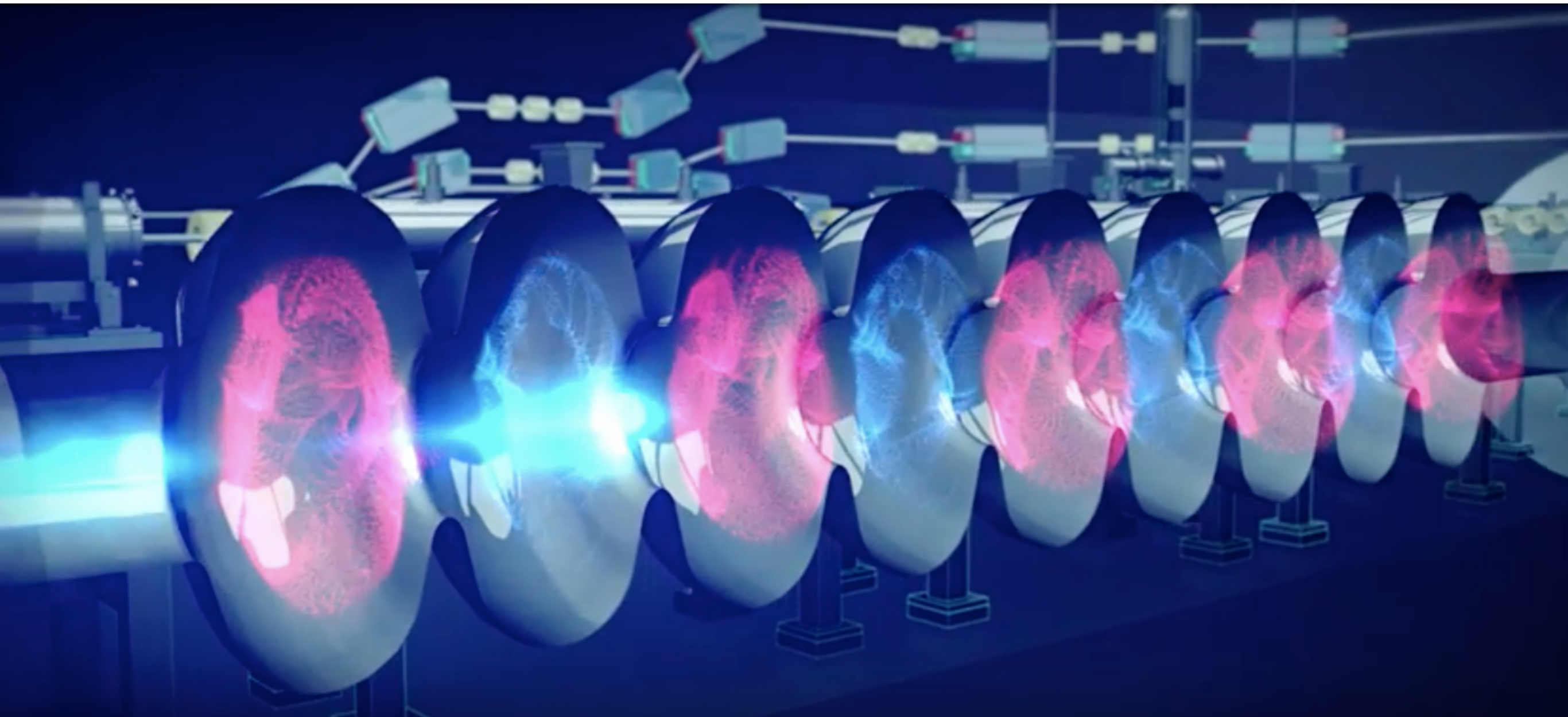


Dark Sector Searches at MESA

Luca Doria

Institut für Kernphysik

Johannes-Gutenberg Universität Mainz



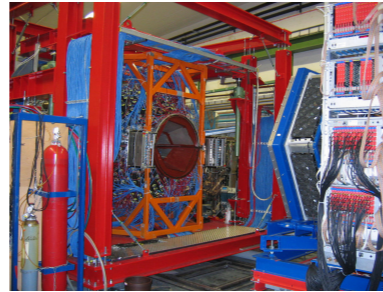
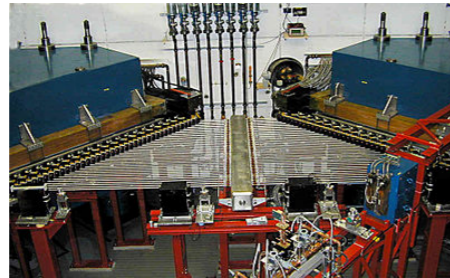
Cluster of Excellence

Precision Physics, Fundamental Interactions
and Structure of Matter



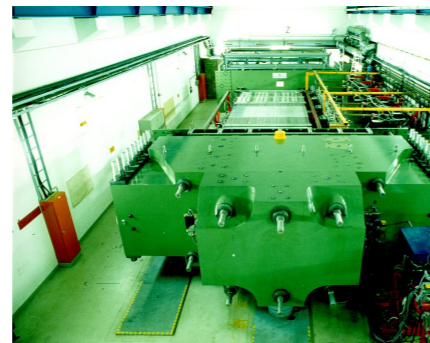
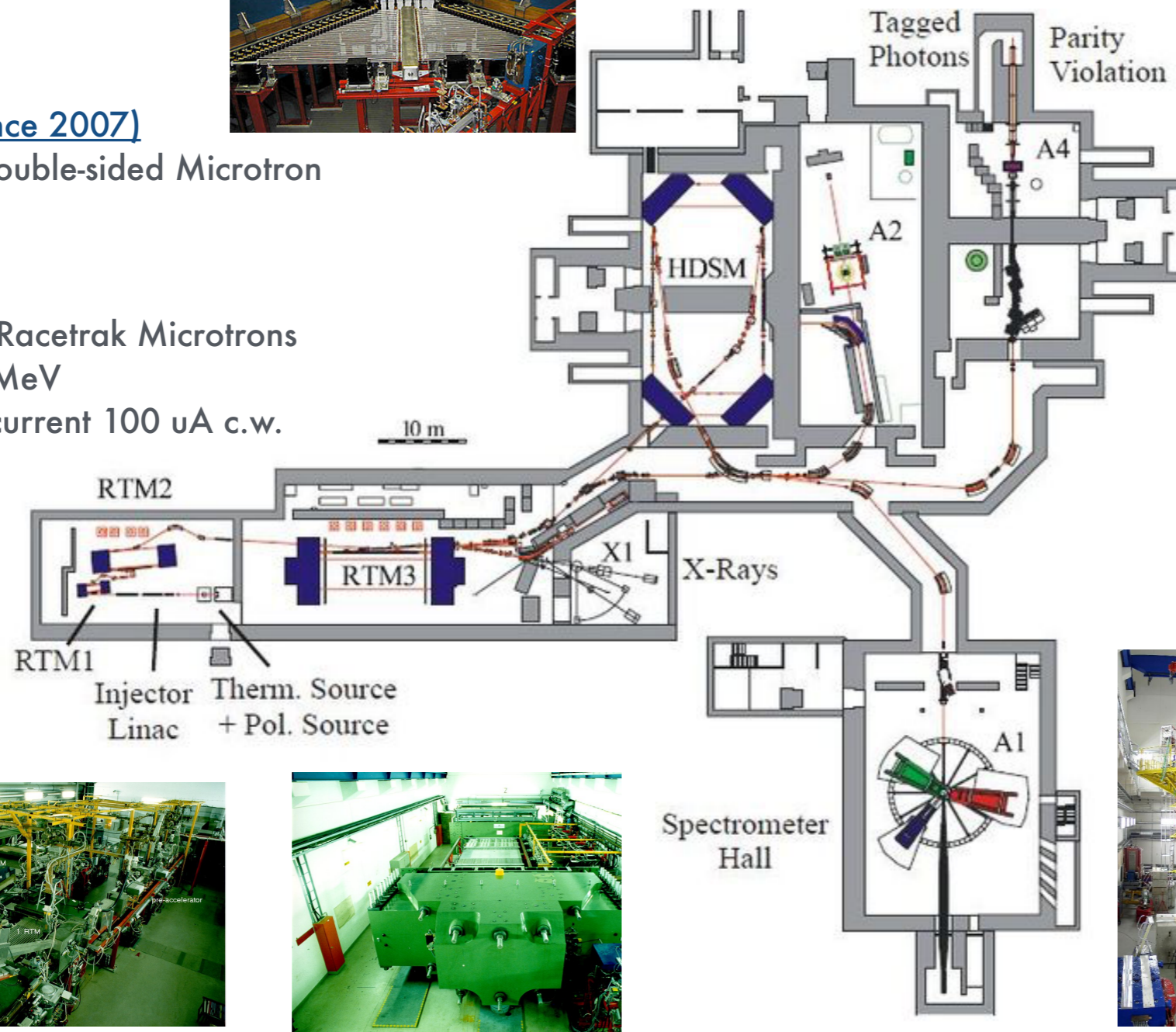
THE LOW-ENERGY FRONTIER
OF THE STANDARD MODEL

A2 Collaboration
Experiments with real photons



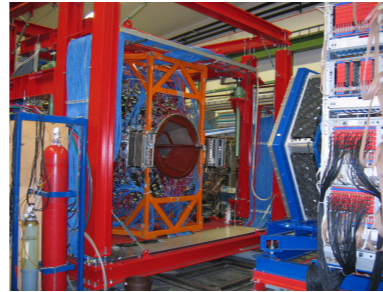
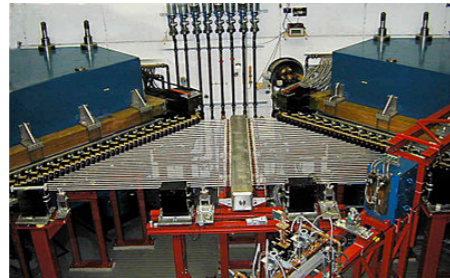
MAMI-C (since 2007)
Harmonic Double-sided Microtron
E= 1.5 GeV

MAMI-B
3 cascaded Racetrak Microtrons
E=180-883 MeV
Max beam current 100 μ A c.w.



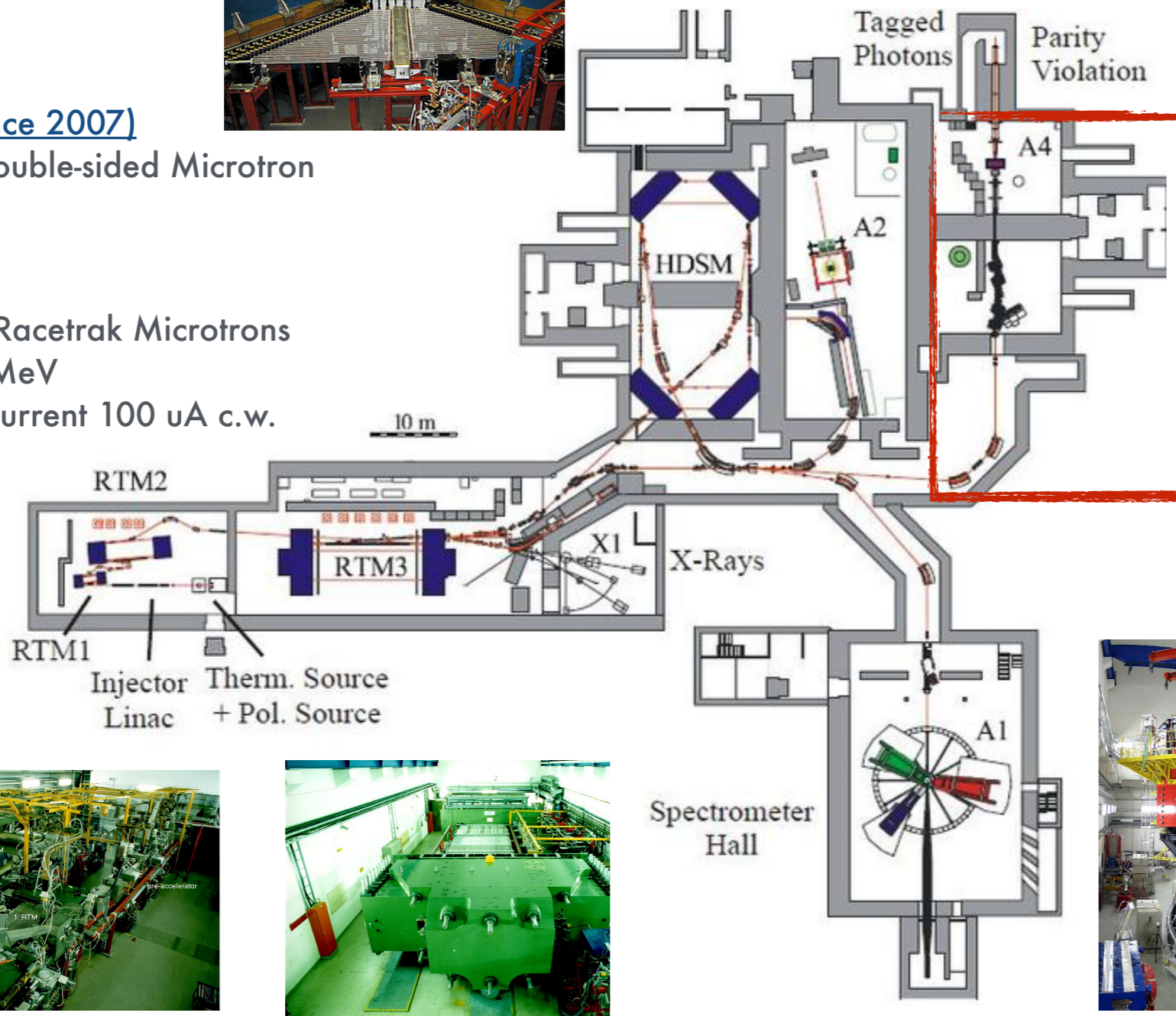
A1 Collaboration
3-spectrometer setup
Experiments with electrons

A2 Collaboration
Experiments with real photons

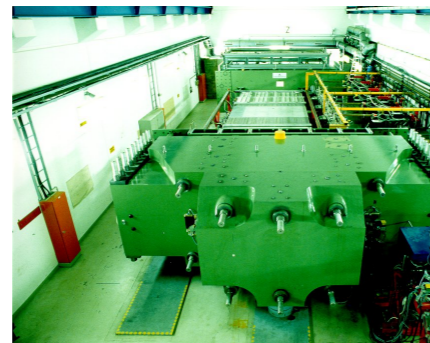


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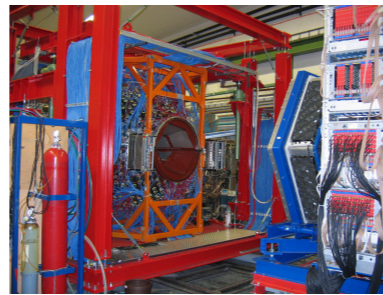
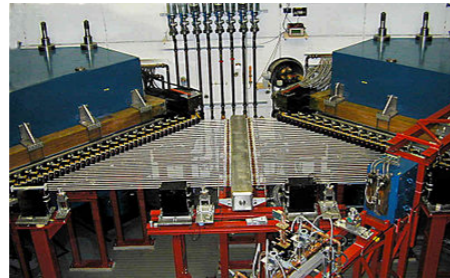


