# Status and prospects of the China Dark Matter Experiment

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### Outline:

- Introduction to dark matter experiments
- CDEX Introduction
- CDEX-1 and CDEX-10
- CDEX nex-stage plan
- Summary



### Dark matter detection methods

# Direct Detection (Underground)



#### Indirect Detection (Space)



#### Accelerator (Man-made)



#### CJPL 🚈 No dark matter particles detected till now!

#### **Direct Detection Principle and Characteristics**



- Event rate quite small, and the deposited energy quite small;
- Event rate exponentially decreases with recoiled energy;
- <1cpkkd @1keV for 10GeV DM.</p>

CIPI



#### The trends for dark matter direct detection

- □ The sensitivities several order improved in last years ;
- Lower background level and larger target mass needed for high mass DM search experiment;
- Lower energy threshold, low background level and larger mass target needed for light DM particle search.



### **CDEX: China Dark matter EXperiment**

#### Established in 2009.

- Tsinghua University (THU)
- Sichuan University (SCU)
- Nankai University (NKU)
- China Institute of Atomic Energy (CIAE)
- Beijing Normal University (BNU)
- Yalong River Company





# CDEX stages

- <u>CDEX-1</u>: Development of pPC-HPGe detector, its background understanding;
- <u>CDEX-10</u>: Performances of HPGe array detector system;
- <u>CDEX-10X</u>: Fabrication of HPGe and Ge crystal growth by CDEX;
- <u>CDEX-100</u>: Ultra-low cosmogenic background and large LN<sub>2</sub> cooling and shielding system ;
- <u>CDEX-1T:</u> Multi-purpose experiment for dark matter and double beta decay.





# CDEX plan for DM



# CDEX-1 experiment

- 1. HPGe technology
  - Designed the first one single module 1kg-scale p-type pointcontact Ge detector (1kg-PPCGe) "prototype" C1A
  - ✓ Improved the second 1kg-PPCGe C1B
- 2. Active shielding technology: NaI(TI) used as anti-Compton detector
  - ✓ C1A 1kg-PPCGe run
  - ✓ C1 20g Ge + NaI(TI) run
  - ✓ C1A 1kg-PPCGe + NaI(TI) run







CDEX-1 CDEX-1A 1kg PCGe

#### CDEX-1B 1kg PCGe

#### 20 cm OFHC Copper 120 cm Lead

# **CDEX-1A&B** physical results

- Data set: >500 kg·day by C1A and >1000 kg·day by C1B:
  - Best SI/SD sensitivities achieved with pPCGe;
  - Best Axion sensitivity achieved below 1 keV;
  - $E_{th}$  improved from 475 eV  $\rightarrow$  160 eV;
  - AM result under preparation.



# CDEX-1 0vbb result

- Published <sup>76</sup>Ge 0vbb result based on CDEX-1 PCGe data;
- Calculation of the level of cosmogenic events @ 2MeV based on cosmogenic characteristic X-ray peaks <10keV;</li>



L. Wang, Q. Yue\*, et al. Sci. China Phys. Mech. Astron. (2017) 60: 071011



### **CDEX-10** experiment

- The important stage towards large-scale Ge experiment;
- Directly immersed into liquid nitrogen for cooling;
- Dataset: 102.8kg·d.



**CIPL** 





#### **CDEX-10** results

- $\checkmark$  E<sub>th</sub> down to 160eV,
- ✓ The best SI results @ 4-5GeV;
- ✓ The best SD results < 4GeV.





# CDEX-10 pPCGe detectors

- Collaborate with commercial companies, such as Canberra, ORTEC, BSI and so on;
- Home-made pPCGe for the lower cosmogenic backgrounds prepared at CJPL now.



T1

- Commercial Ge crystal;
- Structure machining;
- Li-drift and B-implanted;
- Home-made ULB PreAmp;
- Underground EF-Cu;
- Underground assemble;
- Underground testing...

### CDEX-1T@CJPL-II

- A future Ge detector composed of the PCGe detector array and LN shielding and cooling system in the CJPL-II
- Both Dark matter and Double Beta Decay





## Key technologies towards CDEX-1T

- Ge purification and crystal growth;
- HPGe detector fabrication;
- Ultra-low background VFE;
- Ultra-pure Cu for structure and cables;
- Large-volume cooling tank.













### CDEX-1T –Ge crystal growth





- The requirement for making P-type Ge detector
  - ✓ Impurity density: ~10<sup>10</sup> cm<sup>-3</sup>
  - ✓ Dislocation: <5000 cm<sup>-2</sup>
- CDEX are working on this two points.



# **HPGe** detector fabrication

 First 500g home-made pPCGe+ASIC finished testing, energy resolution and energy threshold compared with commercial one.







#### Cosmogenic Background of Ge detector

- Long-time ground preparation of detector induces high cosmogenic background level (~2cpkkd @2-4keV);
- Based on simulation, 2 months ground fabrication and transportation could decrease the <sup>3</sup>H continuous background level to ~10<sup>-2</sup>cpkkd @ 2-4 keV.



#### Cosmogenic Background of U-Ge detector

- Underground germanium crystal growth and detector fabrication could dramatically decrease the cosmogenic backgrounds from non-Ge isotopes, such as <sup>3</sup>He, <sup>65</sup>Zn;
- <sup>76</sup>Ge Enriched germanium material could help to decrease
  <sup>68</sup>Ge(<sup>68</sup>Ga) cosmogenic backgrounds too.



# CDEX ULB-Cu @ CJPL

- Setting up the facilities for ULB-Cu production;
- CDEX copper goal will be the Majorana EFCu purification: Th<0.06 µBq/kg、U<0.17 µBq/kg.</li>
- Shielded by LN2, Structure materials used as little as possible in order to lower the background contribution.







# **CDEX-1T Plan**





# CDEX-1T@CJPL



# Summary

- CDEX has achieved competitive DM physical results from 2013 on;
- The first physical results from CDEX-10 will be published by PRL in 2 weeks;
- The best SI sensitivities on 4-5GeV energy region and the best SD sensitivities below 4GeV;
- CDEX is developing the key technologies of low background Ge detector towards a ton-scale Ge experiment for DM (+DBD+Solar Neutrino).

