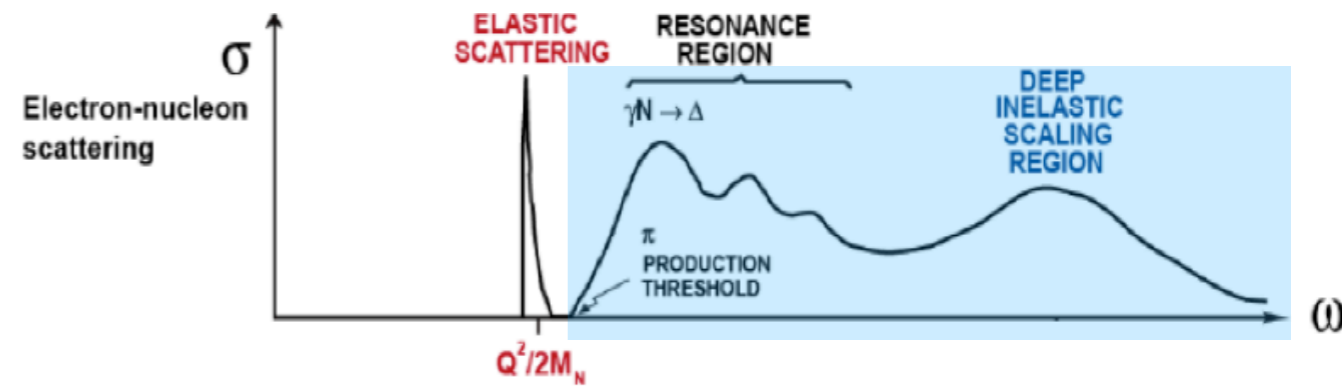
$$\text{GLS SR} = 3 \left[1 - \frac{\alpha_s^{\overline{MS}}}{\pi} - C_2 \left(\frac{\alpha_s^{\overline{MS}}}{\pi}\right)^2 - C_3 \left(\frac{\alpha_s^{\overline{MS}}}{\pi}\right)^3 \right]$$



Physics input

Inelastic contributions beyond DIS

$$\square_{\gamma W}^{\text{low } Q^2} = \frac{\alpha}{\pi} \int_0^{\Lambda^2} dQ^2 \int_{\nu_\pi}^{\infty} \frac{d\nu}{(\nu + q)^2} \frac{\nu + 2q}{M\nu} F_3^{(0)}$$



Physics input

Inelastic contributions beyond DIS

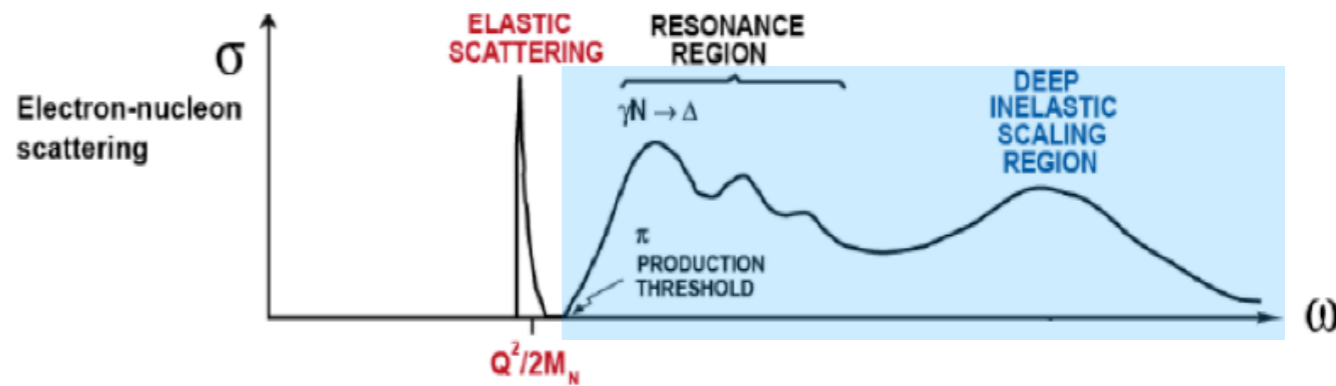
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M&S interpolating contribution

Constraints:

- I $F^{\text{INT}}(\Lambda^2) = F^{\text{DIS}}(\Lambda^2)$
- II $F^{\text{INT}}((0.823 \text{ GeV})^2) = F^{\text{Born}}((0.823 \text{ GeV})^2)$
- III $F^{\text{INT}}(0) = 0.$



$$\square_{\gamma W}^{VA(0)} = \frac{\alpha}{8\pi} \int_{Q_2^2}^{\Lambda^2} dQ^2 F^{\text{INT}}(Q^2)$$

3 Eqs. -> 3 free parameters

$$F^{\text{INT}}(Q^2) = -\frac{1.490}{Q^2 + m_\rho^2} + \frac{6.855}{Q^2 + m_A^2} - \frac{4.414}{Q^2 + m_{\rho'}^2}$$

Physics input

Inelastic contributions beyond DIS

$$\square_{\gamma W}^{\text{low } Q^2} = \frac{\alpha}{\pi} \int_0^{\Lambda^2} dQ^2 \int_{\nu_\pi}^{\infty} \frac{d\nu}{(\nu + q)^2} \frac{\nu + 2q}{M\nu} F_3^{(0)}$$