Existing Halls cleared for MESA



A1 Collaboration
3-spectrometer setup
Experiments with electrons

A2 Collaboration
Experiments with real photons

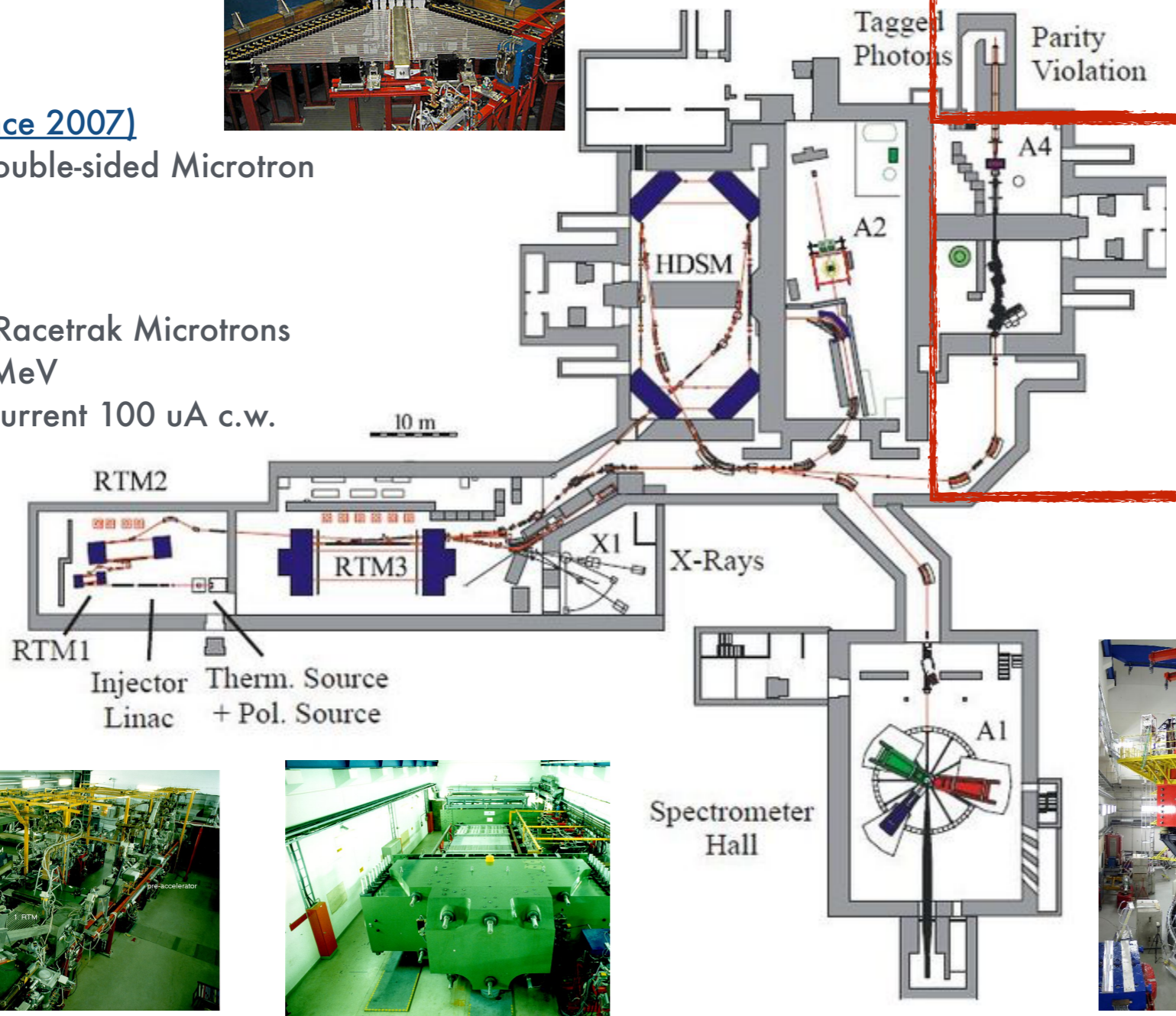


New MESA Hall and Building

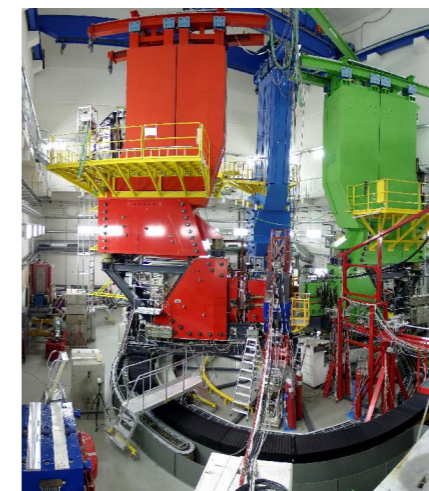
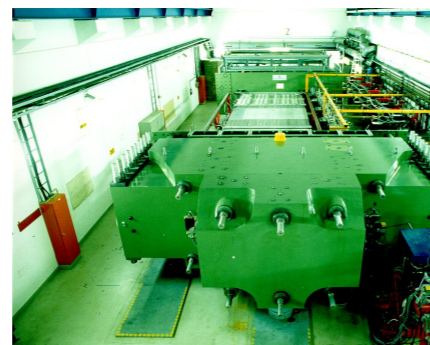
Existing High-power Beam Dump

MAMI-C (since 2007)
Harmonic Double-sided Microtron
E= 1.5 GeV

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Existing Halls cleared for MESA

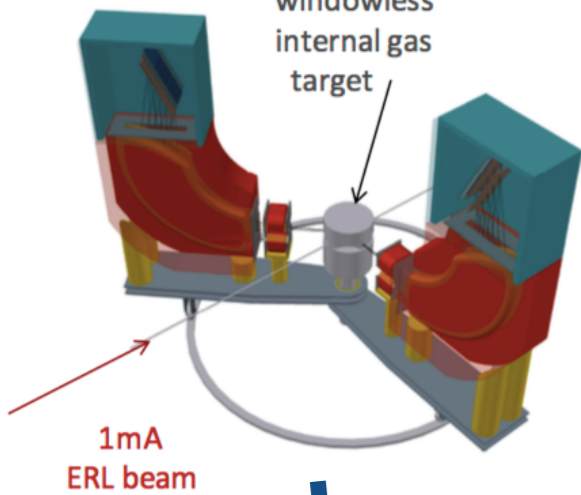


A1 Collaboration
3-spectrometer setup
Experiments with electrons

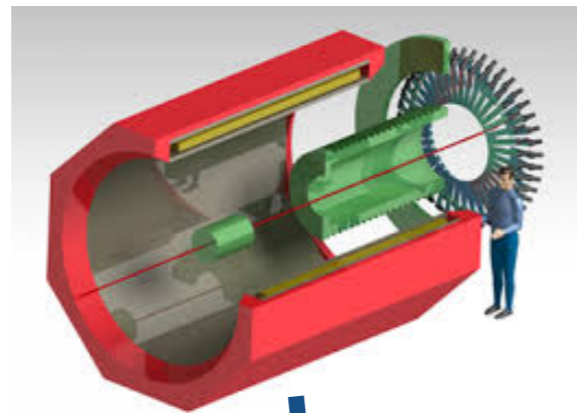
Mainz Energy-Recovering Superconducting Accelerator

MAGIX

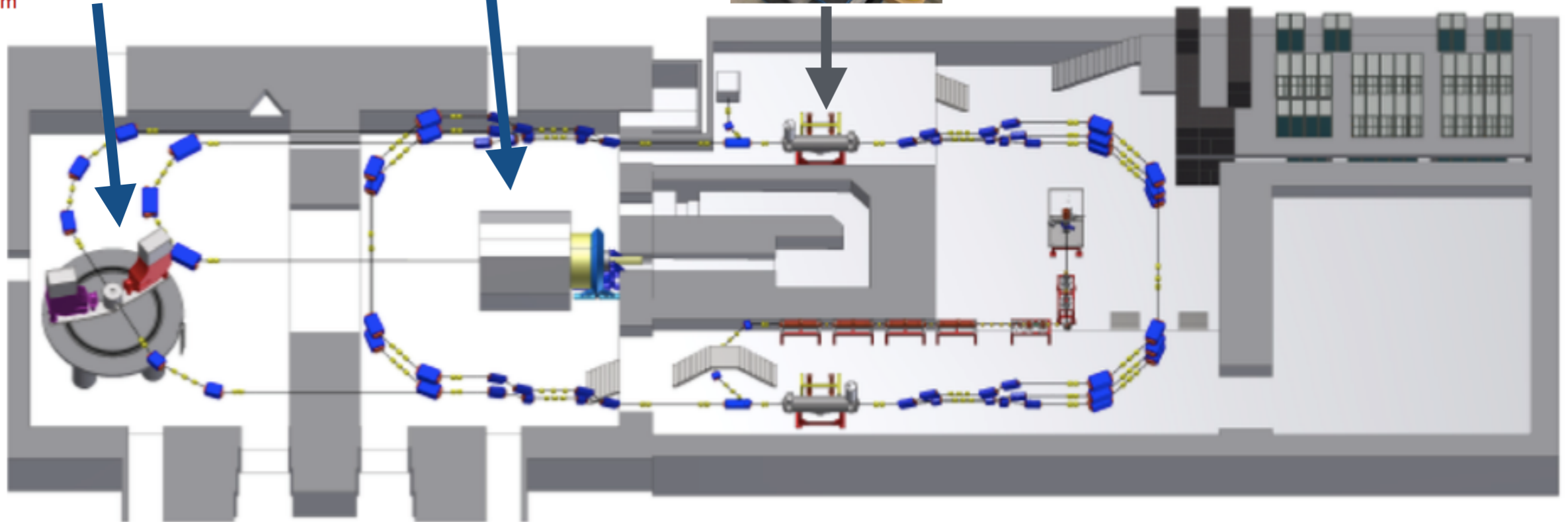
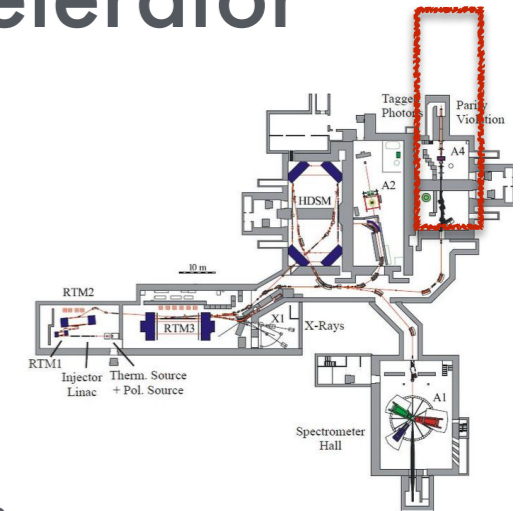
windowless
internal gas
target



P2



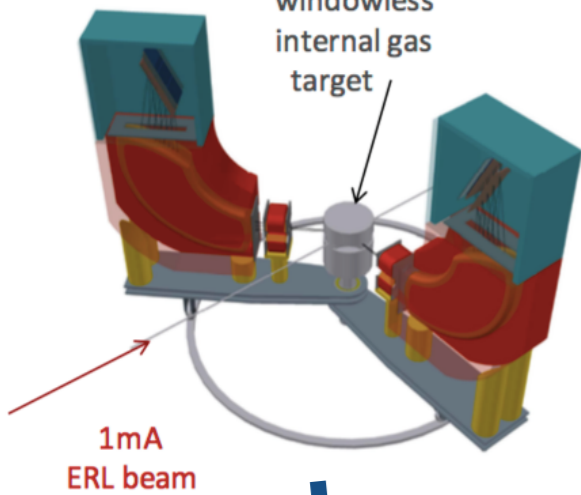
Superconducting Cavities:
 9-cell ,1.3 GHz
 12.5 MeV/m gain
 XFEL-TESLA type
 Modified Rossendorf-type Modules



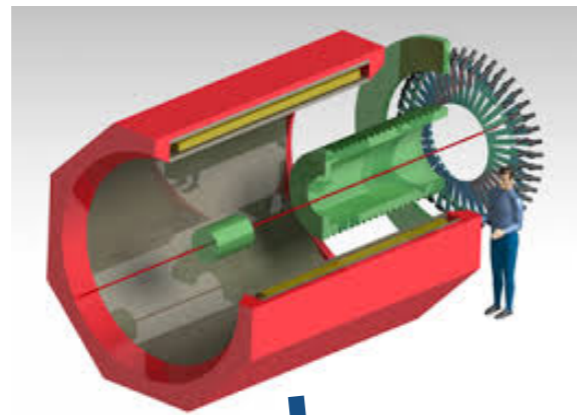
Mainz Energy-Recovering Superconducting Accelerator

MAGIX

windowless
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P2



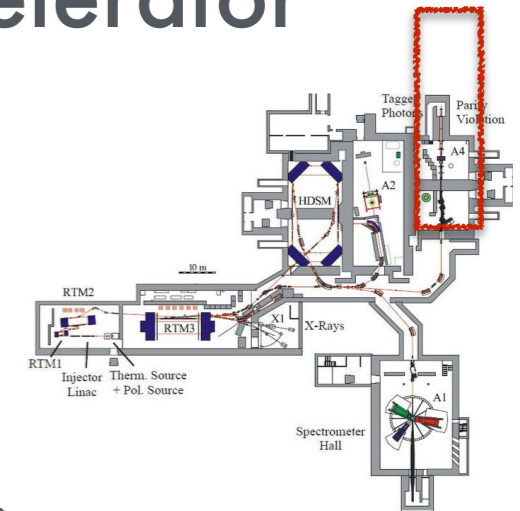
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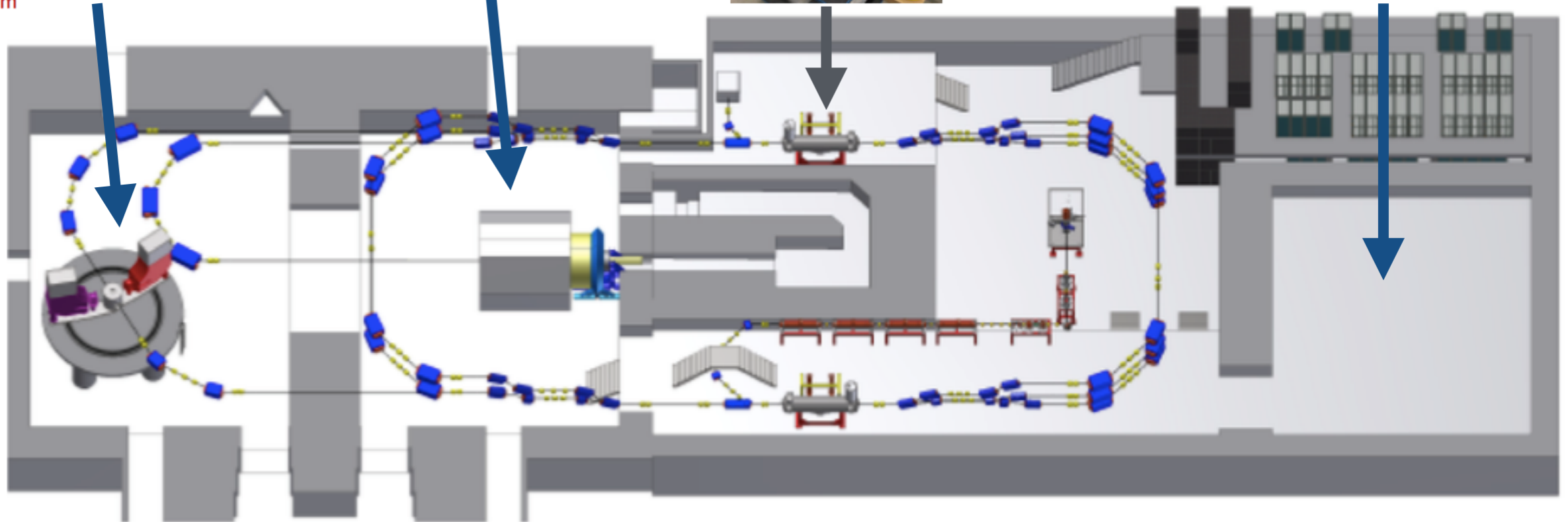
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XFEL-TESLA type

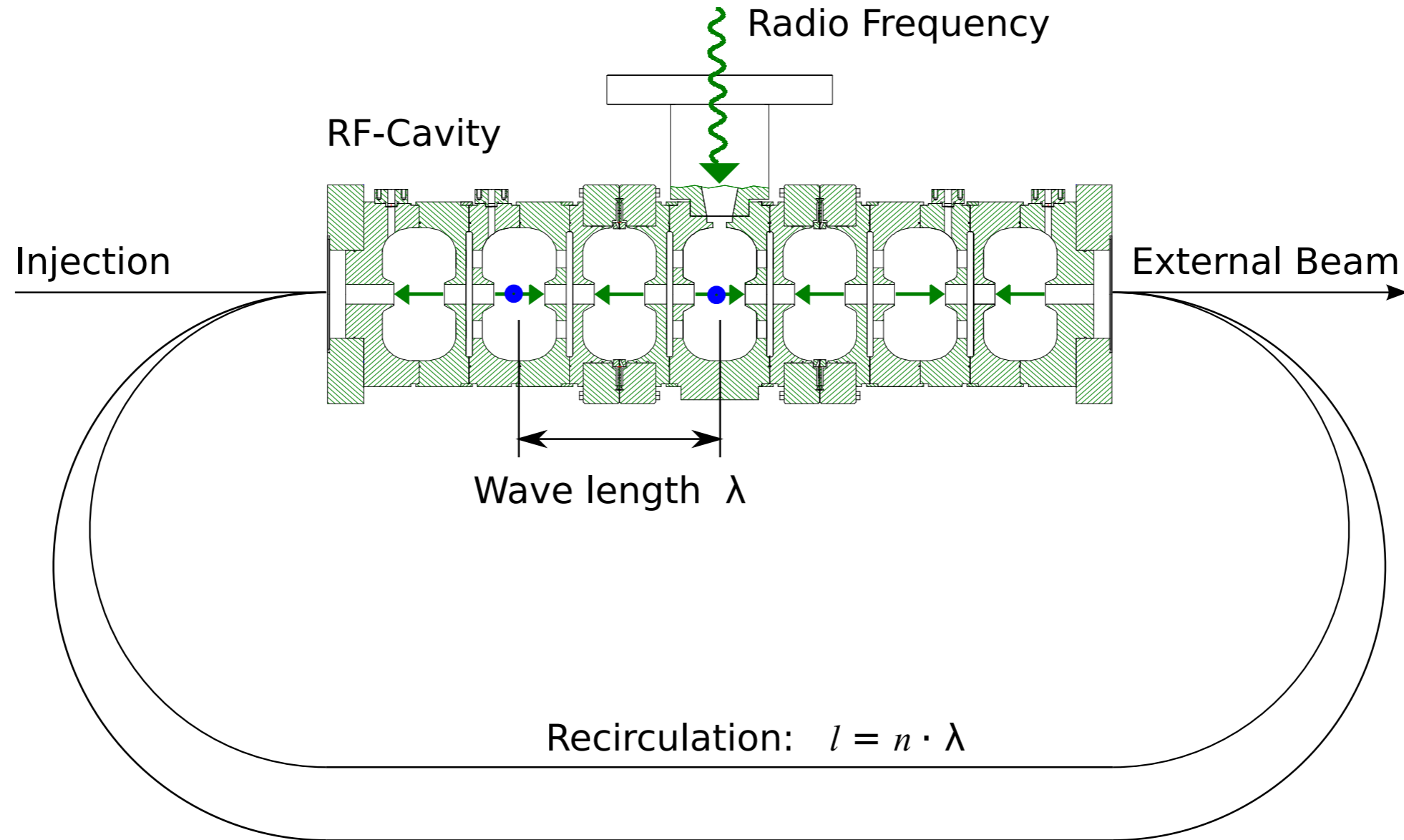
Modified Rossendorf-type Modules



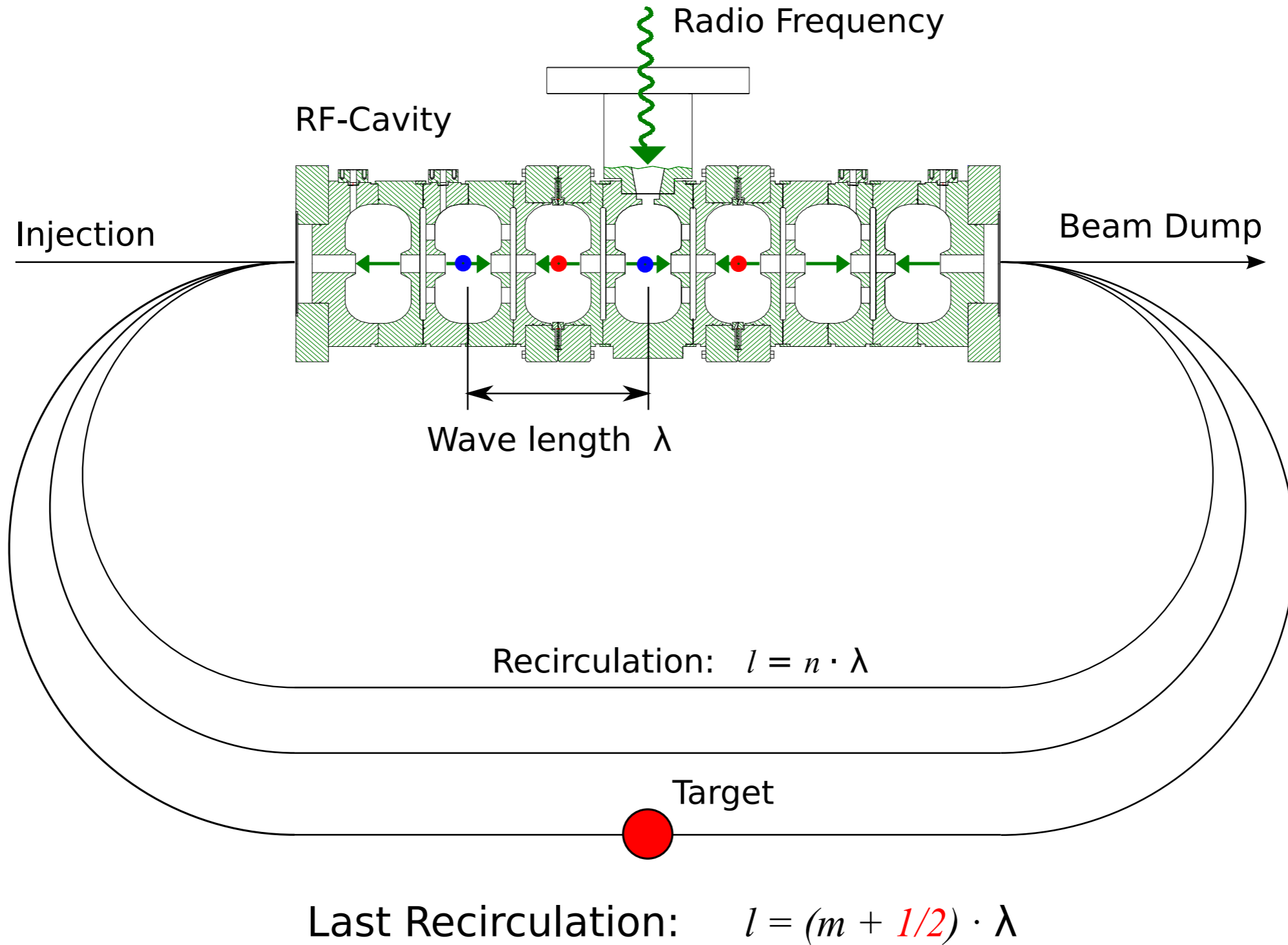
Available Space for an
additional experiment