M&S interpolating contribution

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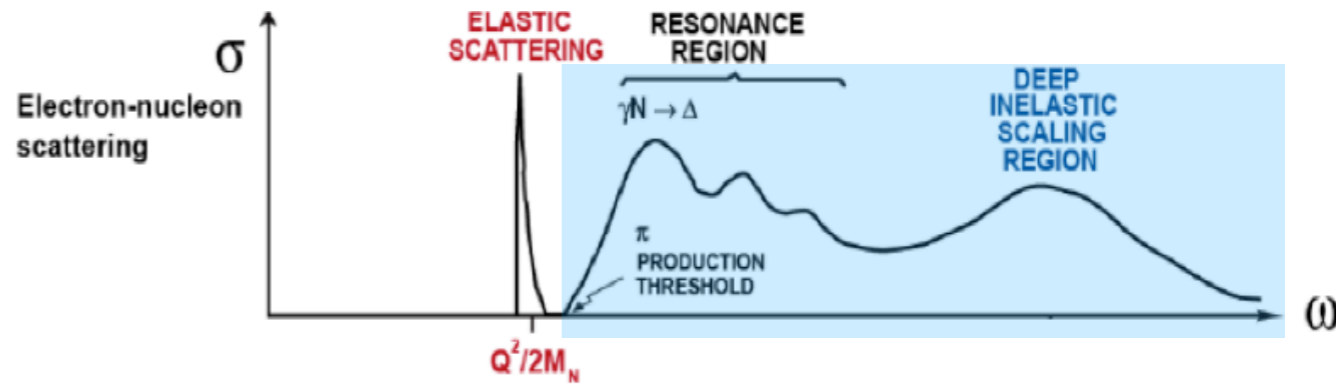
- I $F^{\text{INT}}(\Lambda^2) = F^{\text{DIS}}(\Lambda^2)$
- II $F^{\text{INT}}((0.823 \text{ GeV})^2) = F^{\text{Born}}((0.823 \text{ GeV})^2)$
- III $F^{\text{INT}}(0) = 0$.

Only constraint I is justified!

II: no reason Born is the whole story below some arbitrary Q^2

III: M&S claim it is required by chiral symmetry

Check in ChPT at one-loop $\int_{\nu_\pi}^{\infty} \frac{d\nu}{\nu^2} F_3^{(0)}(\nu, Q^2 = 0) = \frac{2M}{Q^2} \int_0^{x_\pi} F_3^{(0)}(x, Q^2) \Big|_{Q^2 \rightarrow 0} = 0$



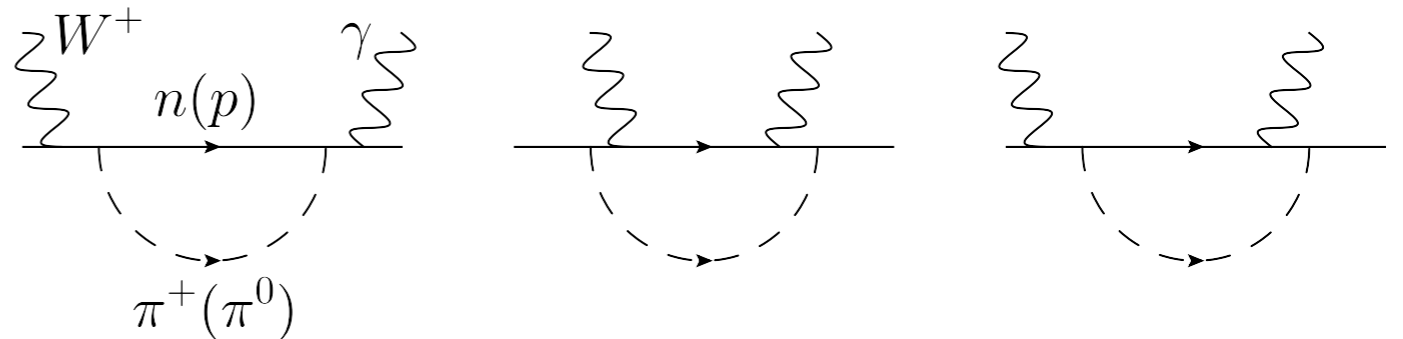
$$\square_{\gamma W}^{VA(0)} = \frac{\alpha}{8\pi} \int_{Q_2^2}^{\Lambda^2} dQ^2 F^{\text{INT}}(Q^2)$$