Energy Recovery Linac

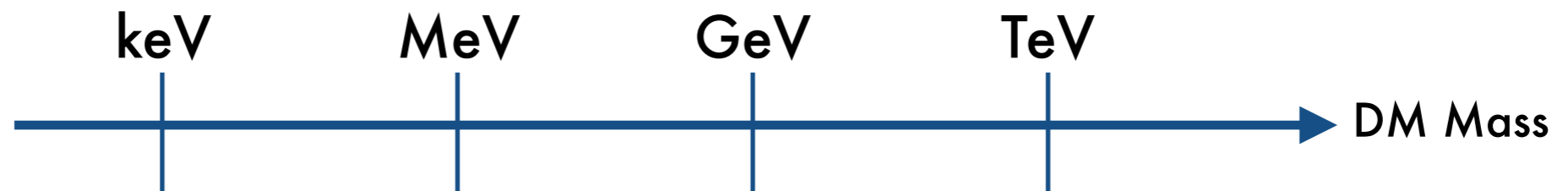


Energy Recovery Linac

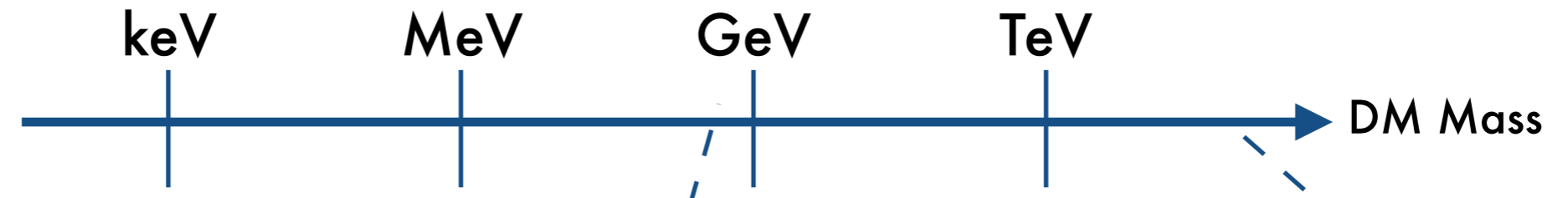


Light Dark Matter

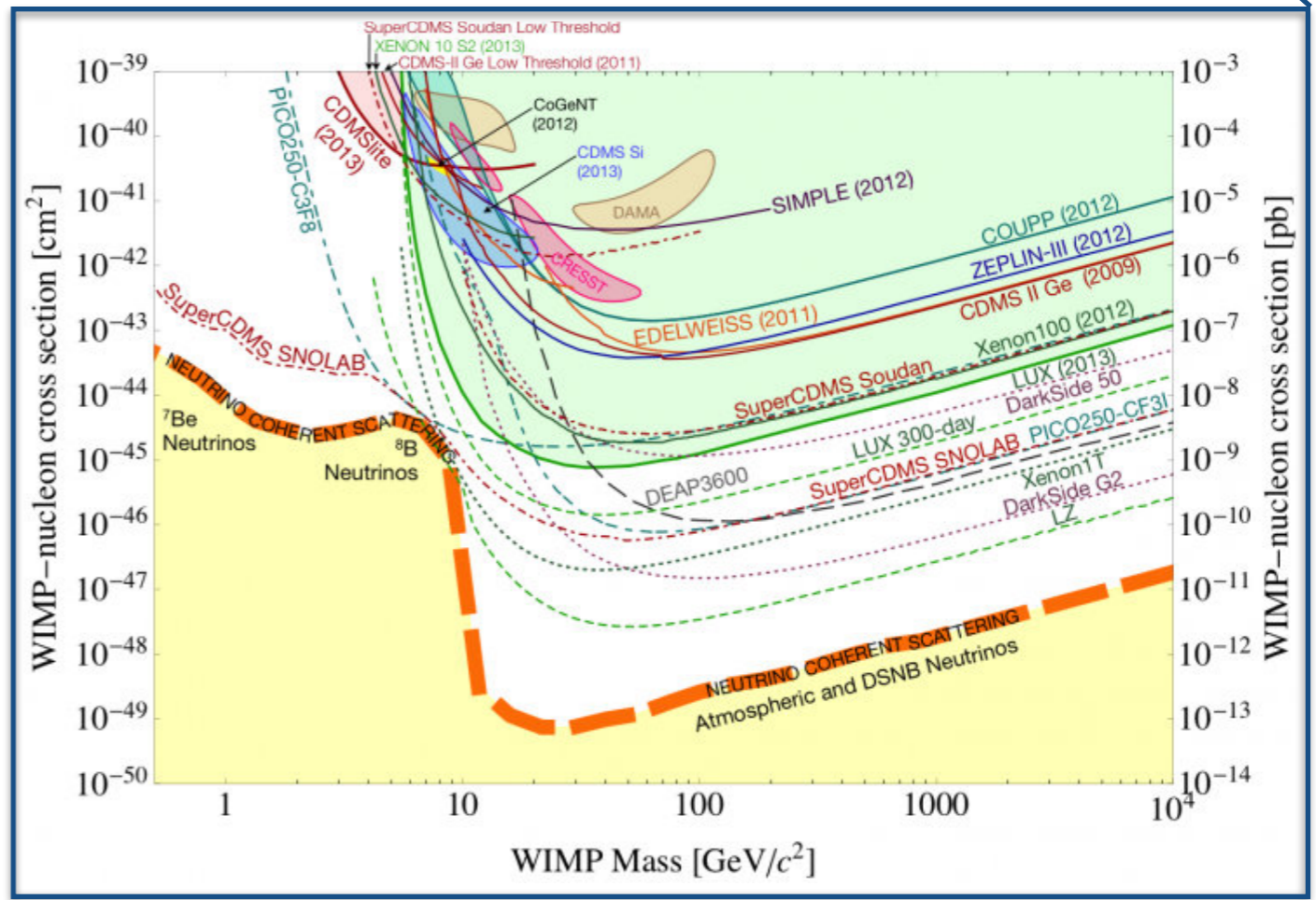
(Light) Dark Matter



(Light) Dark Matter



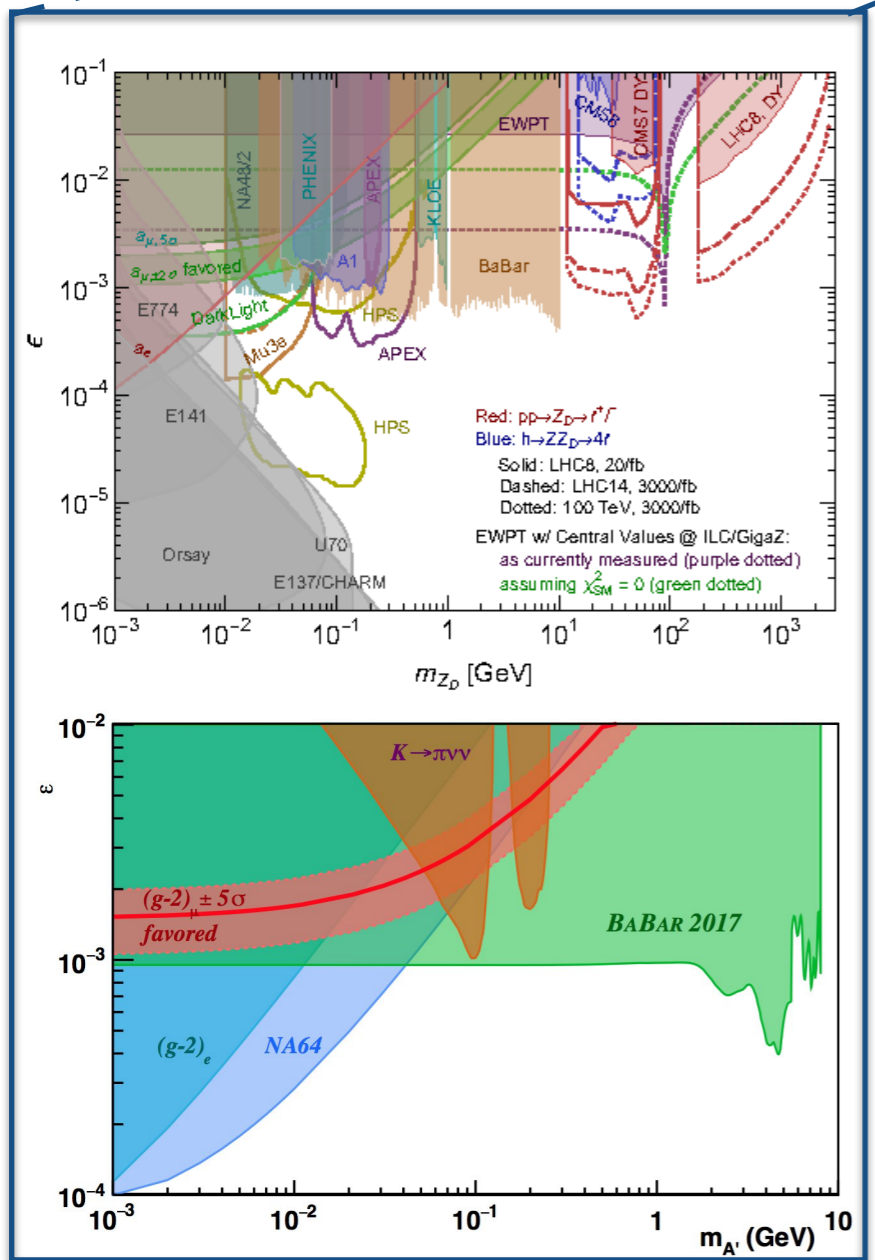
WIMPs



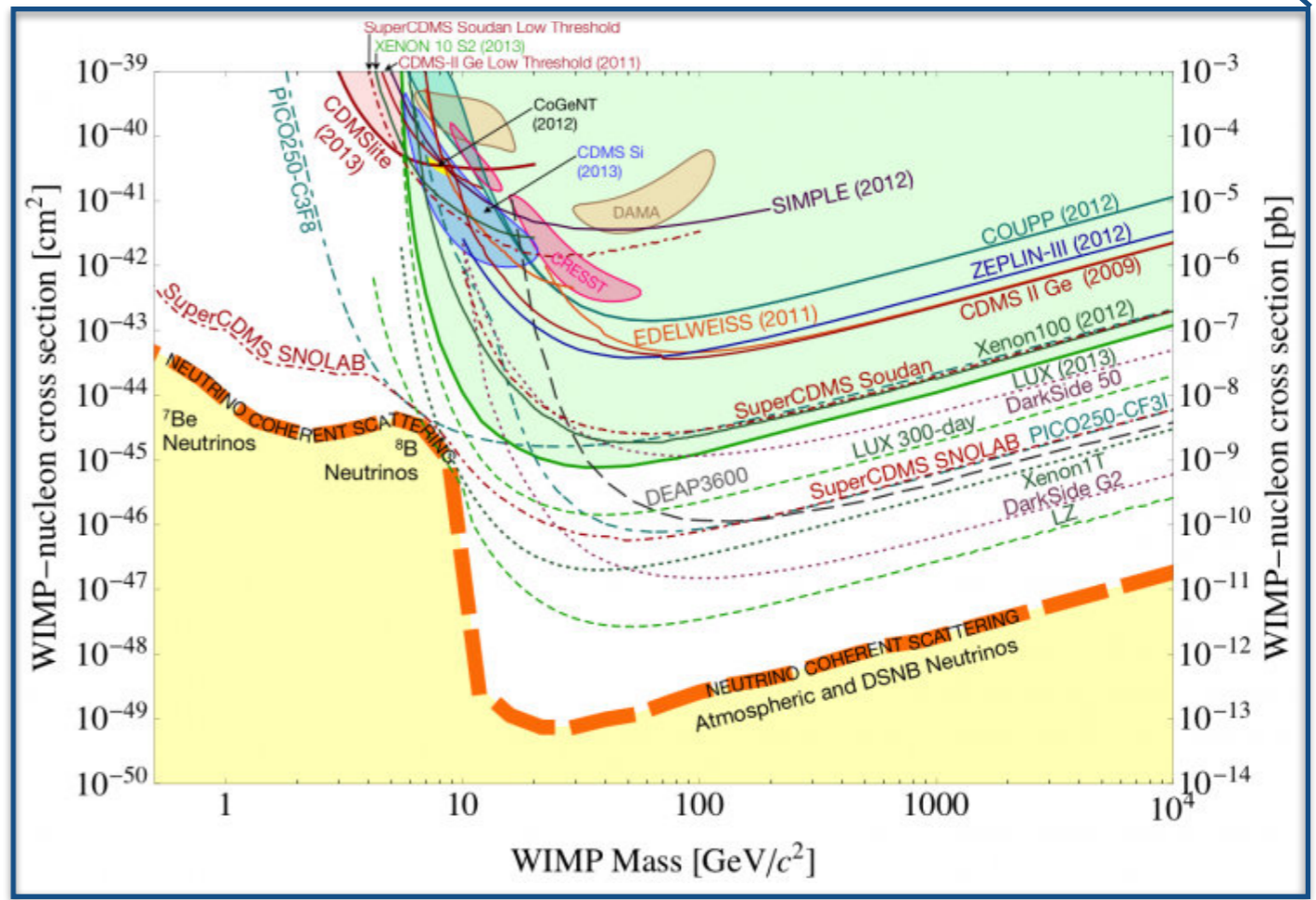
(Light) Dark Matter



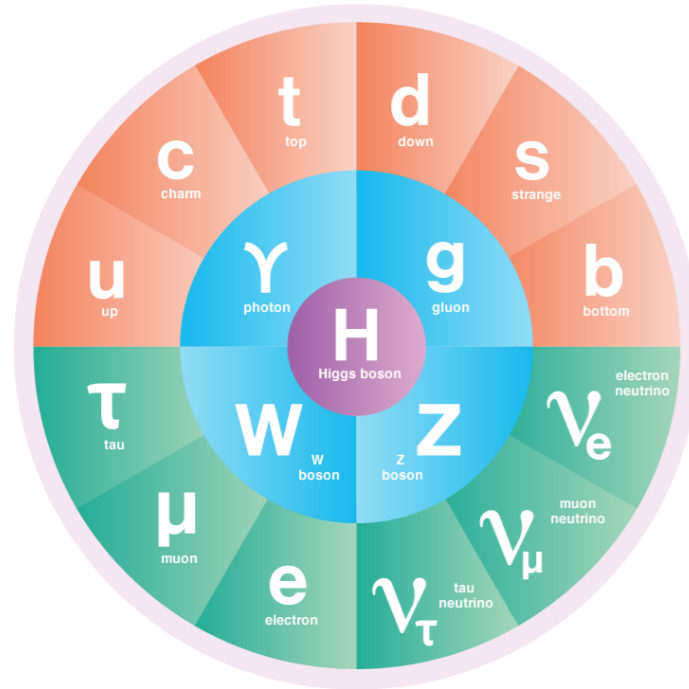
LDM



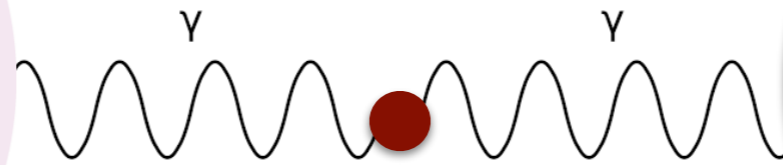
WIMPs



Standard Model



Dark Sector



“Portals”

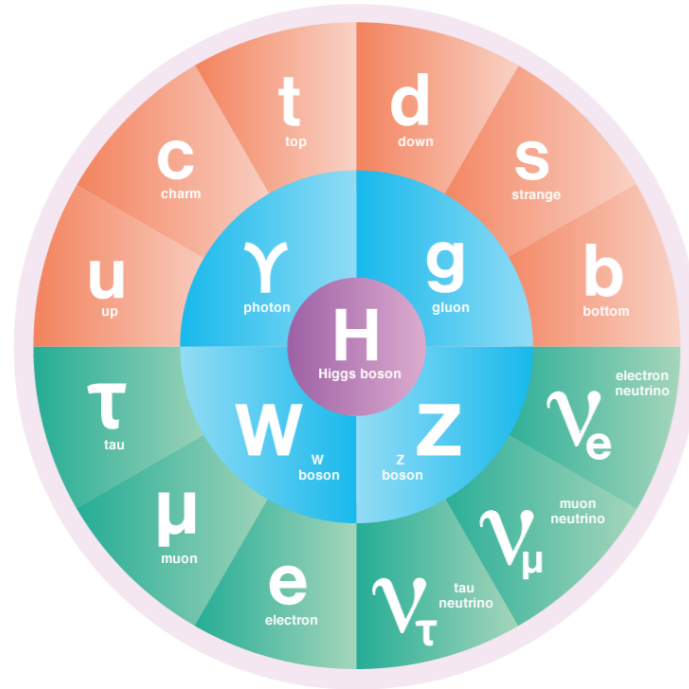
Vector Portal $\frac{1}{2} \epsilon_Y F_{\mu\nu} F'^{\mu\nu}$

Higgs Portal $\epsilon_h |h|^2 |\phi|^2$ Precision Higgs Physics

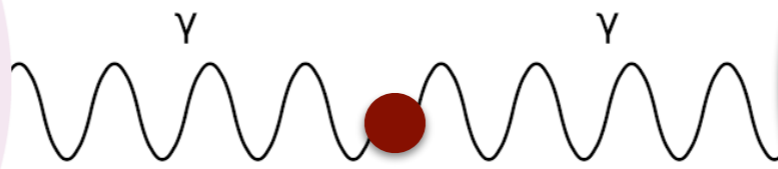
Neutrino Portal $\epsilon_\nu h L \psi$ New Neutrino States

Axion Portal $\frac{G_{a\gamma\gamma}}{4} a F_{\mu\nu} \tilde{F}^{\mu\nu}$

Standard Model



Dark Sector



“Portals”

Vector Portal $\frac{1}{2} \epsilon_Y F_{\mu\nu} F'^{\mu\nu}$

Higgs Portal $\epsilon_h |h|^2 |\phi|^2$ Precision Higgs Physics

Neutrino Portal $\epsilon_\nu h L \psi$ New Neutrino States

Axion Portal $\frac{G_{a\gamma\gamma}}{4} a F_{\mu\nu} \tilde{F}^{\mu\nu}$

Minimal Dark Photon Model

$$\mathcal{L} \sim \bar{\chi}(i\not{D} - m_\chi)\chi + \frac{1}{2} \epsilon_Y F'_{\mu\nu} B_{\mu\nu} + \frac{1}{2} m_{A'}^2 A'_\mu A'^\mu$$

$D_\mu = \partial_\mu + ig_D A'_\mu$ New U(1) massive gauge boson

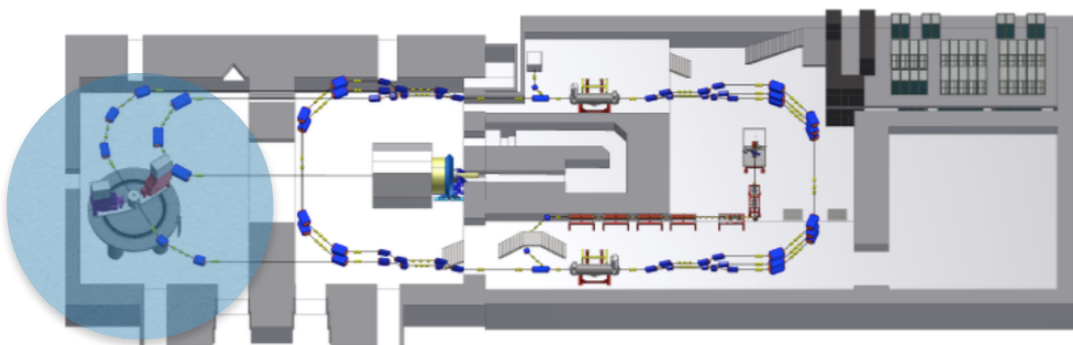
After EW Symmetry Breaking:

$$\epsilon = \epsilon_Y \cos \theta_W \ll 1 \quad \frac{1}{2} \epsilon F'_{\mu\nu} F_{\mu\nu}$$

4 parameters: $m_{A'} \quad \alpha_D = \frac{g_D^2}{4\pi} \quad m_\chi \quad \epsilon_Y$

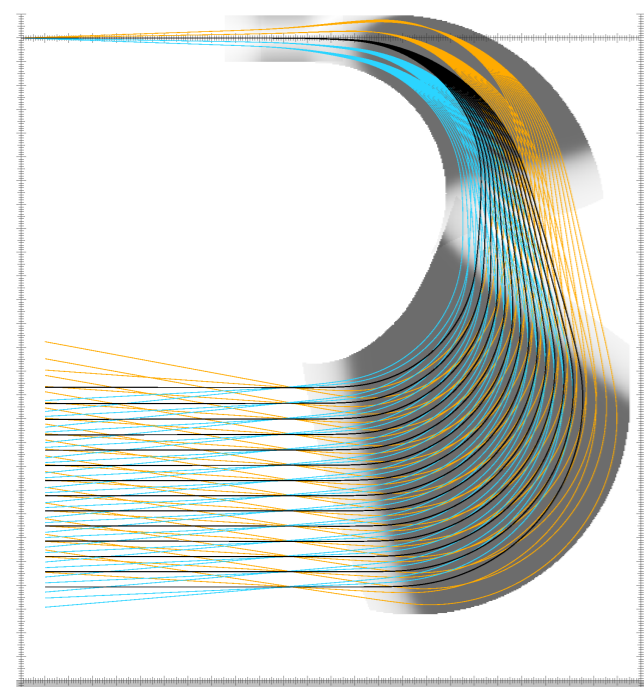
The MAGIX Experiment

The MAGIX Experiment

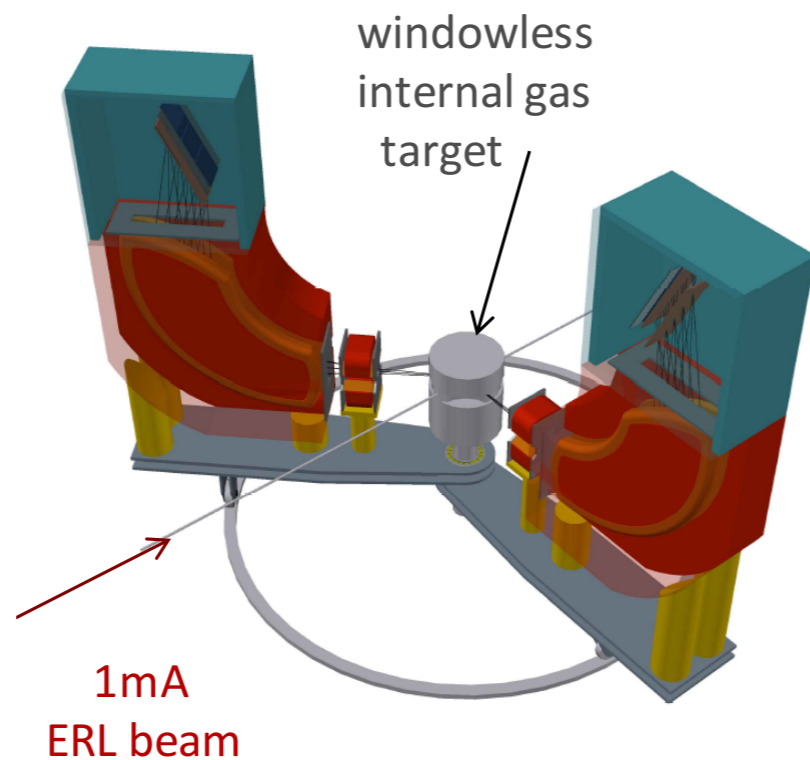


- New technique in low-energy part/nucl phys experiments
- Competitive luminosity (mA beam current)
- Double arm spectrometer (ns time coincidence)
- High-resolution tracking at low energies $dp/p = 10^{-4}$
- Acceptance +/- 50mr
- Focal plane detectors current design options:
 - low material budget GEMs
 - Time Projection Chamber with GEM readout

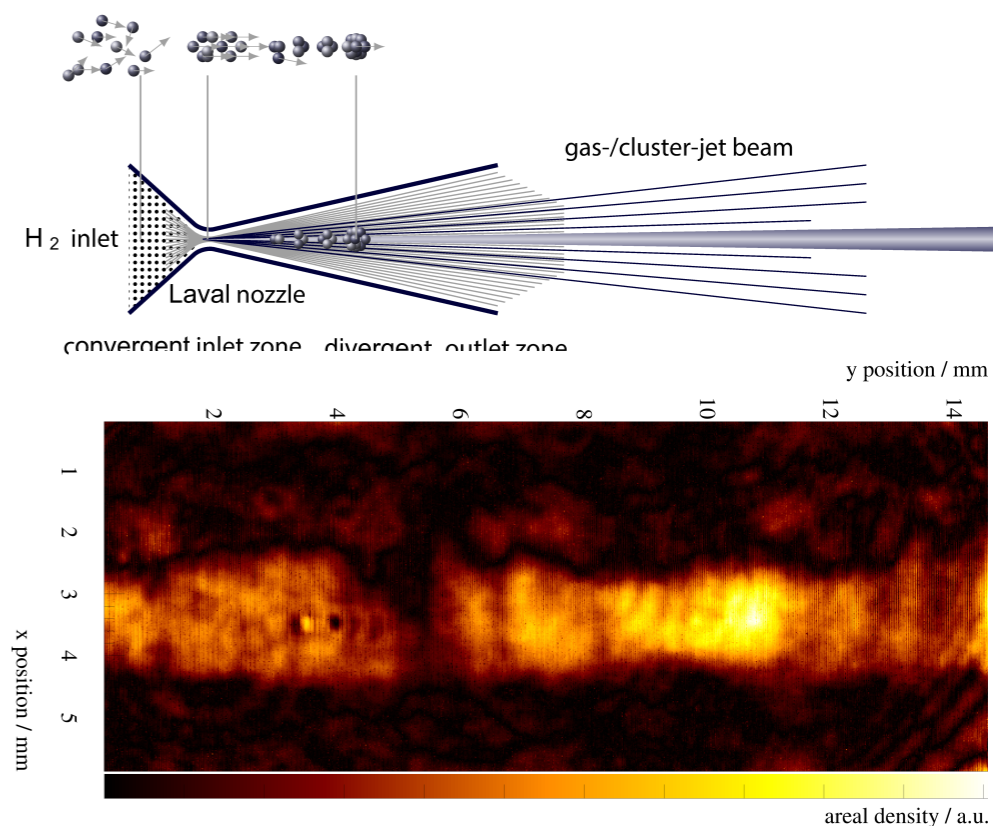
Magnetic optics studies



Experimental Setup



Gas-Jet Target

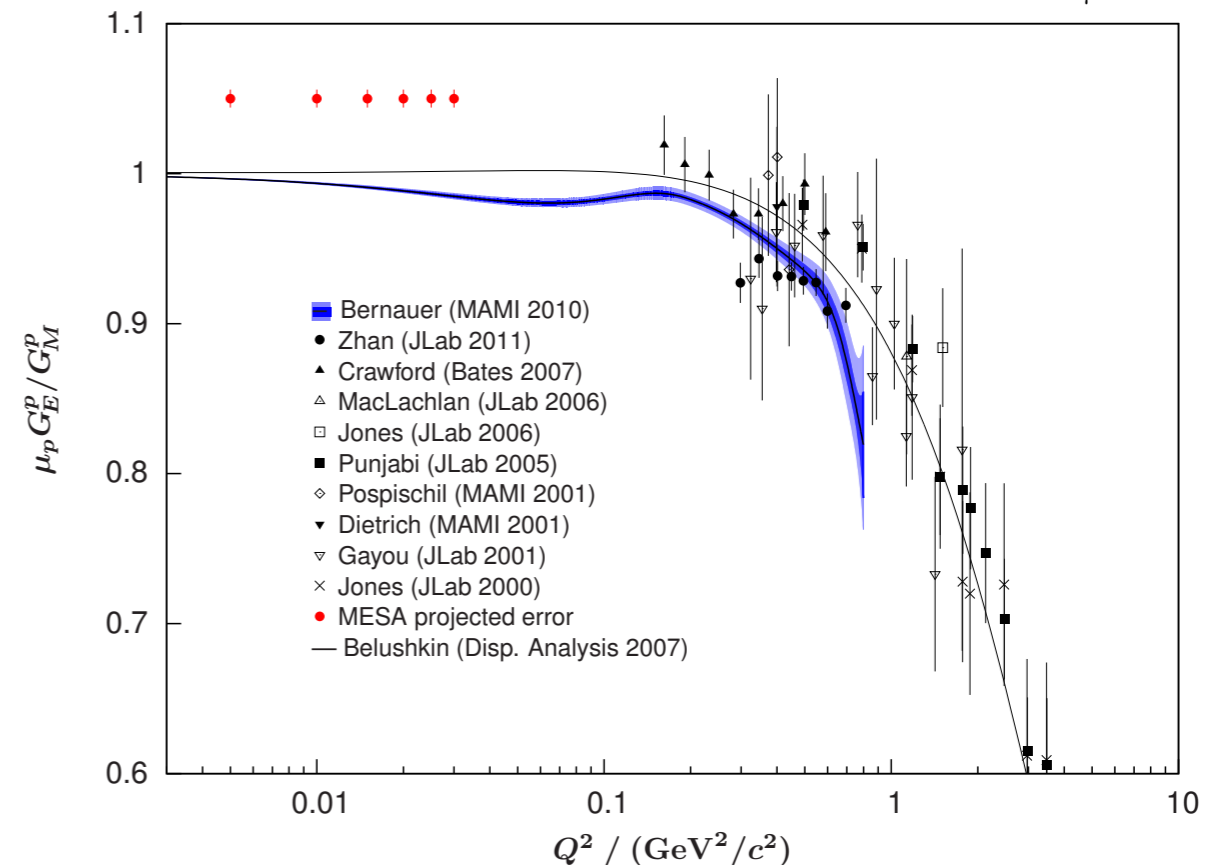
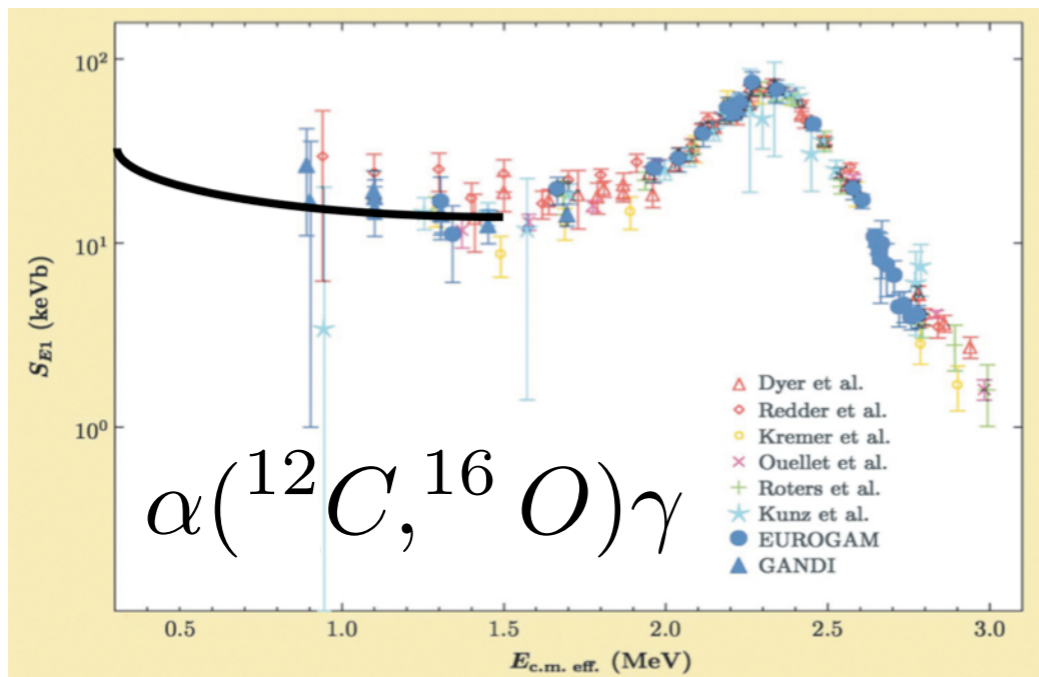
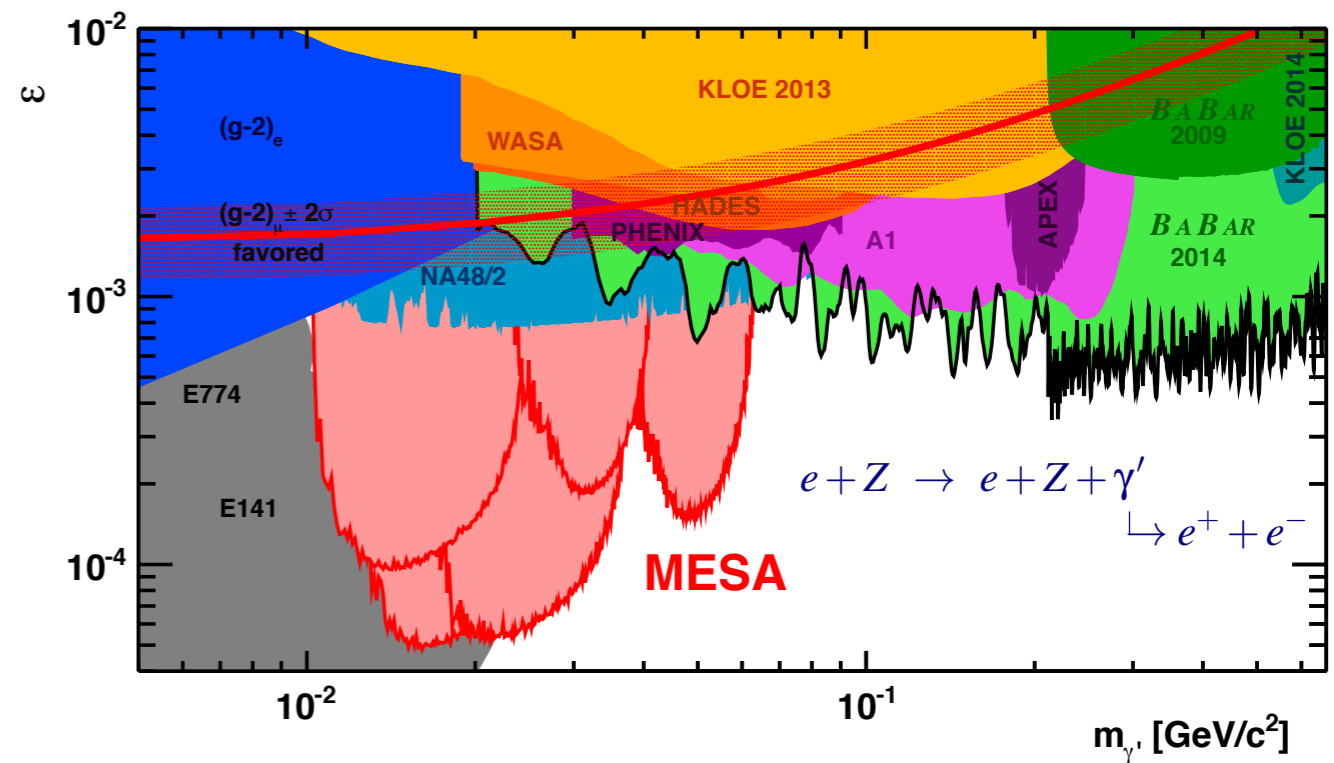


Physics:

- Nucleon Form factors (proton radius puzzle)
- Dark photon (visible and invisible decays)
- Nuclear Astrophysics

Other possibilities

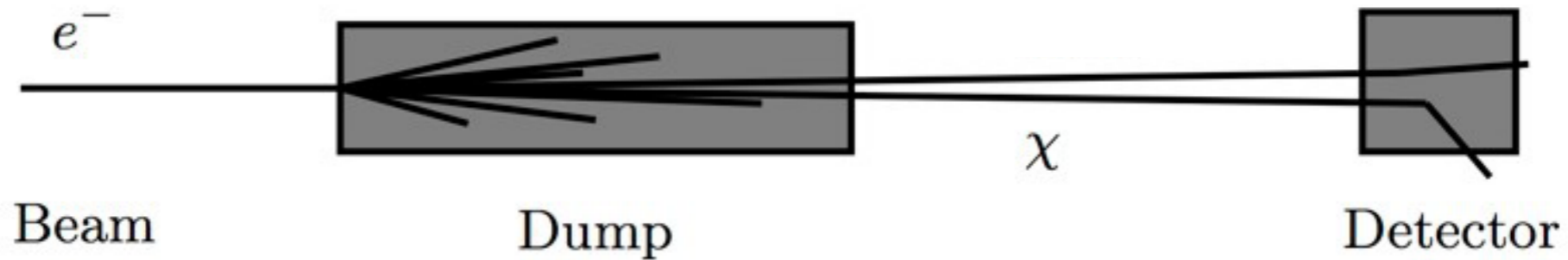
- Nuclear physics (3-body forces)
- Nucleon polarizabilities
- ...