3 Eqs. \rightarrow 3 free parameters

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ChPT - check if $F(0)=0$

Representative Feynman graphs

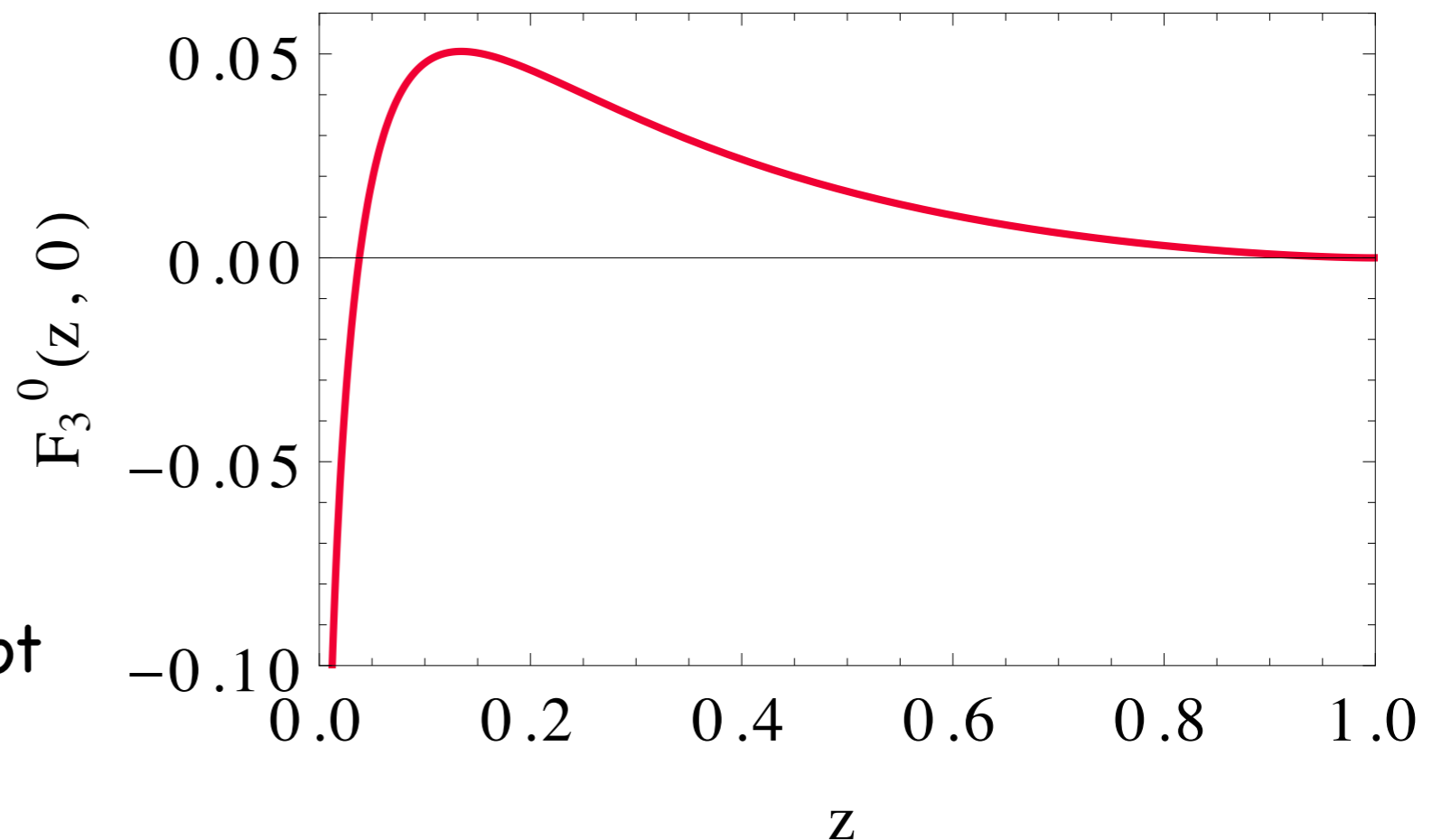


To visualize: change integration variable to $z = \nu_\pi / \nu$

$$\int_{\nu_\pi}^{\infty} \frac{d\nu \nu_\pi}{\nu^2} F_3^{(0)}(\nu, Q^2 = 0) = \int_0^1 dz F_3^{(0)}(\nu_\pi/z, Q^2 = 0)$$

The surface below and above the curve is not the same

The claim of M&S about the superconvergence relation is not supported by ChPT



Inelastic states beyond DIS

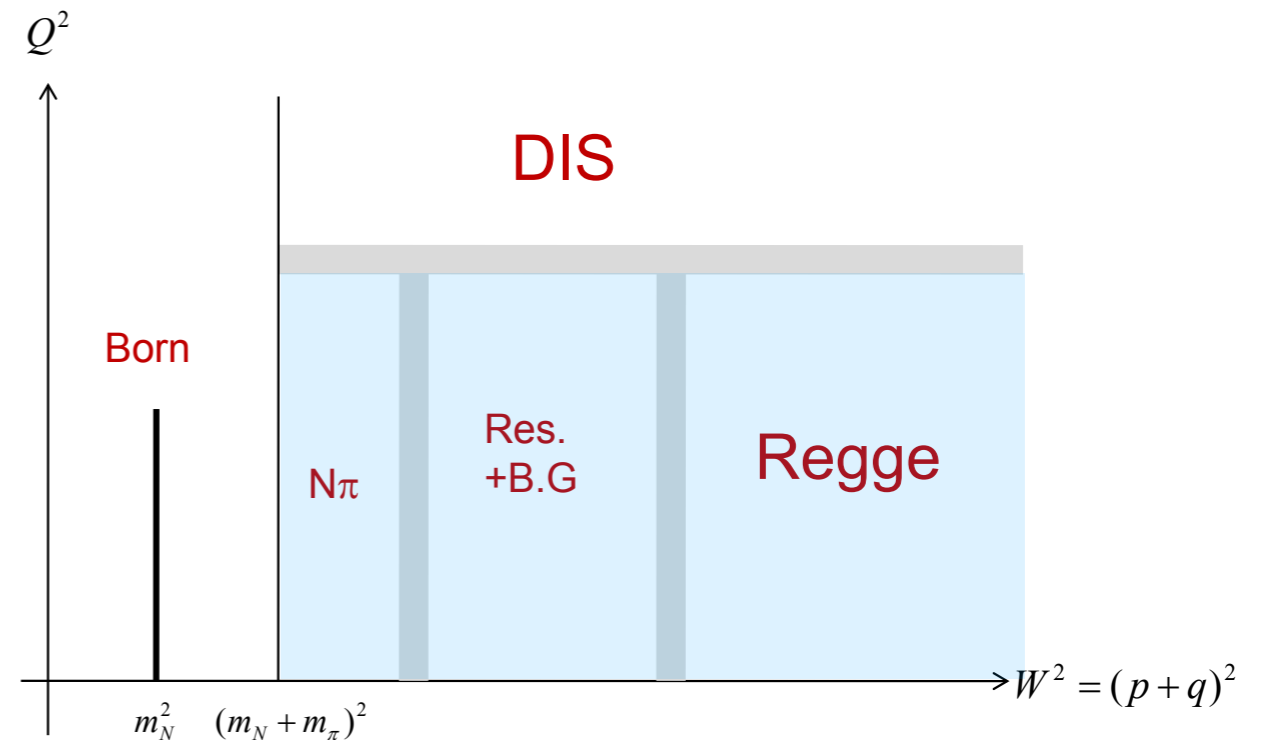
Our approach

- saturate F^{INT} by hadronic states:

low-energy π -N continuum

I=1/2 resonances

dominant Regge exchange (ρ)



π N contribution to the box:

$$\square_{\gamma W}^{\pi N} = \frac{\alpha}{2\pi} 0.044(4)$$

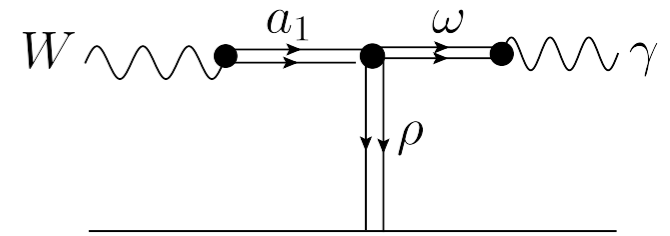
I=1/2 resonances: tiny contribution!

$$\square_{\gamma W}^{\text{Res}} \leq \frac{\alpha}{2\pi} 0.01$$

Inelastic states beyond DIS

Regge exchange

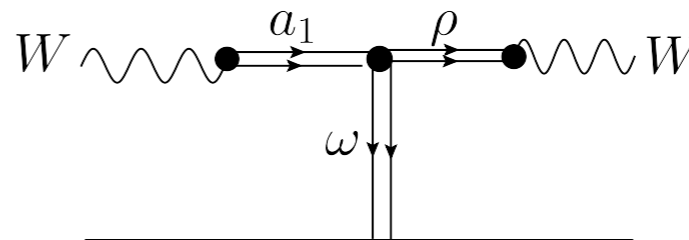
$$F_3^{(0),\text{Regge}}(\nu, Q^2) = C_R(Q^2) \left(\frac{\nu}{\nu_0} \right)^{\alpha_\rho}$$



Vector/axial vector dominance:

Stodolsky, Piketty '70

$$C_R^{\text{VDM}}(Q^2) = C_R(0) \frac{m_\omega^2}{m_\omega^2 + Q^2} \frac{m_{a_1}^2}{m_{a_1}^2 + Q^2}$$



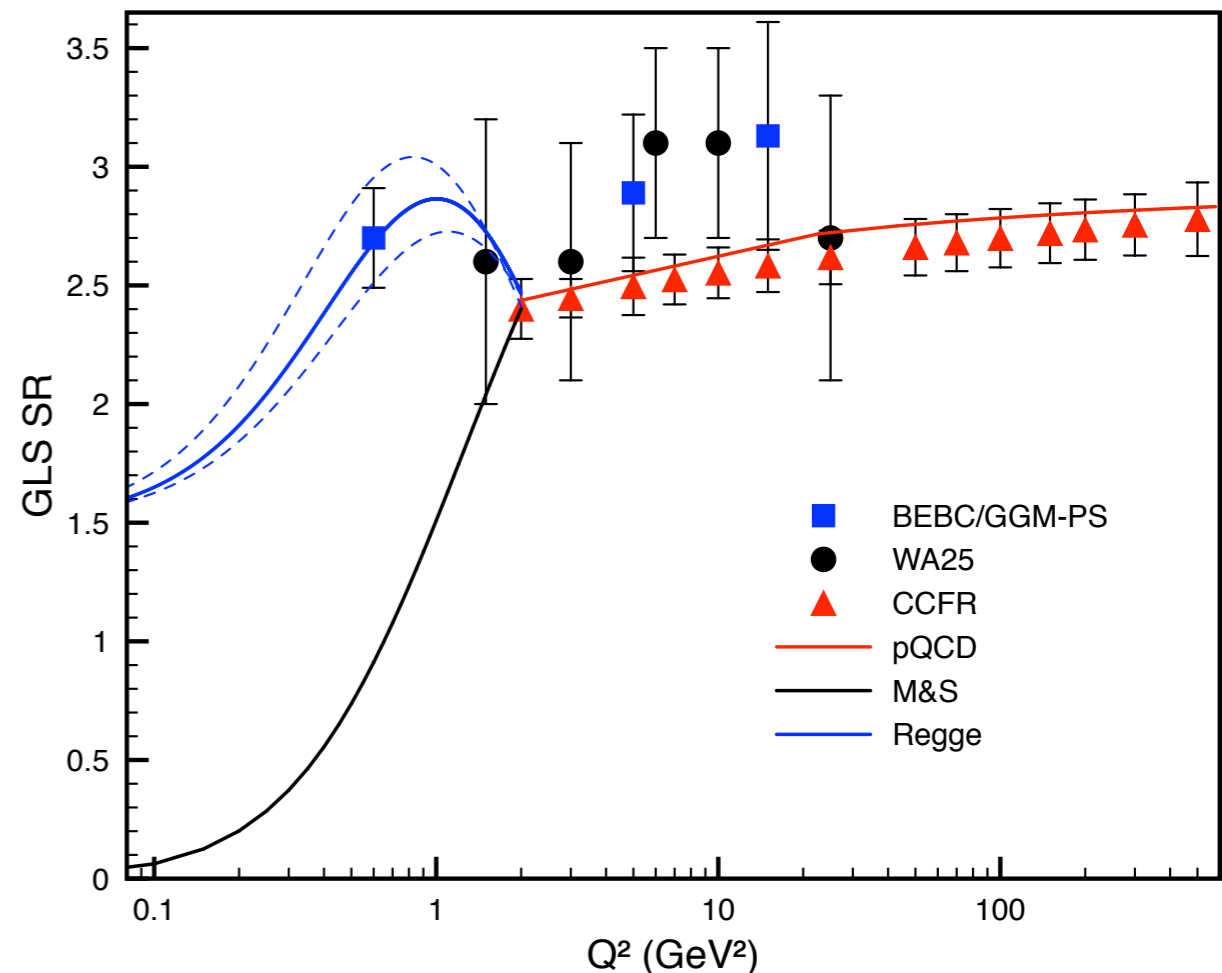
Guidance from GLS sum rule:
CC process, A-V interference