BDX@MESA

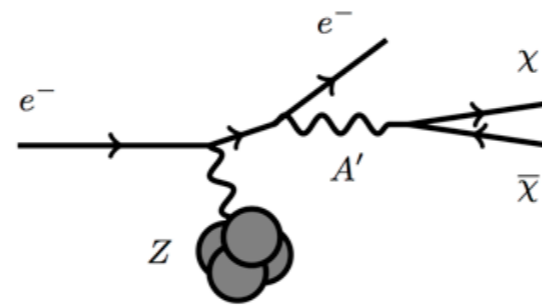
A Beam-Dump Experiment at MESA

Bjorken et al., Phys. Rev. D80, 075018 (2009)

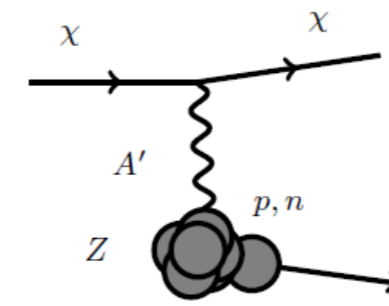


$$\frac{d\sigma}{dx} \approx \frac{8Z^2 \alpha^3 \epsilon^2 x}{m_{A'}^2} \left(1 + \frac{x^2}{3(1-x)} \right) \mathcal{L}og$$

$$\theta_{A' \max} \sim \max \left(\frac{\sqrt{m_{A'} m_e}}{E_0}, \frac{m_{A'}^{3/2}}{E_0^{3/2}} \right)$$

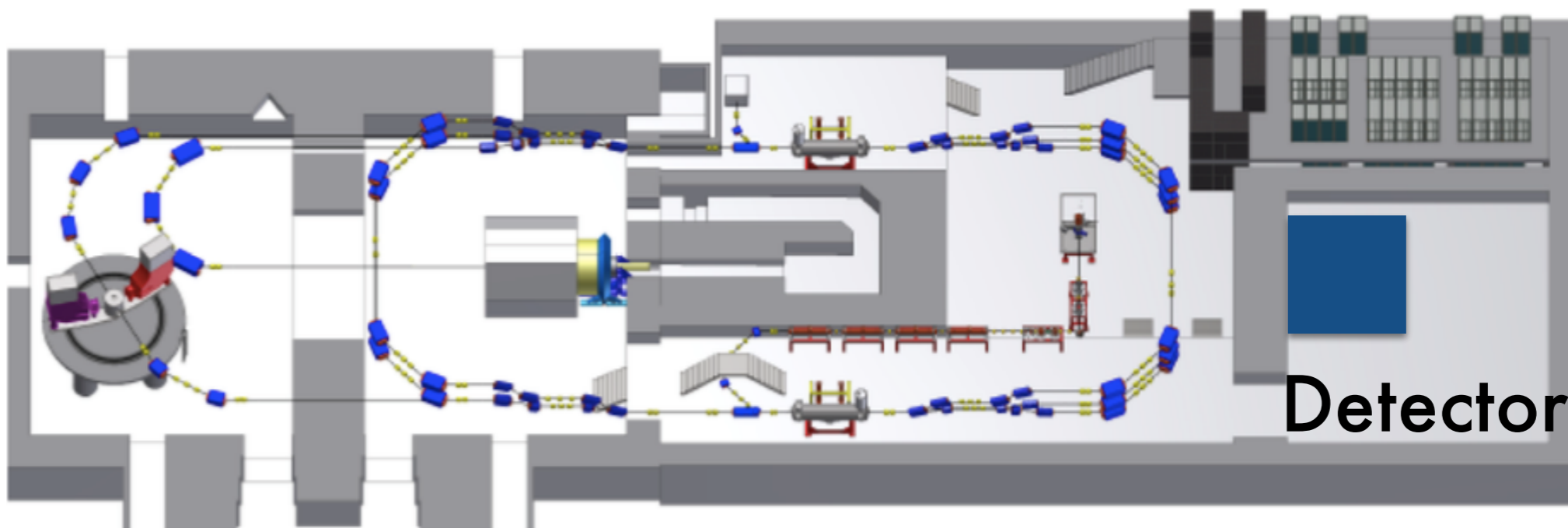


$$Y_{Prod} \sim \epsilon^2 / m_{A'}^2$$



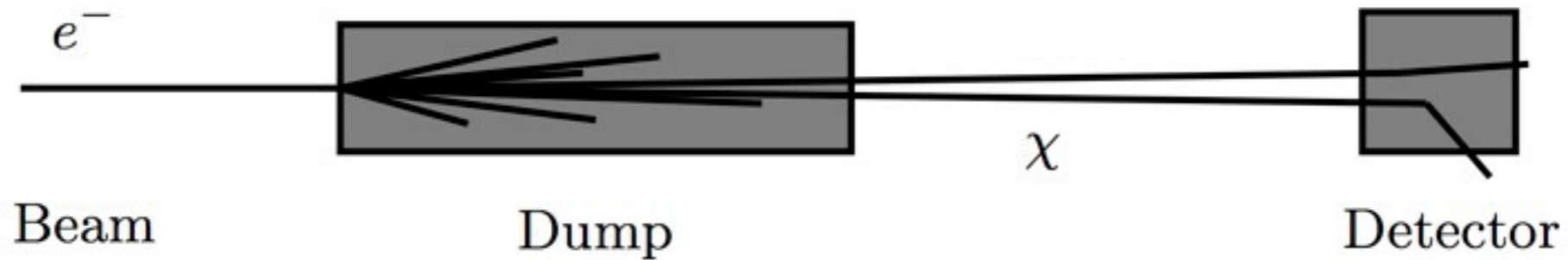
$$Y_{Det} \sim \epsilon^2 \alpha_D / m_{A'}^2$$

Total Yield $Y_{TOT} \sim \epsilon^4 \alpha_D / m_{A'}^4$



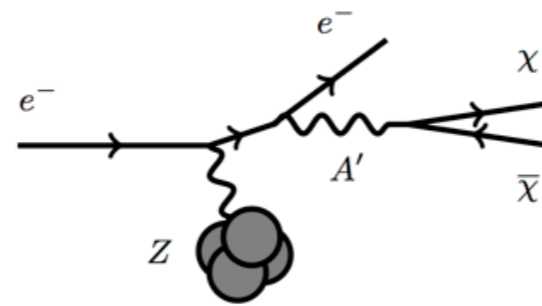
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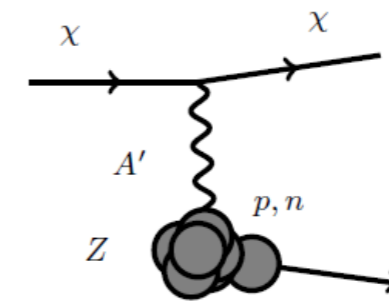


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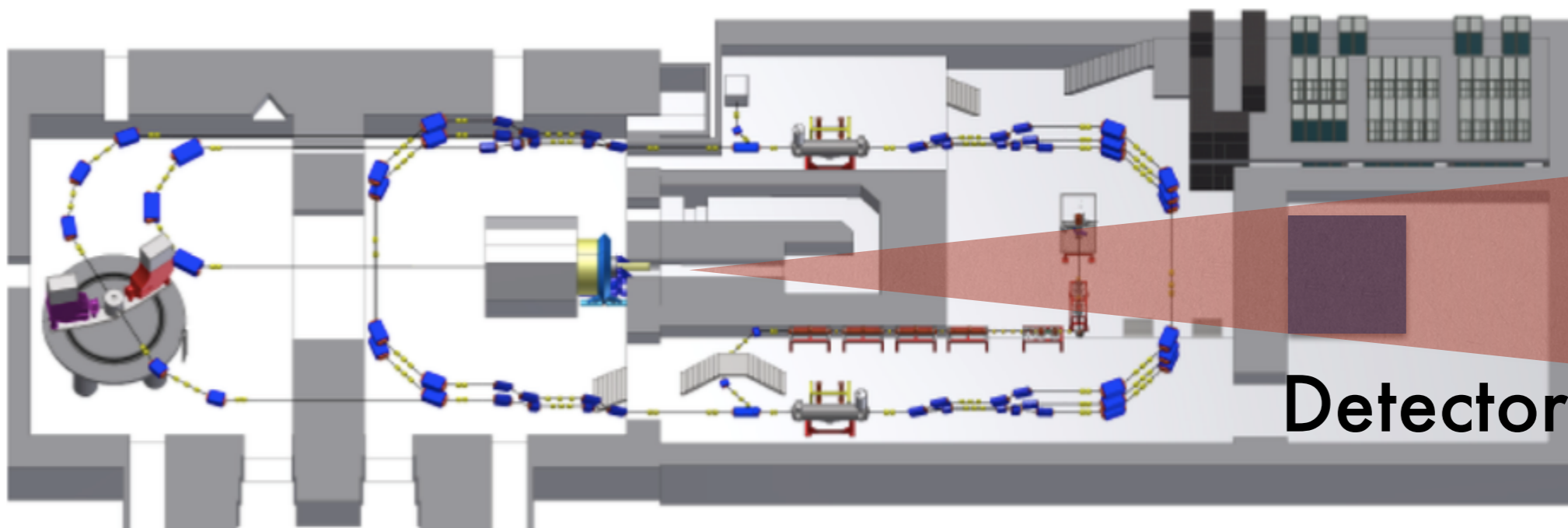


$$Y_{Prod} \sim \epsilon^2 / m_{A'}^2$$



$$Y_{Det} \sim \epsilon^2 \alpha_D / m_{A'}^2$$

Total Yield $Y_{TOT} \sim \epsilon^4 \alpha_D / m_{A'}^4$



$$\theta_{RMS} \sim \frac{m_{A'}}{E_{beam}}$$

Detector

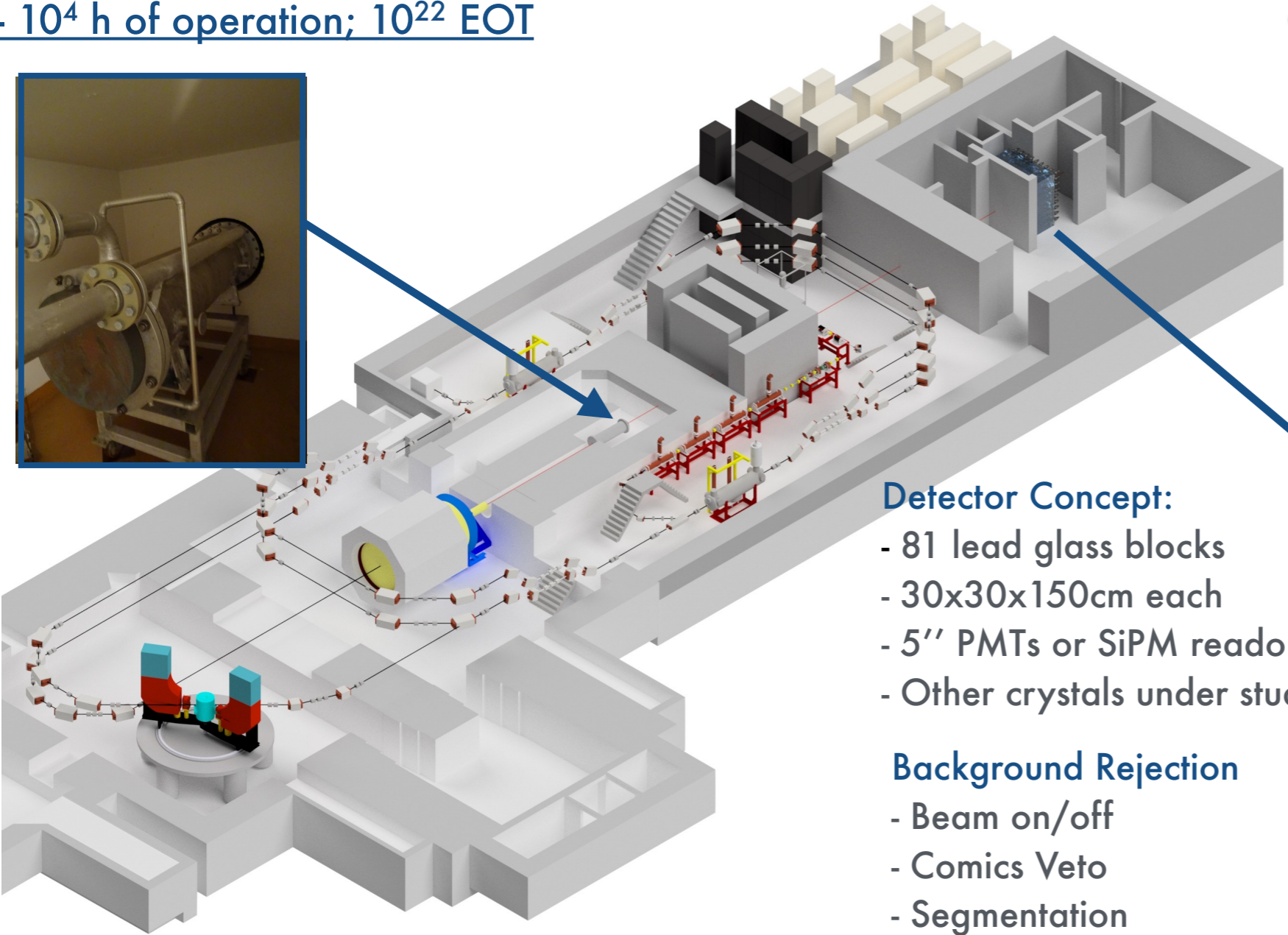
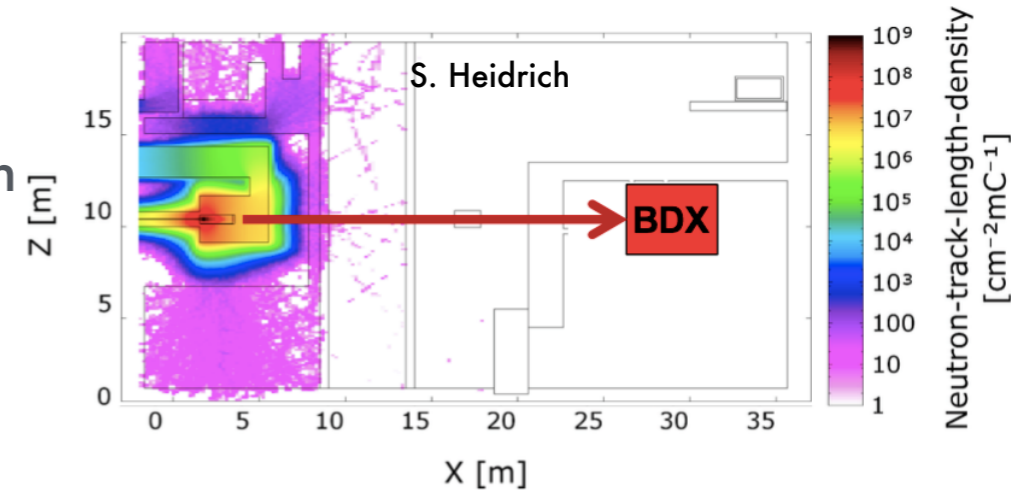
Opportunity: A BD Experiment at MESA