Only overall normalization is changed!
Match to pQCD at $Q^2=2\text{GeV}^2$
VDM + fit to (few) low- Q^2 data below

$$C_R(Q^2) = C_R^{\text{VDM}}(Q^2) \times h(Q^2)$$

Deduce $h(Q^2)$, $\Delta h(Q^2)$ from data!

$$\square_{\gamma W}^{\text{Regge}} = \frac{\alpha}{2\pi} 0.238(14)$$



γW -box on free neutron

Marciano & Sirlin '06

$$\square_{\gamma W}^{VA} = \frac{\alpha}{2\pi} [c_B + c_{int} + c_{DIS}] = \frac{\alpha}{2\pi} [0.83(8) + 0.14(14) + 1.84(0)]$$

$$\square_{\gamma W}^{MS} = \frac{\alpha}{2\pi} 2.79(17) = 3.24(20) \times 10^{-3}$$

New evaluation

$$\square_{\gamma W}^{VA} = \frac{\alpha}{2\pi} [c_B + c_{piN} + c_{Res} + c_{Regge} + c_{DIS}] = \frac{\alpha}{2\pi} [0.91(5) + 0.044(5) + 0.01(1) + 0.238(14) + 1.84(0)]$$

$$\square_{\gamma W}^{New} = \frac{\alpha}{2\pi} 3.03(5) = 3.51(6) \times 10^{-3}$$

V_{ud} from free n: about 1 sigma smaller

Numbers are preliminary but all crucial ingredients are in place.

Central value shifted by 1 sigma; uncertainty is likely to be reduced by factor 3

Currently the uncertainty for neutron decay is dominated by the experiment

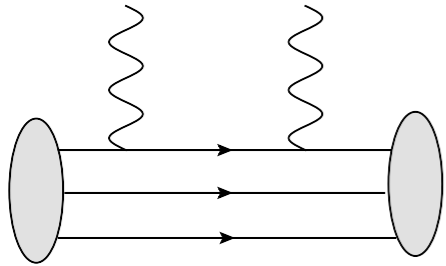
Nuclear β -decay

General structure of RC for nuclear decay (see John's talk)

$$1 + \text{RC} = (1 + \delta_R)(1 - \delta_C)(1 + \Delta).$$

Nuclear β -decay

General structure of RC for nuclear decay (see John's talk)

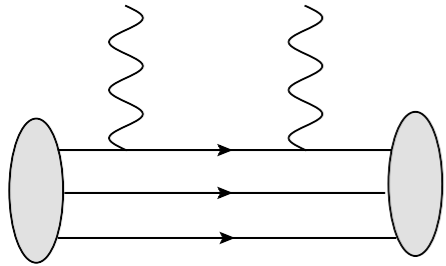


$$1 + \text{RC} = (1 + \delta_R)(1 - \delta_C)(1 + \Delta).$$

Universal (no NS)

Nuclear β -decay

General structure of RC for nuclear decay (see John's talk)



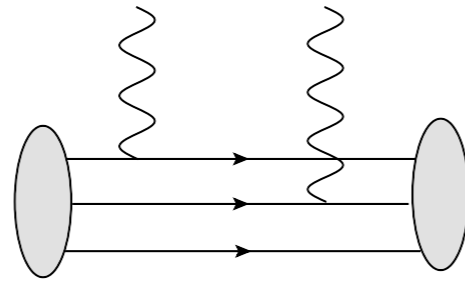
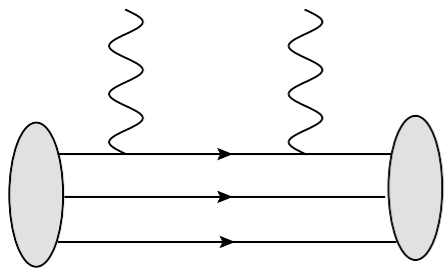
$$1 + \text{RC} = (1 + \delta_R)(1 - \delta_C)(1 + \Delta).$$

Isospin breaking

Universal (no NS)