Beam Dump

- 20 X_0 Beam Dump
- Material: Aluminum (+ Water)
- Addition of a W plate?
- Energy on Dump: ~ 135 MeV
- 10^4 h of operation; 10^{22} EOT

Experimental Area

- 70 X_0 (~ 8 m) barite concrete
- \sim no neutrons at detector position
- no beam dump backgrounds
- No neutrinos



Detector:

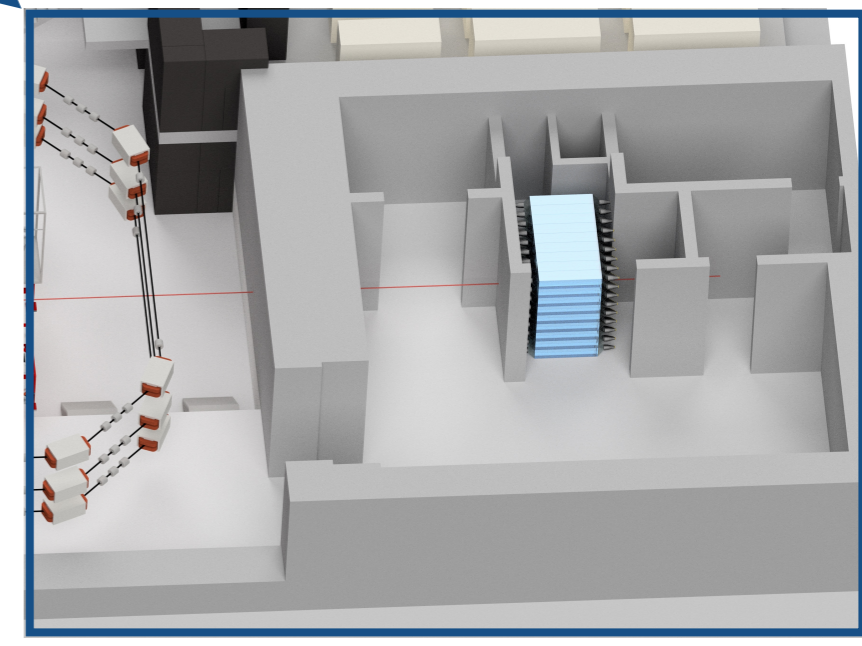
- Maximize active volume
- Maximize Density
- Directionality

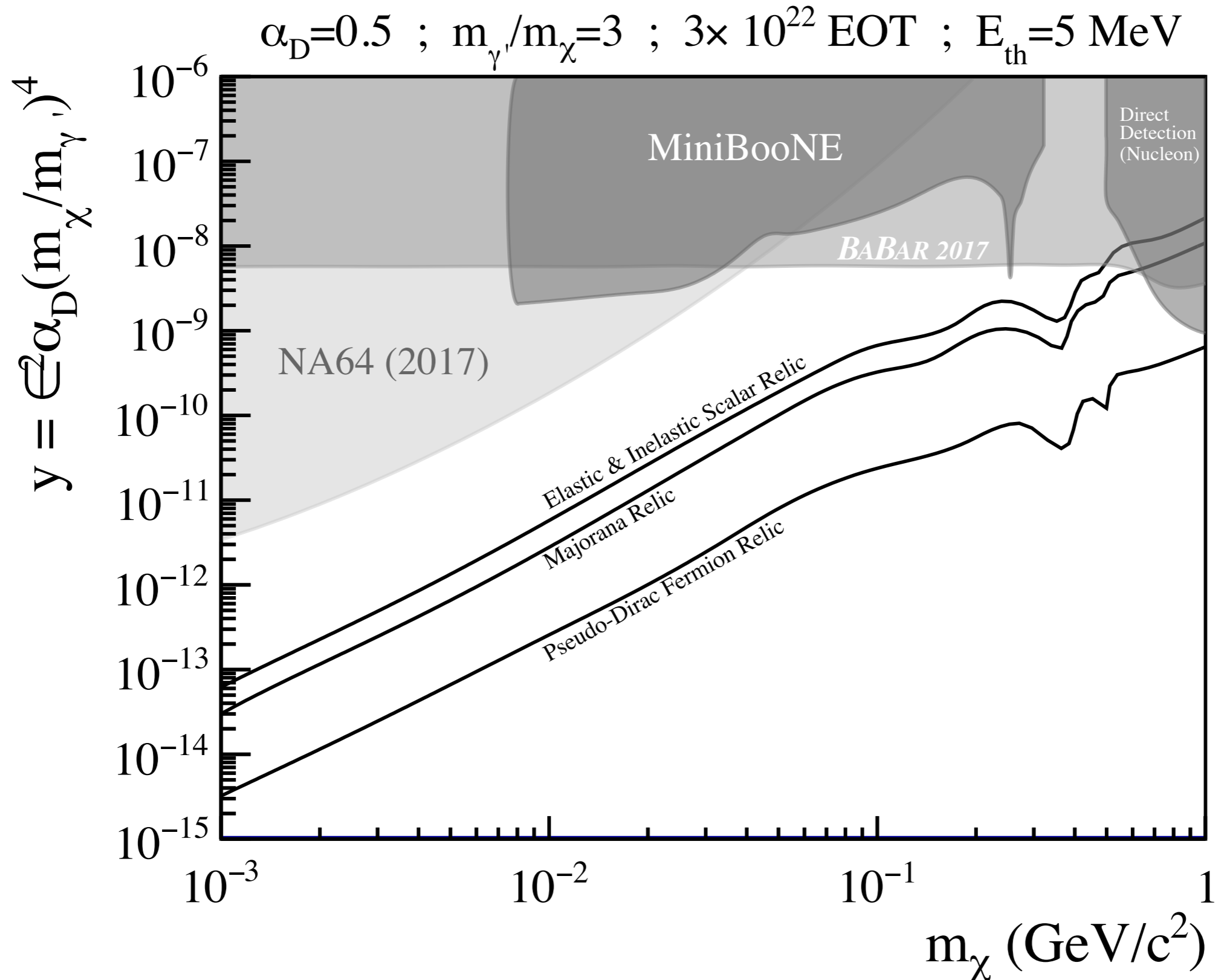
Detector Concept:

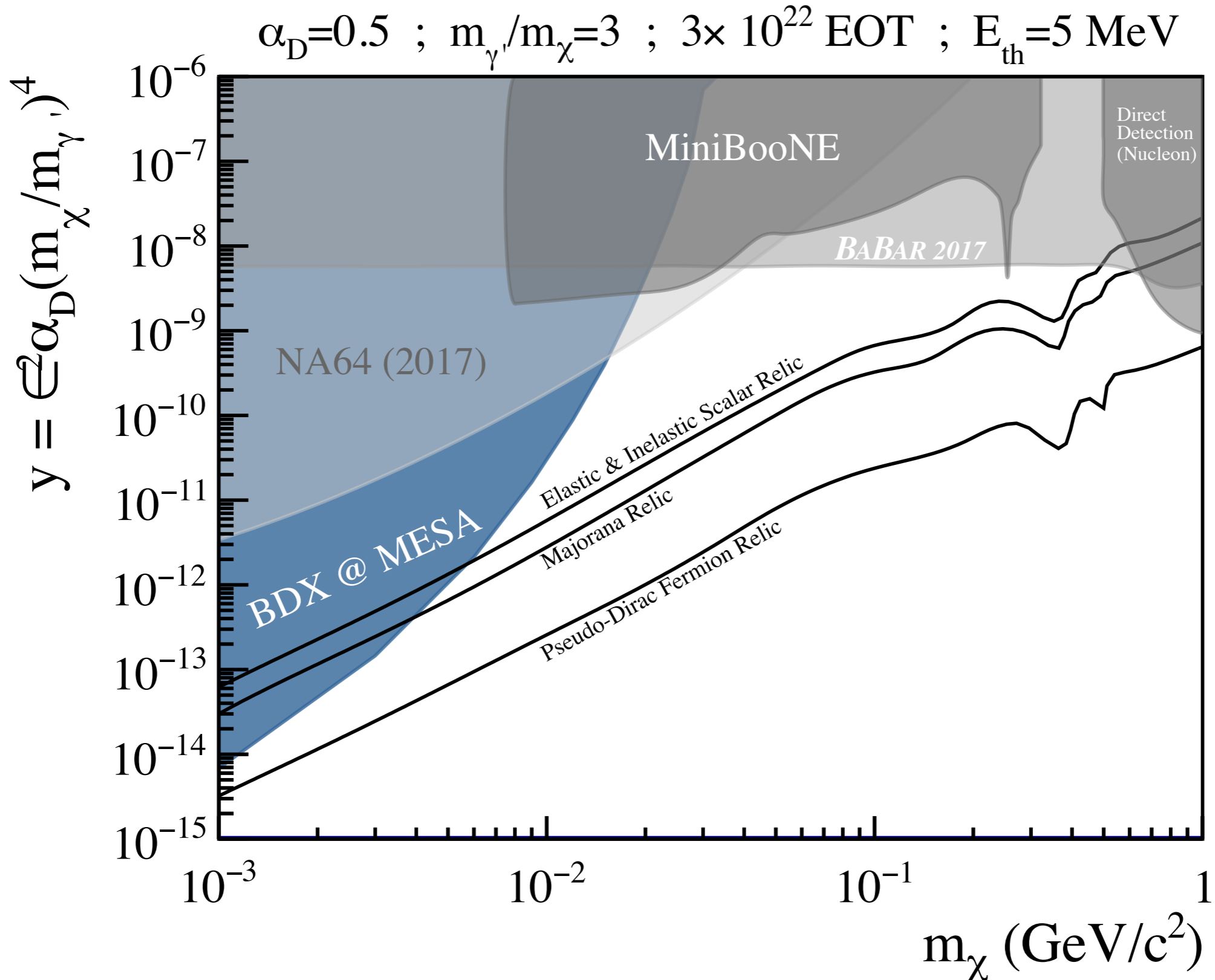
- 81 lead glass blocks
- 30x30x150cm each
- 5'' PMTs or SiPM readout
- Other crystals under study

Background Rejection

- Beam on/off
- Comics Veto
- Segmentation







Potential for significant contributions to LDM searches:

- Visible Decays: expand current limits in new territory (MAGIX)
- Invisible Decays: still large parameter space to explore (MAGIX + BDX@MESA)
- Unique test for both DM production and interaction.
- High sensitivity: potential to “hit” the thermal target

Beam at MESA ideal:

- Very large luminosity, CW beam, and excellent stability
- Recirculation for MAGIX and parasitic operation + beam-off for BDX@MESA
- Low backgrounds (below pion threshold: no neutrinos/muons)

Near-Future Directions:

- MAGIX: Magnetic optics and detector R&D ongoing (GEMs, TPC)
- BDX@MESA: Extend G4 Simulation+MadGraph , Calorimeter R&D started
 - many possibilities to investigate (Cherenkov/Scintillation crystals)

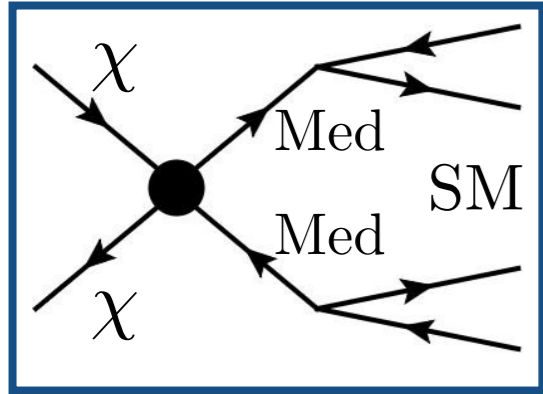
Many Experiments running or planned:

NA62, HPS, SeaQuest, MiniBoone, DarkLight, APEX,
B-Factories, PADME, LDMX, SHiP,

MESA and Dark Sector Experiments could start by >2020 !

Thank You!

Secluded Annihilation

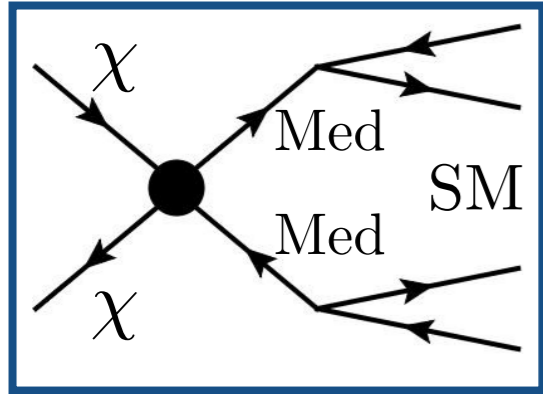


$$m_\chi > m_{Med}$$

$$\langle \sigma v \rangle \sim g_D^4 / m_\chi^4$$

No SM coupling dependence:
arbitrary coupling possible.
Hard to test experimentally.