Nuclear β -decay

General structure of RC for nuclear decay (see John's talk)



$$1 + \text{RC} = (1 + \delta_R)(1 - \delta_C)(1 + \Delta).$$

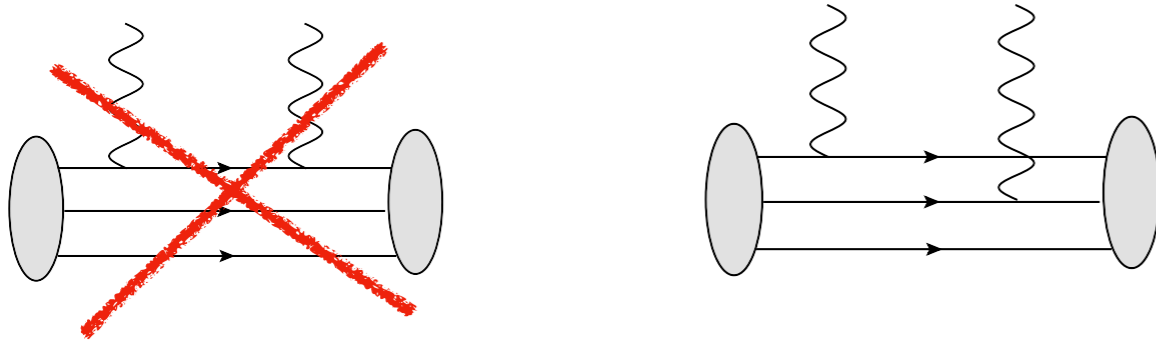
Isospin breaking

Nuclear structure (NS)

Universal (no NS)

Nuclear β -decay

General structure of RC for nuclear decay (see John's talk)



Nuclear Green fn: only with 2 active N

$$1 + \text{RC} = (1 + \delta_R)(1 - \delta_C)(1 + \Delta).$$

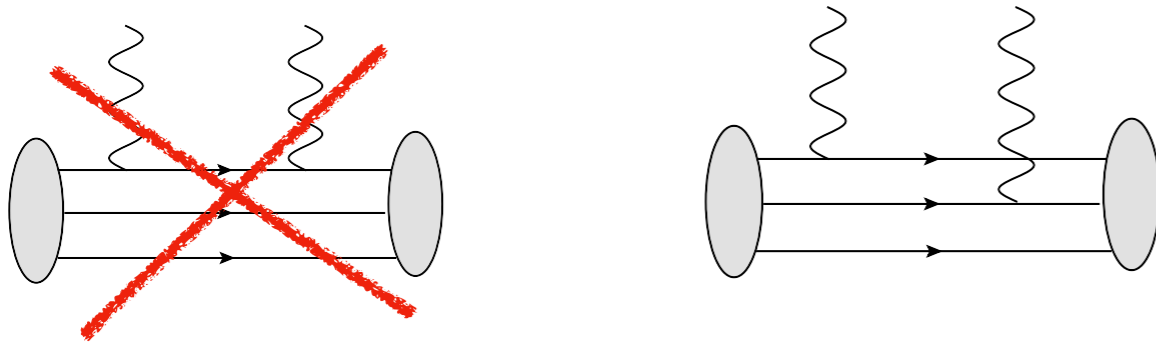
Isospin breaking

Nuclear structure (NS)

Universal (no NS)

Nuclear β -decay

General structure of RC for nuclear decay (see John's talk)



$$1 + \text{RC} = (1 + \delta_R)(1 - \delta_C)(1 + \Delta).$$

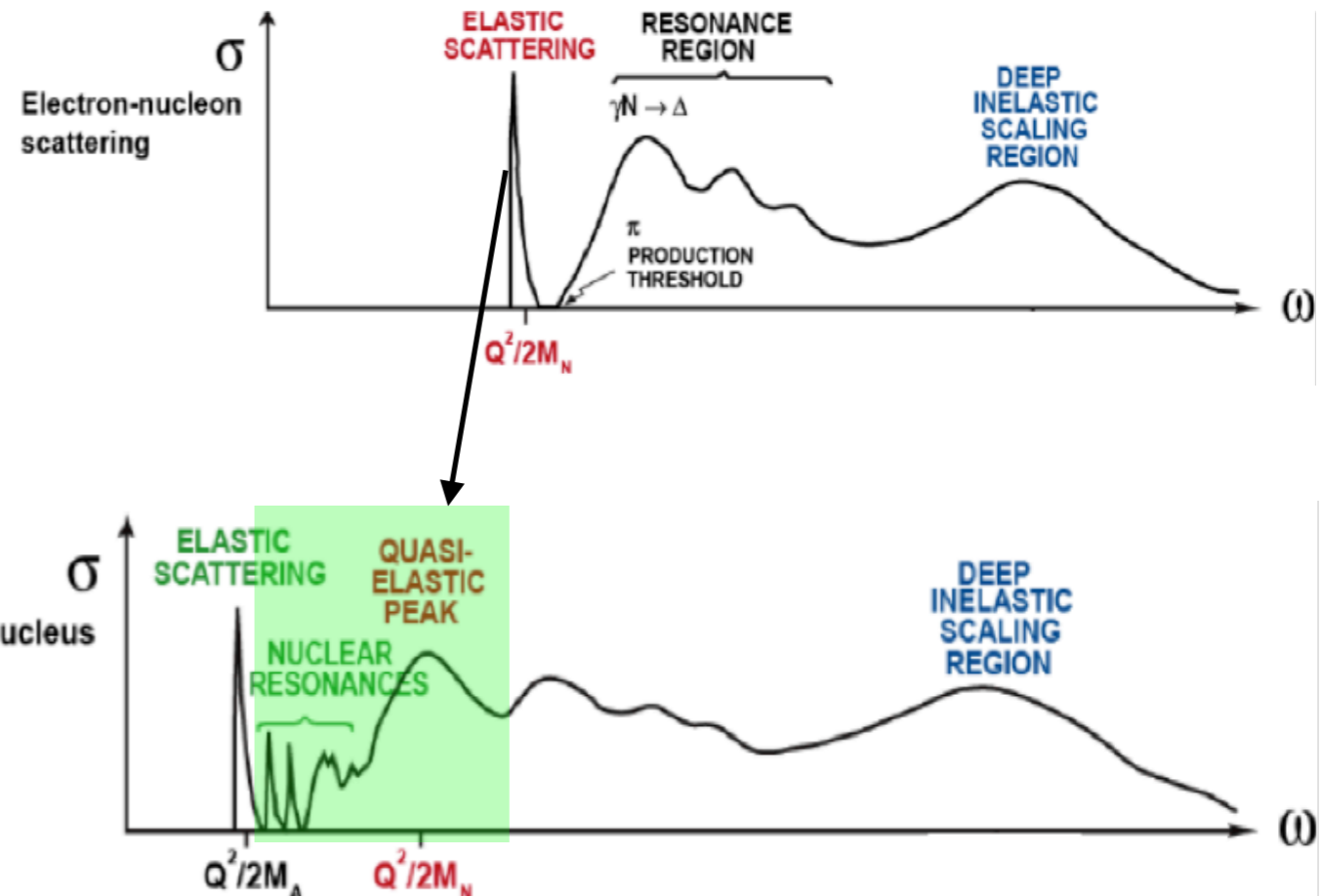
Nuclear structure (NS)
Universal (no NS)

Nuclear Green fn: only with 2 active N

But data tell us differently:
prominent broad QE peak -
mostly 1N knock-out

$$\square_{\gamma W}^{\text{Nucl.}} = \frac{\alpha}{\pi} \int_0^{\Lambda^2} dQ^2 \int_{\nu_N}^{\infty} \frac{d\nu}{(\nu + q)^2} \frac{\nu + 2q}{M\nu} F_3^{(0)}.$$

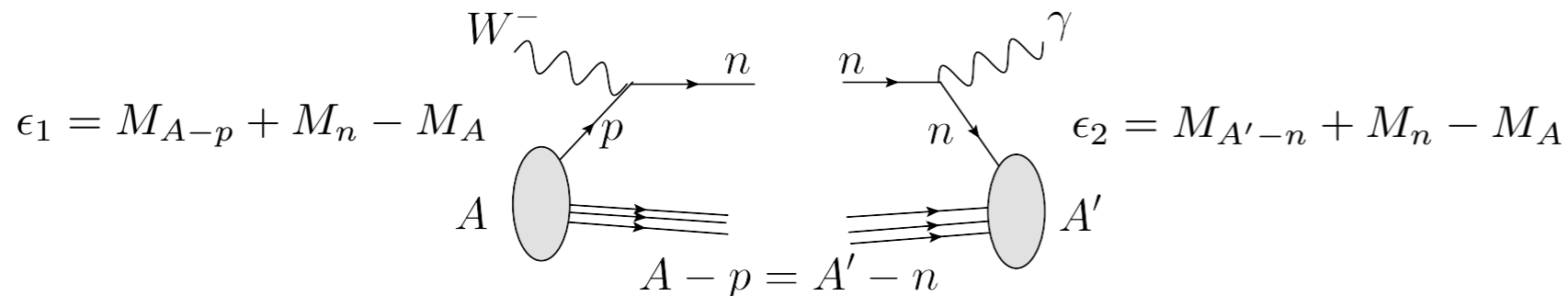
QE peak is common for all nuclei
Modify the universal correction
to account for bulk QE effect



QE contribution to γW -box

Bulk nuclear properties: Fermi momentum and break-up threshold

20 decays: $^{10}\text{C} \rightarrow ^{10}\text{B}$ through $^{74}\text{Rb} \rightarrow ^{74}\text{Kr}$ (Towner&Hardy '14 review)



$$\bar{\epsilon} = \sqrt{\epsilon_1 \epsilon_2}$$

Decay	ϵ_2 (MeV)	ϵ_1 (MeV)	$\bar{\epsilon}$ (MeV)
$^{10}\text{C} \rightarrow ^{10}\text{B}$	8.44	4.79	6.36
$^{14}\text{O} \rightarrow ^{14}\text{N}$	10.55	5.41	7.55
$^{18}\text{Ne} \rightarrow ^{18}\text{F}$	9.15	4.71	6.56
$^{22}\text{Mg} \rightarrow ^{22}\text{Na}$	11.07	6.28	8.34
$^{26}\text{Si} \rightarrow ^{26}\text{Al}$	11.36	6.30	8.46
$^{30}\text{S} \rightarrow ^{30}\text{P}$	11.32	5.18	7.66
$^{34}\text{Ar} \rightarrow ^{34}\text{Cl}$	11.51	5.44	7.91
$^{38}\text{Ca} \rightarrow ^{38}\text{K}$	12.07	5.33	8.02
$^{42}\text{Ti} \rightarrow ^{42}\text{Sc}$	11.55	4.55	7.25
$^{26m}\text{Al} \rightarrow ^{26}\text{Mg}$	11.09	6.86	8.72
$^{34}\text{Cl} \rightarrow ^{34}\text{S}$	11.42	5.92	8.22
$^{38m}\text{K} \rightarrow ^{38}\text{Ar}$	11.84	5.79	8.28
$^{42}\text{Sc} \rightarrow ^{42}\text{Ca}$	11.48	5.05	7.61
$^{46}\text{Va} \rightarrow ^{46}\text{Ti}$	13.19	6.14	9.00
$^{50}\text{Mn} \rightarrow ^{50}\text{Cr}$	13.00	5.37	8.35
$^{54}\text{Co} \rightarrow ^{54}\text{Fe}$	13.38	5.13	8.28
$^{62}\text{Ga} \rightarrow ^{62}\text{Zn}$	12.90	3.72	6.94
$^{66}\text{As} \rightarrow ^{66}\text{Ge}$	13.29	3.16	6.48
$^{70}\text{Br} \rightarrow ^{70}\text{Se}$	13.82	3.20	6.65
$^{74}\text{Rb} \rightarrow ^{74}\text{Kr}$	13.85	3.44	6.90