Secluded Annihilation

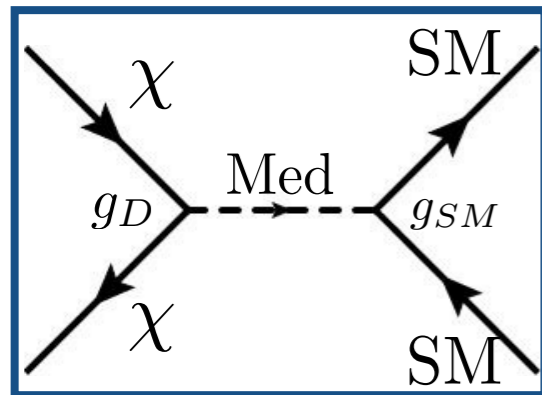


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Direct Annihilation

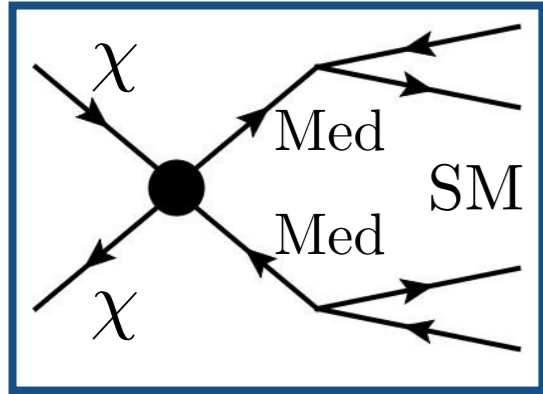


$$m_\chi < m_{Med}$$

$$\langle \sigma v \rangle \sim \frac{g_D^2 g_{SM}^2 m_\chi^2}{m_{Med}^4}$$

There is a minimum SM
coupling compatible with
thermal history.

Secluded Annihilation

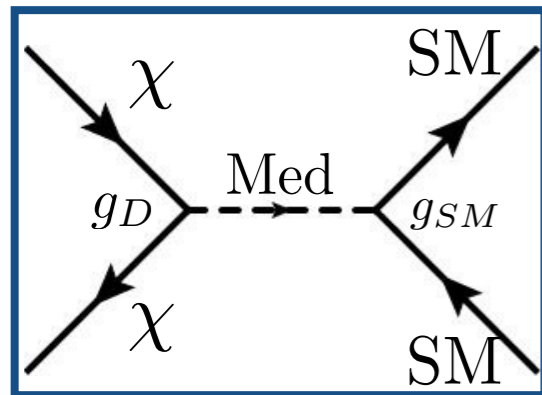


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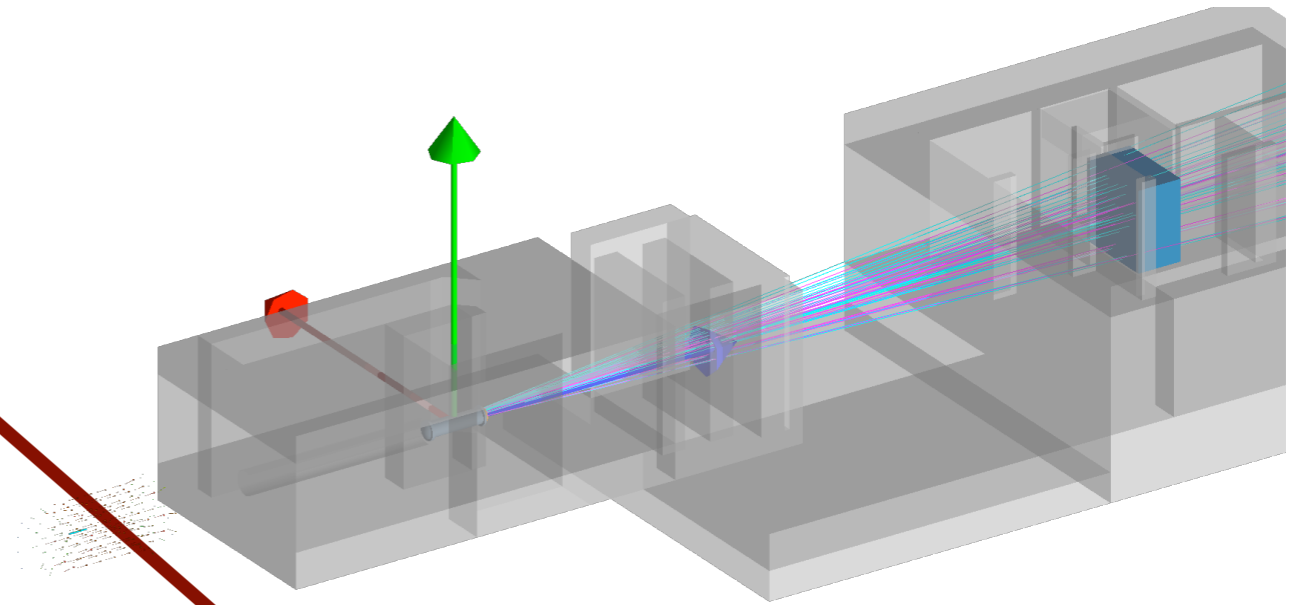
There is a minimum SM
coupling compatible with
thermal history.

Experimentally testable target:

$$y = \frac{g_D^2 g_{SM}^2}{4\pi} \left(\frac{m_\chi}{m_{Med}} \right)^4 \gtrsim m_\chi^2 \langle \sigma v \rangle_{Relic}$$

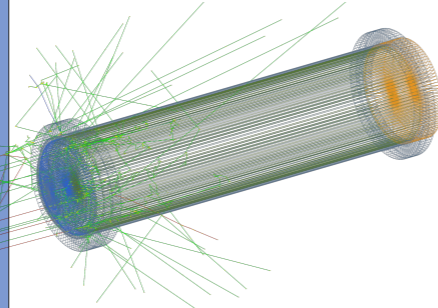
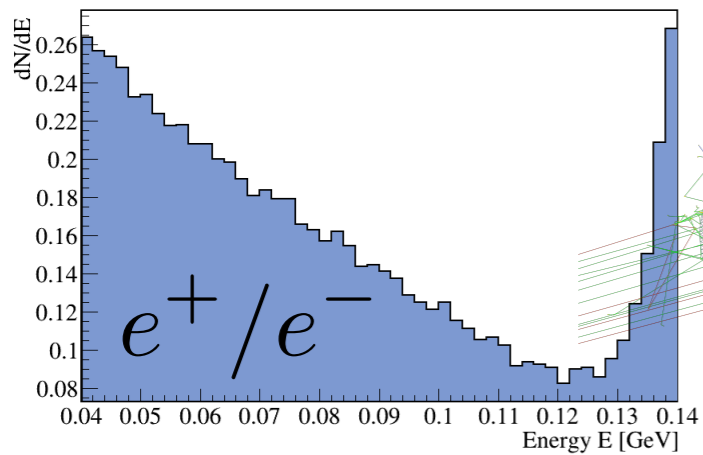
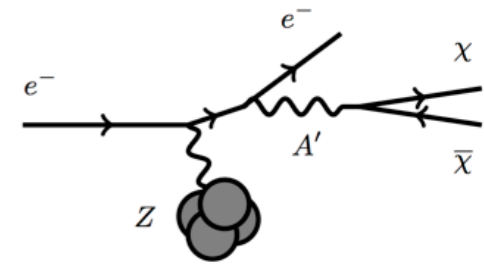
Geant4 Simulation

P2 Target, Magn. Field
Beam Dump, Exp. Halls



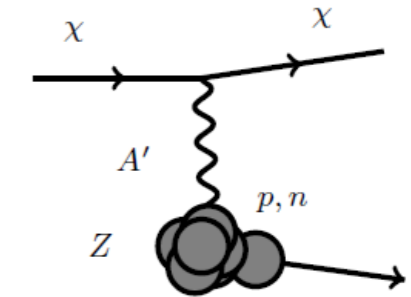
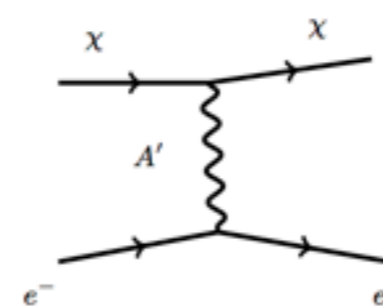
MadGraph

M. Battaglieri
A. Celentano
(INFN Genova)

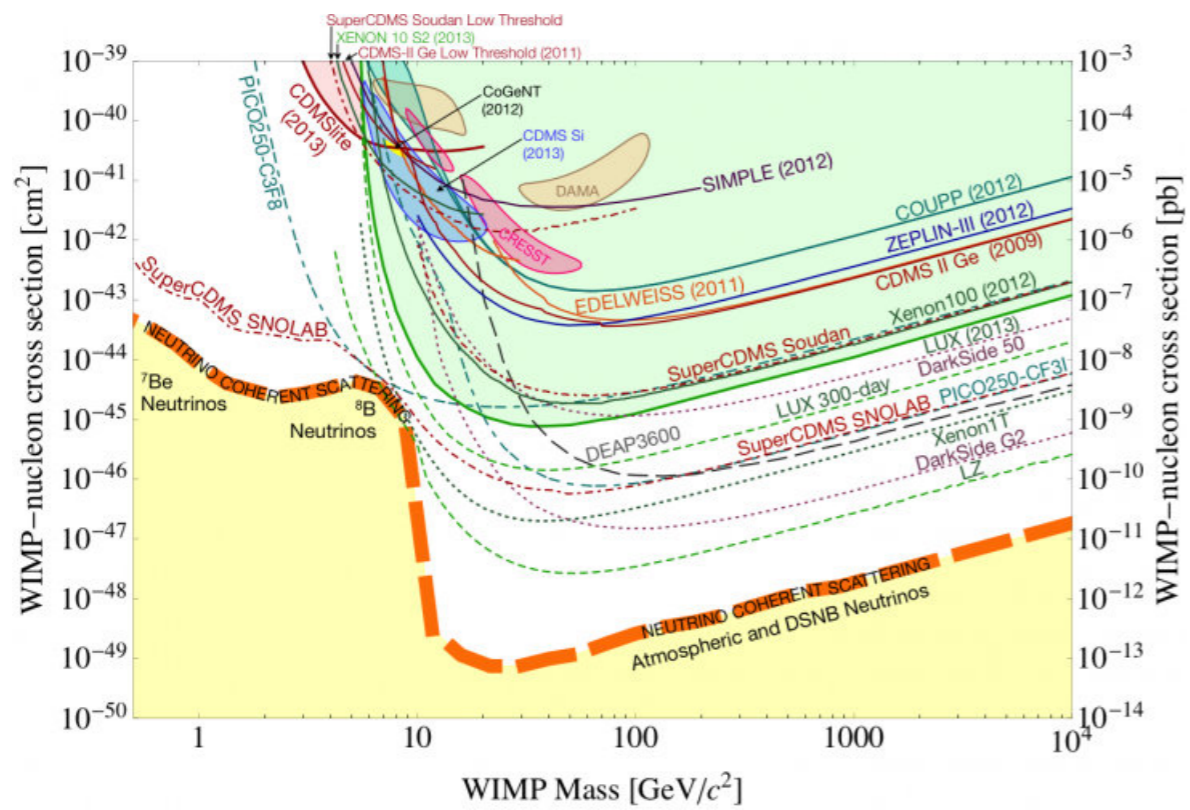


e/N DM Scattering

Detector Response



(WIMP) Direct Dark Matter Searches



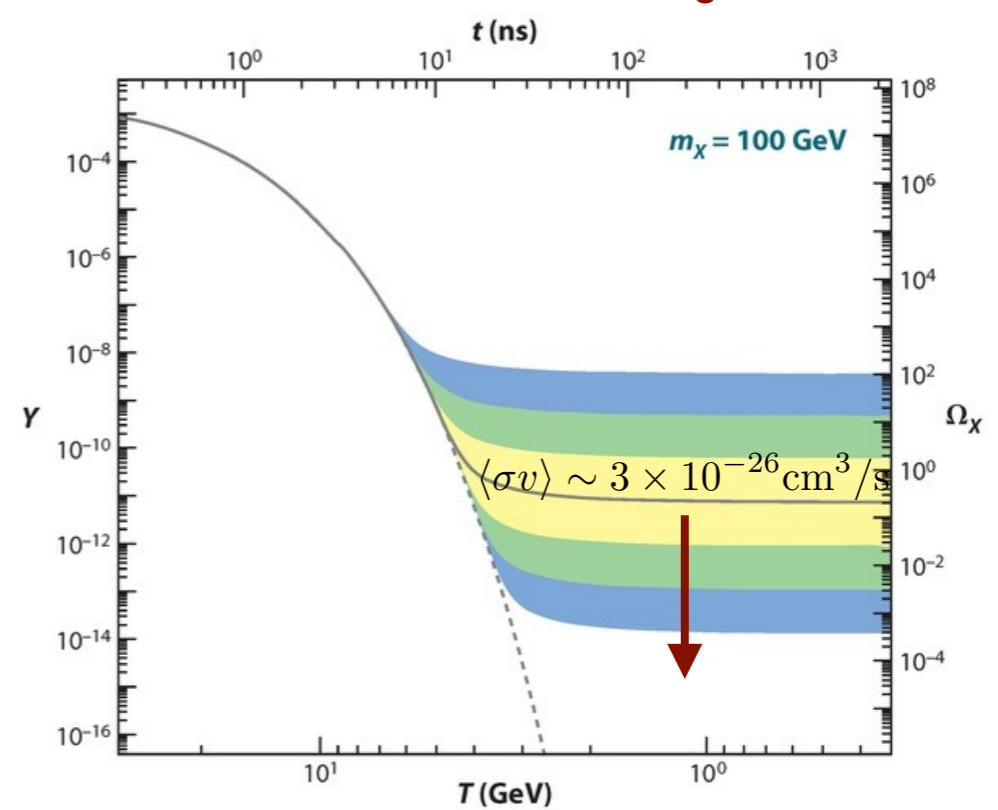
Light Dark Matter

- If light, smaller annihilation CS
- DM overabundance
- "Overclosed" Universe

The way out:

- postulate a new interaction
 - annihilation via a new force carrier
- If coupling small enough, DM can be light!

Thermal Relic Paradigm



Freeze out: $n\langle\sigma v\rangle(T_{FO}) \sim H(T_{FO})$

WIMP "miracle":

$$\left. \begin{aligned} m/T \gg 1 \\ \sigma_{EW} \sim G_F^2 T_{FO} \sim 10^{-8} \text{ GeV}^2 \end{aligned} \right\} \Rightarrow \Omega_{DM} \sim 0.2$$

But:

$$\sigma \sim g^4/m_{DM}^2 \Rightarrow m_{DM} \gg 0.1 \text{ eV}$$

And in general:

$$10 \text{ GeV} < m_{DM} < 120 \text{ TeV}$$

Lee-Weinberg Unitarity Limit