Effective removal energies - all in a small range

$$\bar{\epsilon} = 7.68 \pm 1.32 \text{ MeV}$$

Fermi momentum also not too different for all A

$$k_F(A = 10) = 228 \text{ MeV}, \quad k_F(A = 74) = 245 \text{ MeV}$$

Can define a universal correction that correctly represents bulk nuclear effect!

Further ingredients:

Free Fermi gas model (or superscaling)

+ Pauli blocking

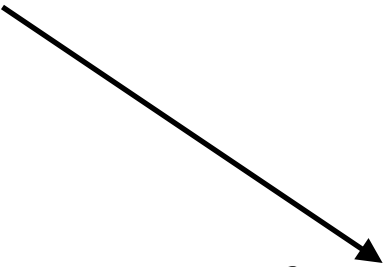
QE contribution to γW -box

γW -box for bound neutron: $\square_{\gamma W}^{\text{free n}} = \frac{\alpha}{2\pi} 0.91(5) \rightarrow \square_{\gamma W}^{\text{QE}} = \frac{\alpha}{2\pi} 0.44(4)$

Reduction: finite breakup threshold $\int_{\frac{Q^2}{2M}} \frac{d\nu}{\nu^2} F_3^n \rightarrow \int_{\frac{Q^2}{2M_A} + \bar{\epsilon}} \frac{d\nu}{\nu^2} F_3^{\text{Nucl}}$

New formulation of the γW -box:

$$\square_{\gamma W}^{MS} = \frac{\alpha}{2\pi} 2.79(17) = 3.24(20) \times 10^{-3}$$


$$\square_{\gamma W}^{\text{Nucl. New}} = \frac{\alpha}{2\pi} 2.56(4) = 2.97(5) \times 10^{-3}$$

A mere shift by 1 sigma; uncertainty significantly reduced.

Nuclear Structure corrections should be revisited and possibly redefined

V_{ud} from superallowed β : 1 sigma larger

Summary

- New dispersive representation of the γW -box
- Data driven uncertainties
- Crucial input: GLS sum rule
- New formulation of RC for V_{ud} extraction:
overall small effect; uncertainty significantly reduced
- Nuclear Structure corrections may need to be reformulated
- Backup: can nuclear structure effects lead to additional energy dependence?

Turn “inner” correction inside-out?

γW -box correction at zero energy

$$\square_{\gamma W}^{VA(0)} = \frac{\alpha}{\pi M} \int_0^\infty \frac{dQ^2 M_W^2}{M_W^2 + Q^2} \int_0^\infty d\nu \frac{(\nu + 2q)}{\nu(\nu + q)^2} F_3^{(0)}(\nu, Q^2),$$

$$\square_{\gamma W}^{VA(3)} = 0,$$

γW -box correction with linear E-dependence

$$\text{Re } \square_{\gamma W}^{even} = \frac{\alpha_{em}}{\pi} \int_{\nu_{thr}}^\infty d\nu \int_0^\infty dQ^2 \frac{F_3^{(0)}}{2M\nu} \left(\frac{1}{E_{min}} - \frac{\nu}{4E_{min}^2} \right),$$

$$\text{Re } \square_{\gamma W}^{odd} = \frac{\alpha_{em}}{\pi} E \int_{\nu_{thr}}^\infty d\nu \int_0^\infty dQ^2 \left[\frac{F_1^{(0)}}{6ME_{min}^3} + \left(\frac{\sqrt{\nu^2 + Q^2}}{2E_{min}^2 \nu Q^2} - \frac{1}{12E_{min}^3 \nu} \right) F_2^{(0)} + \frac{F_3^{(-)}}{2M\nu} \left(\frac{1}{2E_{min}^2} - \frac{\nu}{6E_{min}^3} \right) \right]$$

$E_{min} = (\nu + \sqrt{\nu^2 + Q^2})/2$

Common wisdom: E-dep. negligible because should come as $(\alpha/2\pi) E/m_\pi < 10^{-5}$

But nuclear excitations live at few MeV \rightarrow large nuclear polarizabilities

$$\alpha_E + \beta_M = \frac{2\alpha_{em}}{M} \int \frac{d\omega}{\omega^3} F_1(\omega, Q^2 = 0) = 2\alpha_{em} \int \frac{d\omega}{\omega^2} \frac{F_2(\omega, Q^2)}{Q^2} \Big|_{Q^2=0}$$

New energy scale: polarizability/radius²

$$\text{Re } \square_{\gamma W}^{odd} \sim \frac{2}{\pi} E \frac{\alpha_E + \beta_M}{R_{Ch}^2}$$

$$R_{Ch} \sim 1.2 \text{ fm } A^{1/3}$$

$$\alpha_E \sim (2.2 \times 10^{-3} \text{ fm}) A^{5/3}$$

Expect

$$\text{Re } \square_{\gamma W}^{odd} \sim 1 \times 10^{-3} \left(\frac{E}{5 \text{ MeV}} \right) \left(\frac{A}{30} \right)$$